

The NS Equation of State: basics

Tidal effects add linearly to point-particle $h(f) = \mathcal{A}e^{i[\psi_{PP}(f) + \psi_T(f)]}$

GW modelling

Inspiral phase: the analytic side

 Effect of electric-static tides successfully included within templates [6pN + 6.5pN leading tail contribution]
Flanagan, Hinderer 08; Hinderer 08;

Damour Nagar 09, Binnington, Poisson 09; Vines, Flanagan, Hinderer 11; Damour, Nagar, Villain 12

- Improved form fo tidal-EOB
- Spin-tidal coupling [6.5 pN]

Damour, Nagar, Villain 12; Bernuzzi, Nagar et al. 12

> Abdelsalhin et al. 18; Forteza et al. 18

• Rotational Love Numbers [6.5 pN]

Pani er al. 15; Landry, Poisson 15

• Magnetic Love numbers [6 pN]

Damour, Nagar 09; Binnington Poisson 09; Yagi 17 Inspiral phase: source of systematics

- *higher-order tidal effects?*
- Lack of complete point-particle knowledge of the waveform up to 5 pN?

Favata 14; Messina et al. 19

O Inclusion of spin-tidal Love number?

Forteza et al. 18

O Rotational Love numbers? Spin magnitudes for NS?

Inspiral phase: extra effects to take into account?

O may be relevant for signals at high-frequencies

• Dynamical tides, aka non constant Love numbers

Maselli et al. 12; Steinhoff, Hinderer et al.16; Schmidt, Hinderer 19

• Resonant oscillations, for circular and hyperbolic encounters

Ho, Lai 99; Flanagan Racine 07; Tsang 13; Yu, Weinberg 16

• Non-linear couplings of higher modes (pressure, gradient modes)

Weinberg 16; Essick, Vitale, Weinberg 16

GW modelling: new physics

Inspiral phase: dynamical tides

Hinderer et al. 16



Inspiral phase: radius constraints

- **O** LOFT, eXTP, NICER: measurement of $R_{\rm NS}$ at the order of ~%
- **O** LVC for GW170817 measurement of $R_{\rm NS} \sim 10\%$

GW170818 - LVC 18

- number of events should beat this limit
- 3G single observation aim to be comparable and better than future satellites, $R_{\rm NS}$ ~ 100 meters
- multi-messenger pipeline to join GW and EM signals with the stellar EoS as a goal?

Fasano et al. 19

Multi-messenger

Multi-messenger on the AdS



Post-merger phase:

• Signals depend noticeably on the NS structure

Bauswein et al. 19; Most et al.19

• Highly-rewarded challenge: possible formation of new-states of matter

BNS-spectroscopy

- *high-frequency content: 3G goal*
- non linear evolution: status of numerical simulations?
- *full-understanding of physical process?*

BNS - Spectroscopy

Takami et al. 14 t [ms] 10 -515 20 5 0 $h_+ \! \times \! 10^{22}$ 5 0 -5 $\log[2\widetilde{h}(f) f^{1/2}] [Hz^{-1/2}, 50 Mpc]$ APR4 \mathbf{f}_{2} -22 GNH3 -22.5 adLIGO 12 -23 f₃ EТ -23.5 2 3 5 0 1 4 f[kHz]



Bauswein et al. 19

BNS - Spectroscopy

- correlation between frequency lines and NS properties?
 - phenomenological relations between peaks and radius/ compactness of the cold EoS Bauswein, Janka 2012; Stergioulas et al.11;

Takami et al. 15

• correlation of the frequency peak with the Love number

Bernuzzi et al. 15

• post-merger frequencies evolve in time. How much SNR needed to resolve them?

Neutron stars

 $\mathcal{C} \in [0.1 \div 0.2]$

 $k_2 \neq 0$

GW170817

constrain the NS equation of state

Black holes $\mathcal{C} = 1/2$ $k_2 = 0$

Pani, Gualtieri, Maselli, Ferrari PRD 92 '15



Maselli + grqc: 181103689 '18