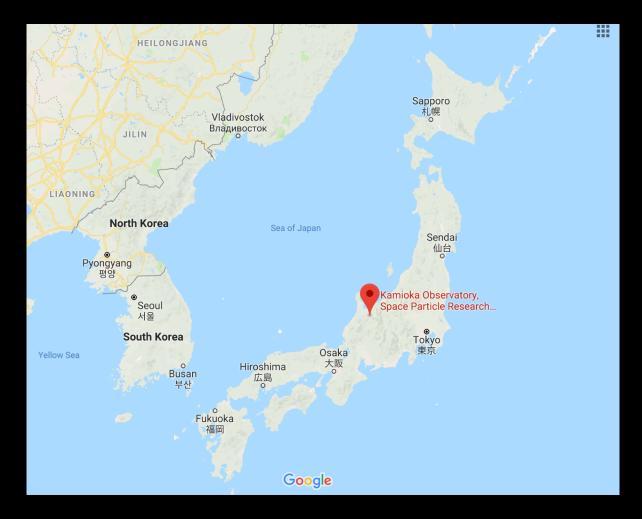
Status of KAGRA: the underground- and cryogenic gravitational-wave detectors

Keiko Kokeyama CRR, University of Tokyo On behalf of the KAGRA collaboration Feb 14th, 2019 KIW5, Perugia, Italy

KAGRA Project

- Underground and Cryogenic interferometric gravitational-wave detector
- Project started in 2010, funded by MEXT, Japan
- International collaboration (97 institutions, 13? countries, 470 members)
- Rapidly being prepared to join the next LIGO-VIRGO observation, O3

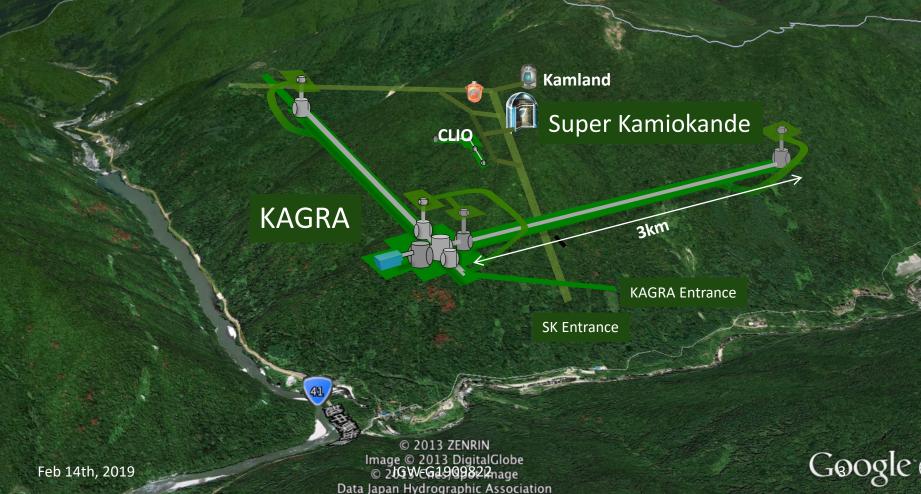
Location of the KAGRA Observatory



Toyama Bay

Toyama City

KAGRA in Kamioka, Japan



Kamioka Site



WELEPLAN.

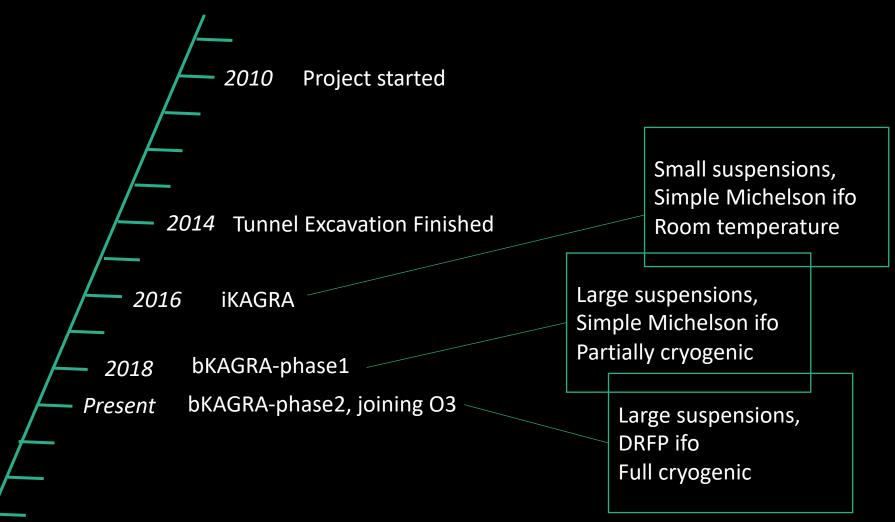
3km arm tunnel



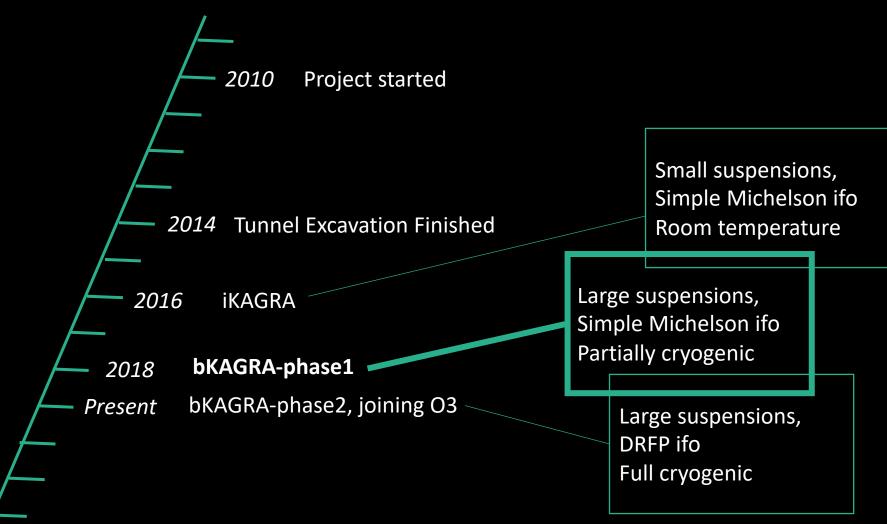
Corner Station



Timeline of the Project

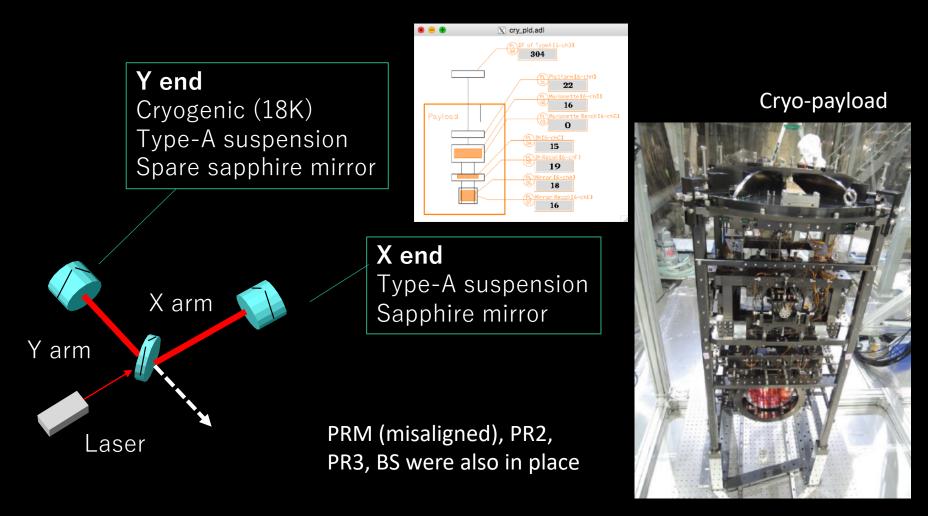


Timeline of the Project

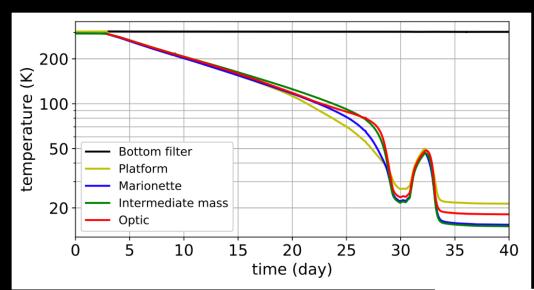


bKAGRA-Phase1 Operation

April 28th – May 6th 2018



Temperature of ETMY

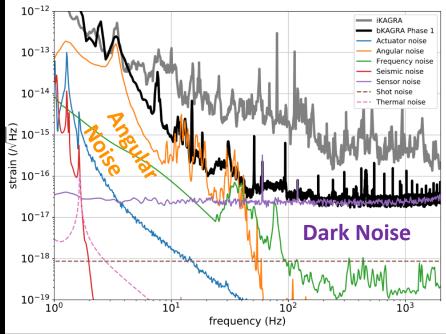


Phase-1 Sensitivity ~3e-17 /rtHz at 100Hz

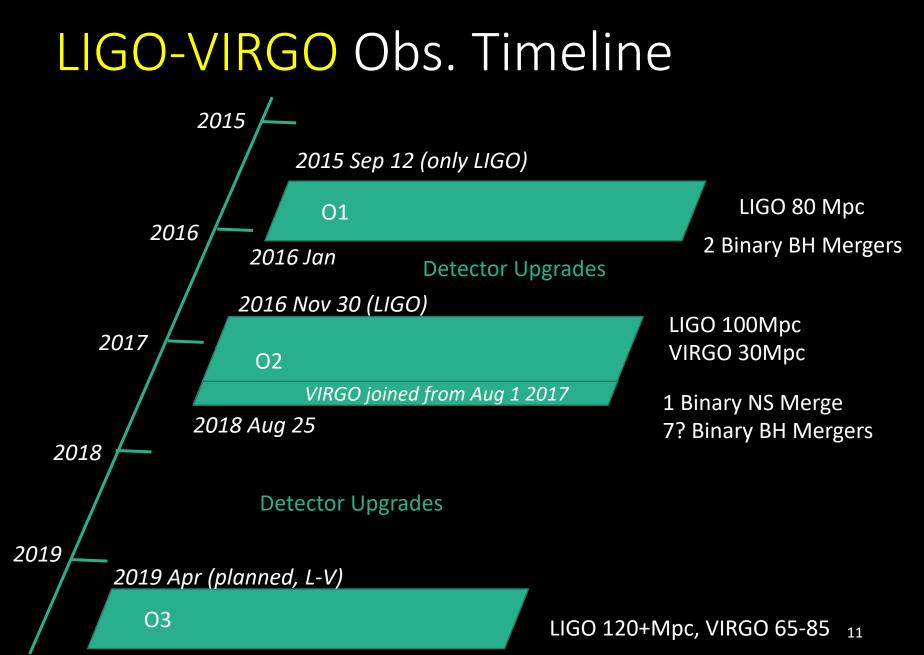
https://arxiv.org/pdf/1901.03569.pdf

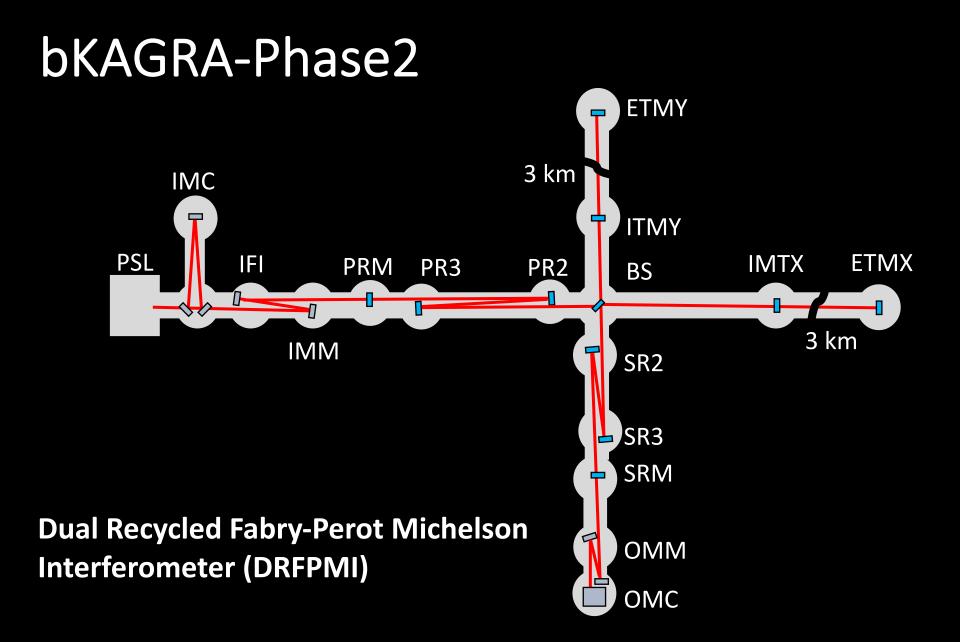
We gained the experiences of...

- First cryogenic operation
- First operation with large suspensions

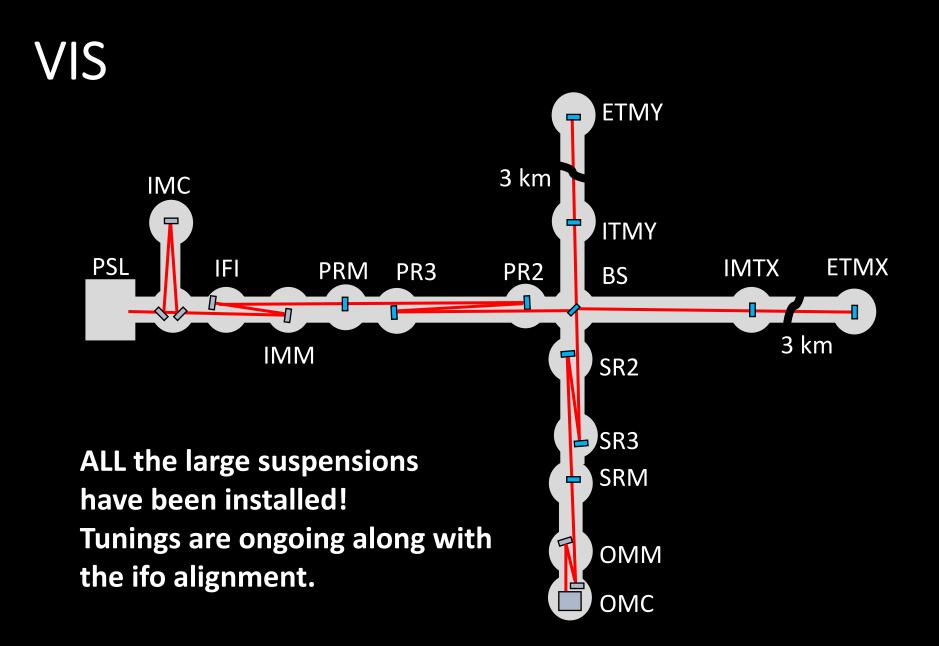


Before going to the current status...



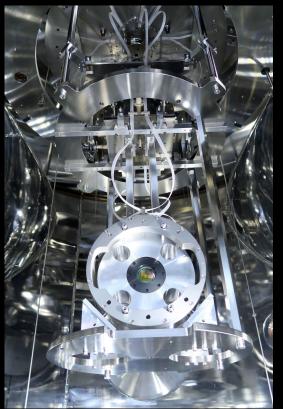


Recent Achievements and News



Most recently, all SRs have been installed!

SRM



SR2



SR3



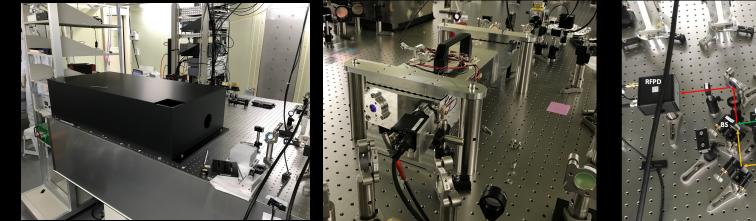
* SRM has a temporary 2-inch mirror

→ See, Burton, Tapia, Fujii, Kozu's talks

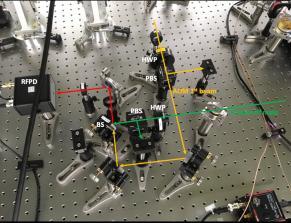
Input Optics

40W laser installed

PMC installed



Mach-Zehnder ifo type modulation system, PM&AM monitor system



- Input mode cleaner was tested with 10W
- Intensity stabilization is being commissioned
- Frequency stabilization (mode cleaner & reference cavity) has been operating since phase1

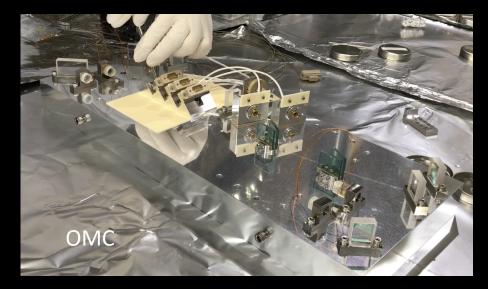
ightarrow See, Nakano's talk

Output Optics



- Output mode cleaner (OMC)
- Output Faraday Isolator (OFI)
- Output mode-matching telescopes (OMMTs) installed!

Nov-Dec 2018



Feb 14th, 2019

Auxiliary Optics

WAB EX



Transmon EX



NAB IY



- WAB 3 of 4 installed! The last one delivered at the X end
- Transmitting monitor (TRANSMON) installed in both of the X and Y ends!

All the sapphire mirrors has installed in Nov 2018

ITMX and ITMY

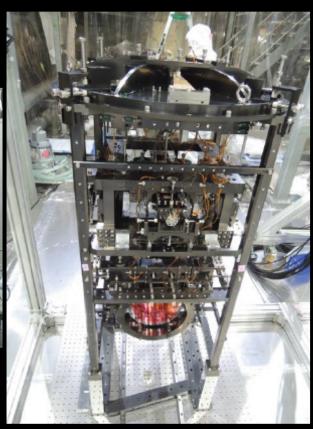
Pictures from K Yamamoto

→ See, Ushiba, Yamada, Fukunaga's talks



Ears were bonded at Toyama U

Bonded mirror is integrated Into the cryo-payload And the type-C suspension at the site

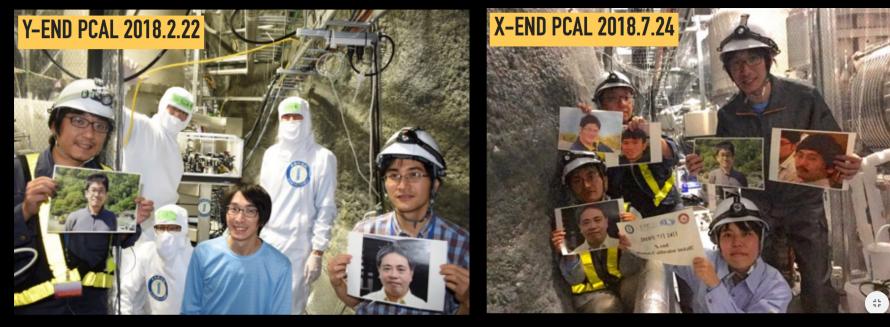


Feb 14th, 2019

CAL

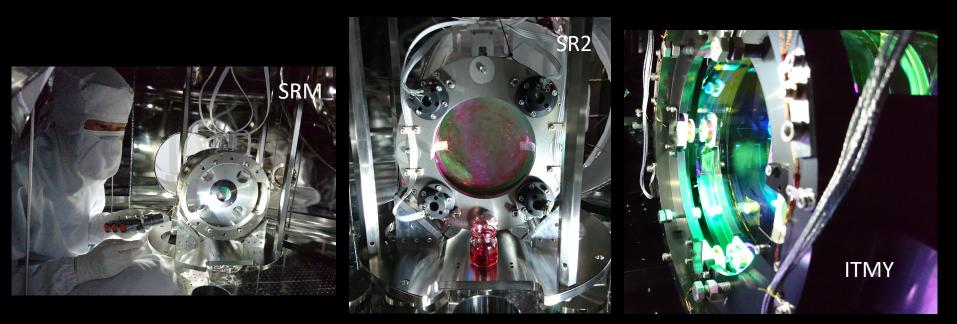
- Photon calibrator modules installed at the both ends
- Calibration pipelines are being constructed

See, T Yamamoto's talk



JGW-D1807705, JGW-G1909581

Mirror



Many mirrors were cleaned with FC before starting the DRMI commissioning

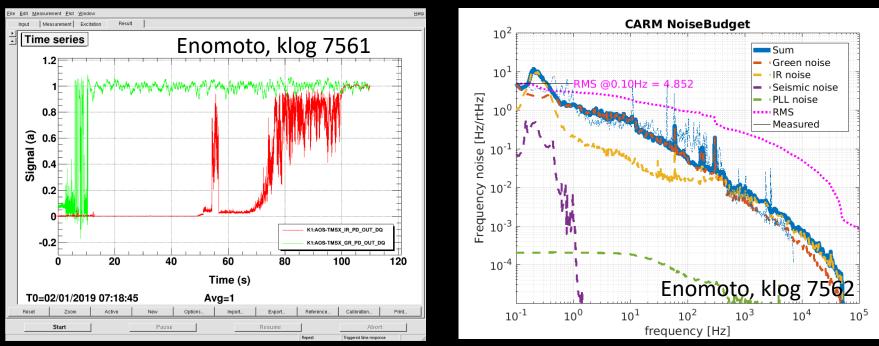
- 3 IMC mirrors as preparation for the high power
- PR3 (condensation), SRM, SR2 (condensation), SR3 (dusts)

Test masses were found to have particles too

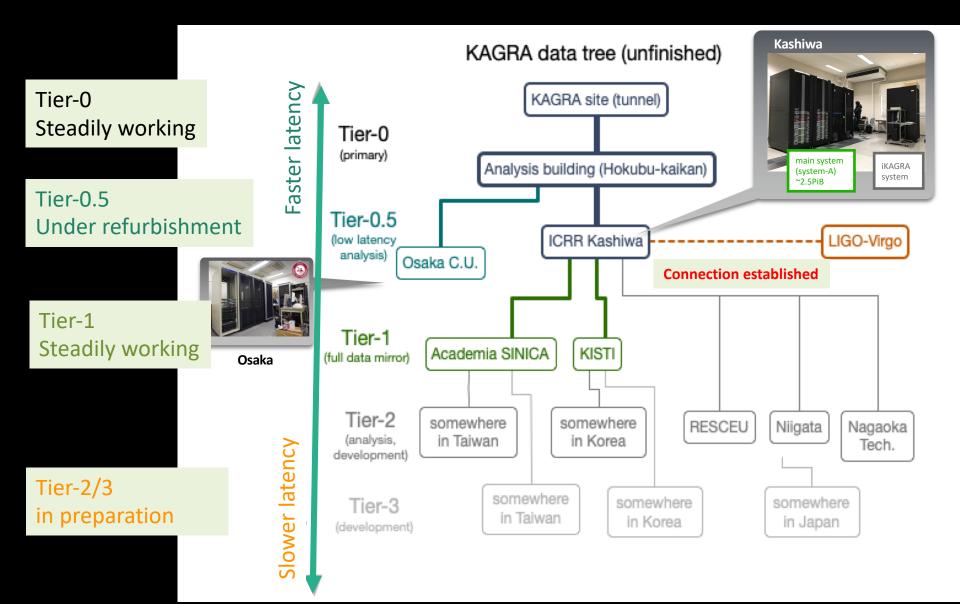
X-arm Locking Test

- X-arm test has completed
 - X-arm locked with the axillary (green) laser, then successfully handed off to the IR laser
 - Noise budgeting

ightarrow See, K Izumi's talk



Data Management



Slie by Kanda-san Also see, [Kagra-cal 00392]

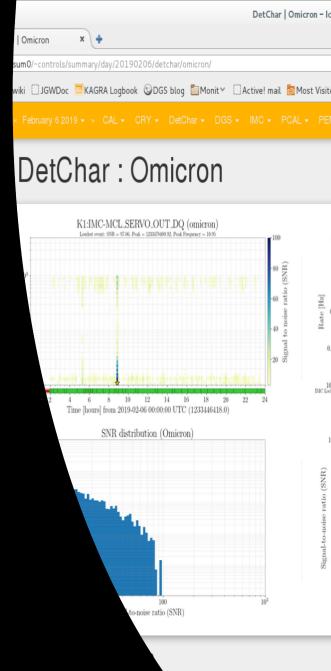
Data Management

Milestones to achieve by the end of Feb:

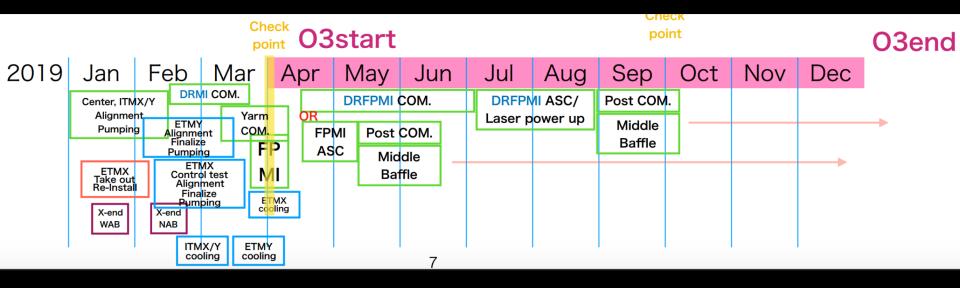
- Low Latency h(t) transfer
 - From KAGRA tunnel to the surface test : OK
 - From KAGRA tunnel to Kashiwa server : Testing
- Low Latency connection with LV
 - The connection had established
 - 6.2 sec from LLO to Kashiwa
 - Status information has been exchanged
- Bulk data connection with LV
 - Implementation of LDR (LIGO Data Replicator) software is on going

Detector Characterization

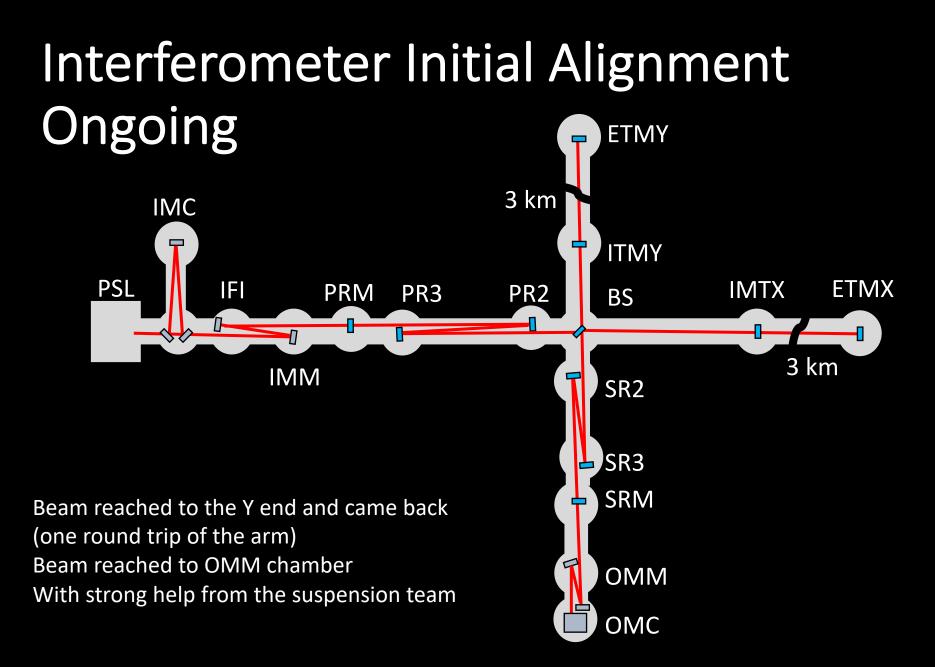
- Had a kick-off meeting with L-V Detchar teams
- A bunch of the Detchar tools were imported from L-V!
 - Summary Pages, Omicron, Bruco, Sus drift mon, ligoDV...
 - Line finding tools, ligoDV-web to be installed
 - iDQ by KGWG (see, Young-min's talk)
- Data quality generation and sharing to be discussed
- Strong collaboration with KGWG



Commissioning Plans and Next IFO Checkpoint



Uchiyama, JGW-E1809209



Prospects of joining O3

- Improves sky coverage (→ see, JGW- G1809082 by Tagoshi)
- Network duty factor
- source parameter estimation (→ see, JGW-G1808212 by Haino)
- Some parameter degeneracy can be resolved with four detectors (e.g. polarization, → see, PRD 98, 022008 by Takeda)

Summary

- KAGRA detector rapidly being integrated towards joining O3
- So far, no major delay happened
- Dual-recycled Michelson ifo commissioning is about to start, followed by the full ifo commissioning and noise hunting
- For O3, KAGRA will contribute for the source localization and the GW polarization studies