# LVK Communications

Susanne Milde, AEI/Milde Science Comms, June 30th 2022

## AGENDA

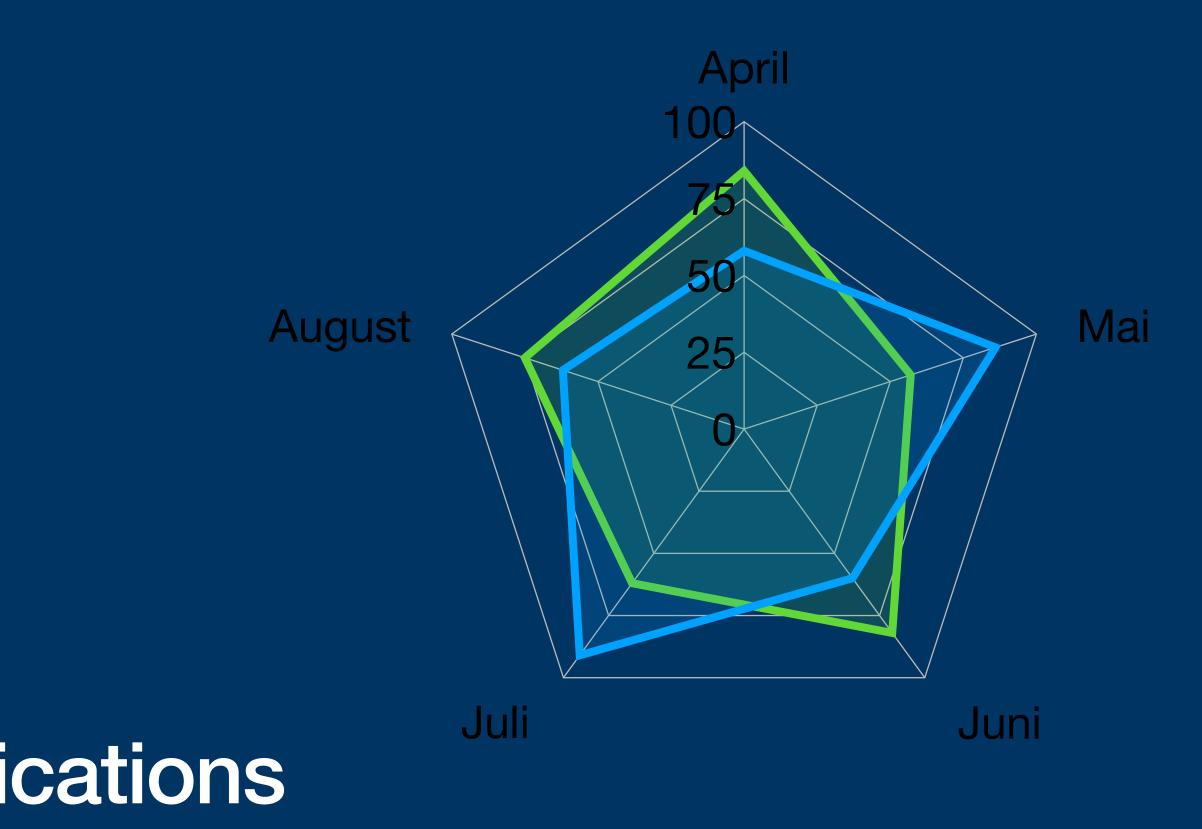
Science Communications: Why, how and what? Tools and essential steps

brainstorming

Science Communications for LVK - exchange of information and experiences,

# Science Communications

Outreach Education Media actives Lobby work Internal communications 



# Local, regional, national

# International

## **SCIENCE COMMUNICATIONS – WHY?**

Crucial if we want to be successful in

- getting recognized,
- getting funded,
- attracting the best researchers to the field

## Additionally

- Inform the public about LVK research

Increase the public knowledge and understanding of (fundamental) research

## The Process

## **Definition of basics:** Goals, target groups, key messages

## **Development of an overarching strategy**

## **Realisation including coordination**

## **Evaluation and adjustment**







# MEDIA: OFF- AND ONLINE

"All the News There's Fit to Prior"

### The New Hork Times

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NEW YORK, TUESDAY, OCTOBERS, 2020

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Venice Turns Cameras on Tourists to Count and Corral Big Crowds Streets and Storefronts Empty.

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# SOCIAL MEDIA

### **New LISA Instagram Account**

Photo post + caption with further information, hashtags & links to learn more Categories: LISA Sources, LISA Science Goals, LISA Signals, LISA Instrumentation



LISA SOURCE HOLE BINARY

THE STRONGEST SOURCES OF GRAVITATIONAL WAVES IN THE UNIVERSE. FROM THEM, WE WILL LEARN MORE ABOUT GALACTIC FORMATION AND STRUCTURE.

Image Credit: NASA

## SUPERMASSIVE BLACK



# LISA SCIENCE GOA EXPANSION OF THE UNIVERSE and the second second

LISA WILL ALLOW US TO DETERMINE THE HUBBLE CONSTANT WITH UNPRECEDENTED PRECISION AND SO, THE EXPANSION RATE OF THE UNIVERSE.





MAX PLANCK INSTITUTE FOR GRAVITATIONAL PHYSICS (Albert Einstein Institute)

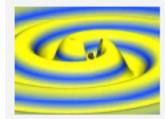
INSTITUTE | RESEARCH | CAREER | NEWS ROOM

Home > Research

### Research at the Albert Einstein Institute 0

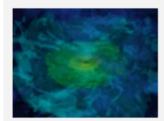
The Max Planck Institute for Gravitational Physics (Albert Einstein Institute/AEI) is one of the world's leading centers for gravitational physics, with a unique breadth and depth of its approach to the subject. Scientists at the AEI focus on all aspects of Einstein's theory of general relativity. The research topics range from the theoretical, observational and experimental aspects of cravitational-wave physics and astrophysics, to unification of general relativity and guantum mechanics, all the way to geometrical and analytical aspects of the theory. The AEI has two branches, one in Potsdam, and one in Hannover where it is closely related to the Leibniz Universität Hannover.

### Departments



### Astrophysical and Cosmological Relativity (Potsdam) Director: Prof. Dr. Alessandra Bucnanno

The Astrophysical and Cosmological Relativity department develops accurate analytical and numerical models of gravitational-wave sources, and uses them to analyse data, improving our ability to extract unique astrophysical and cosmological information from the observed signals, and testing Einstein's theory of general relativity. > more



### Computational Relativistic Astrophysics (Potsdam) Director: Prof. Dr. Masaru Shibata

Research in the department Computational Relativistic Astrophysics covers mergers of binary neutror stars and mixed binaries as well as stellar core collapse that form black holes. The department also focuses on studying more fundamental aspects of General Relativity using numerical tools. > more



### Laser Interferometry and Gravitational Wave Astronomy (Hannover)

Director: Prof. Dr. Karsten Danzmann

The department Laser Interferometry and Gravitational Wave Astronomy focuses on the development of gravitational wave detectors on Earth as well as in Space. This comprises also a full range of supporting laboratory experiments in guantum optics and laser physics. > more

### HUMANS OF LIGO

### LATEST POSTS



## **ONLINE INFORMATION**

"Geodetic" redirects here. For other uses, see Geodetic (disambiguation).

This article has multiple issues. Please help to improve it or discuss these [hide] issues on the talk page. (Learn how and when to remove these template messages)

- This article includes a list of general references, but it remains largely unverified because it lacks sufficient corresponding inline citations. (February 2009)
- This article needs additional citations for verification. (December 2018)
- This article's lead section may not adequately summarize its contents. (June 2018)

Geodesy (/dʒi:'odɪsi/<sup>[1]</sup> jee-OD-ih-see) is the Earth science of accurately measuring and understanding Earth's geometric shape, orientation in space, and gravitational field.<sup>[2]</sup> The field also incorporates studies of how these properties change over time and equivalent measurements for other planets (known as planetary geodesy). Geodynamical phenomena include crustal motion, tides and polar motion, which can be studied by designing global and national control networks, applying space and terrestrial techniques

SEARCH



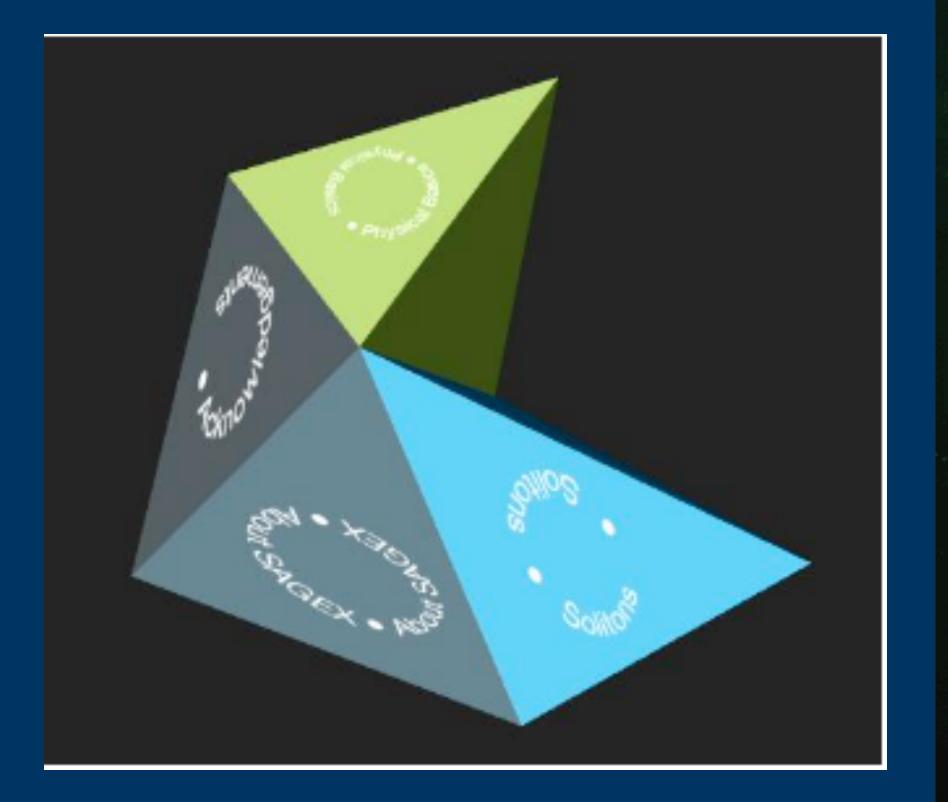
## **Outreach Activities**

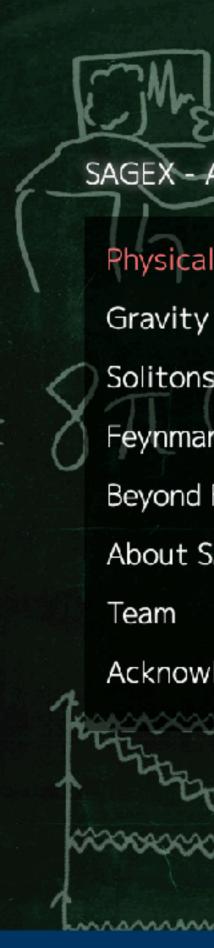






# https://exhibition.sagex.org/#/exhibition-hub





### SAGEX - AT THE FRONTIER OF PHYSICS

### Physical Basics

- Solitons
- Feynman Diagrams
- **Beyond Feynman Diagrams**
- About SAGEX
- Acknowledgements



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# Materials



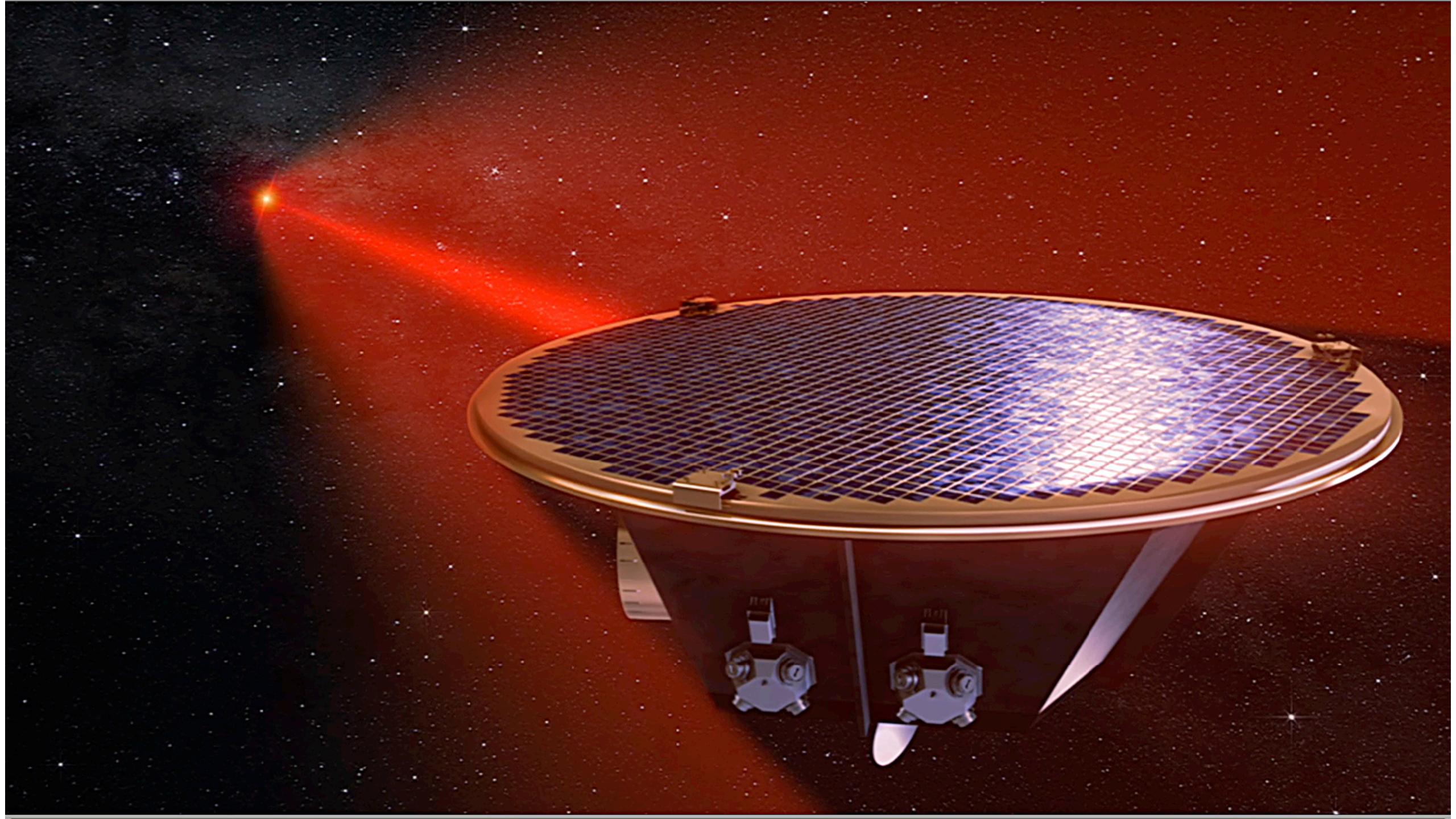
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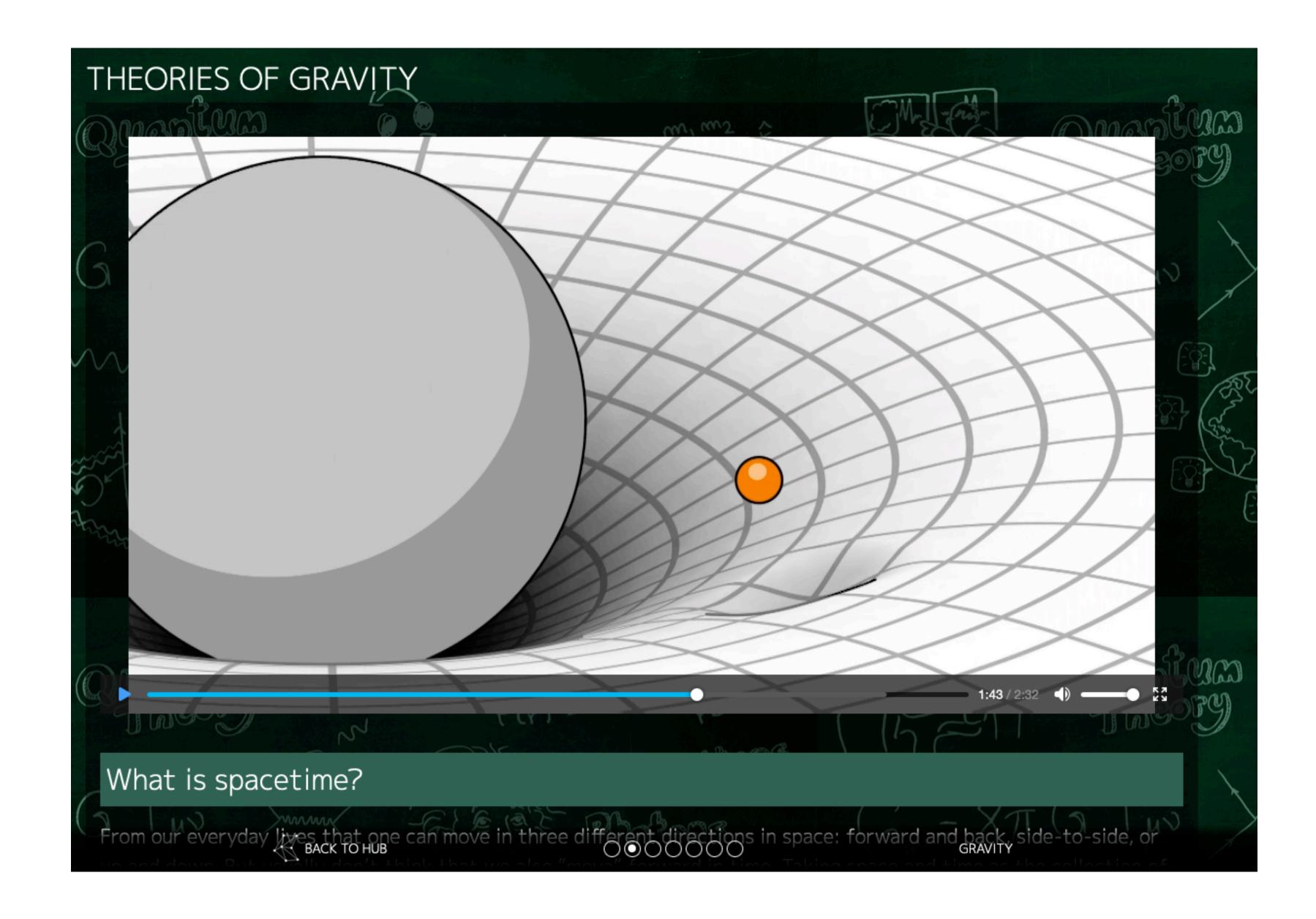
## Texts, images, animations

## Skilled experts

# Content

based on —> goal, target groups, key messages





### WHY DOES THE SUN SHINE?

The Sun is an ordinary middle-aged but also in the variety of elements the iron in our blood, and rare and star, one of 100+ billion stars in around us - the oxygen we breathe, precious elements such as gold, our Galaxy alone. The Sun consists the carbon that is the basis of life as platinum or uranium. largely of hydrogen (a nucleus made well as fossil fuel. of one proton) and helium (two proof one proton) and helium (two pro-tons, two neutrons). These, the light-est elements, are thought to have been made in primordial process-es; just minutes after the Big Bang, nearly 14 billion years ago. Such primordial gas gradually condensed through dravity into massive indense ther hand eventually condensed ther hand eventually condense

In stars, tremendous pressure and material far into space, heat drive the nuclear-fusion reac-tions of hydrogen nuclei to produce Within the supernova, violent nucle- evolution more helium, as well as carbon, oxy- ar processes between transient exgen, and other nuclei, with the re- otic nuclei lead to the synthesis of lease of huge amounts of energy - more of the important elements that on the MPG/ESO 2.2-metre telescope at ESO's La Sille the 'nuclear binding energy'. We see shape our life ('nucleosynthesis'). Observatory in Chile shows the bright open star cluster NGC 25%2. Between the bright stars, far avery in the the manifestation of this in the sun- Such supernovae make the calolum background of the image, many remote galaxies can be shine and warmth that sustains life, in our bones, the oxygen we breathe, seen some with clearly spiral shapes.

through gravity into massive incan- other hand, eventually collapse under Nuclear astrophysicists investigate descent balls of gas – the first stars, their own gravity, and most of them the processes underlying the cre-explode as a supernova, throwing out ation of the elements and their influ-

ence on broader cosmic phenomena in gas, stars, galaxies, and in their

### **THE ELEMENTS** AND THEIR ISOTOPES: MADE BY NUCLEAR REACTIONS

In atomic nuclei, protons and neutrons are bound together by the strong nuclear force against the electrostatic repulsion of the electric charge of protons. A nucleus is very compact, and ten thousand times smaller than the electron cloud that determines the size of the atom. The charge of the electron cloud determines the characteristic chemical properties of each element. The different number of neutrons that can be bound to the same number of protons make up the variety of isotopes, and these determine the characteristics of nuclear reactions. These reactions re-arrange the mix of protons and neutrons, thus creating new isotopes from existing ones. In cosmic environments, nuclear reactions often involve unstable and rare isotopes. Thus, from the primordial elements hydrogen and helium, elements such as carbon, oxygen, iron, and gold, and all their isotopes, are made.

A three-dimensional periodic table showing the various isotopes of the elements along the third axis.

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