**AdVirgo Suspensions and Mirrors** 





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#### Introduction Seismic Noise on Earth

- Seismic noise has both natural and human origins and can vary by few orders of magnitude from site to site.
- Al ground motion displacement spectra observed worldwide share some common characteristics: they have essentially the same amplitude in all three orthogonal space directions and they exhibit a low pass behavior that follows the empirical law for f > 0.1 Hz





#### Introduction Harmonic Oscillators as Mechanical filters

At frequencies higher than the oscillator resonance, the transfer function of an harmonic oscillator is equivalent to a second-order low pass filter.



#### **Transfer Function**

## **AdVirgo Superattenuator**



The AdVirgo superattenuator (SA) is a complex mechanical device capable of providing more than 10 orders of magnitude of passive seismic isolation in all six degrees of freedom above a few Hz

• The SA is a passive mechanical system constituted by a 5 stage pendulum supported by a 3-leg elastic pre-isolator called inverted pendulum (IP).

• All the normal mode resonance frequencies of the SA are kept below 2 Hz.

• The SA mechanical structure, consists of three fundamental parts: the inverted pendulum, the chain of standard filters, the payload.

 Mechanical design for AdVirgo is essentially the same of Virgo except for the payload.



### AdVirgo Superattenuator The inverted pendulum



• A low frequency pre-isolator constituted of three 6 m-long hollow legs, each one connected to the ground through a flexible joint and supporting an interconnecting structure (the top ring) on its top.

• The structure horizontal normal modes are tuned at about 30-40 mHz.







Since the system is very soft, it requires very low forces to be moved:

for f<<f0  $F \simeq M \omega_0^2 x$ 

• The top ring is a mechanical support for an additional seismic filter, called filter 0, similar to those used in the chain.

• The filter 0 is equipped with a set of sensors and actuators, placed in a pinwheel configuration, that are used to actively damp the IP resonance modes.

#### AdVirgo Superattenuator Vertical Attenuation: Standard filters



The first four pendulum stages of the SA are denominated Standard Filters (SFs).

The SF is essentially a rigid steel cylinder supporting a set of maraging steel cantilevered triangular blades clamped along the outer surface of the filter body.

A magnetic anti-spring system, assembled on each filter, is designed to reduce its fundamental vertical frequency from about 1.5 Hz down below 0.5 Hz.



Magnetic antispring working principle

#### AdVirgo Superattenuator Vertical attenuation: Blades



- All the maraging steel blades have a thickness of 3.5 mm, a length of 385.5 mm, while the width of the triangular base changes according with the load to be supported.
- The number of blades ranges from 12 (in the first filter of the chain) to 4 (in the filter 7) according to the suspended load. A total of 52 blades is needed for a long tower.
- The load M depends by the base width b, by the thickness t and length I with this law

$$M = \frac{Ebt^3}{12R_cgl}$$

#### AdVirgo Superattenuator The payload





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## AdVirgo Superattenuator SA control system setup



21 Motors

• 5 F0 Coils 6 F7 Coils

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#### AdVirgo Superattenuator Hierarchical Control



The control system has been designed using a hierarchical strategy regulated by the dynamic range of the actuators.

- 23 Coil Magnet Actuators in 3 points (actuation stages) of the SA:
- Filter 0:

Large displacements (hundreds of microns) for f<10 mHz.

- Filter 7 + Marionette:

Small payload displacements (1 micron) in the 10 mHz<f<1 Hz band.

- Filter 7 + Mirror:

Small residual mirror displacements (a few nm), for f>1 Hz

- >20 Local Sensors
- Accelerometers
- 4 Displacements Sensors (LVDTs) on Filter 0
- 6 Displacements Sensors on Filter 0
- 6 Displacements Sensors on Filter 7
- Optical Readout of Marionette and Mirror Position

For 50 mHz<f<5 Hz

For f>1 Hz

#### AdVirgo Superattenuator SA Sensors

• There is a total of 5 Accelerometer (Accs) installed on the suspension F0 of 2 different types with sensitivity of about  $3*10^{-10}$  m/s<sup>2</sup>/sqrt(Hz) for f < 3 Hz

• There are 18 LVDTs installed on long tower suspensions of 3 different types with a sensitivity of about  $10^{-8}$  m/sqrt(Hz) for f > 0.1 Hz

• All the LVDTs are operated using a digital demodulation scheme at 320 kHz sampling frequency





#### AdVirgo Superattenuator SA Control system hardware





Visit of Swedish Academy

#### AdVirgo Superattenuator SA control sytem hardware

#### Electronics Design based on Texas Instruments DSP

- TMS320C6678
  - Eight TMS320C66x DSP Core Subsystems
  - 320 GMAC/160 GFLOP @ 1.25GHz
  - Four Lanes of SRIO 2.1 5 Gbaud Per Lane Full Duplex
  - Two Lanes PCIe Gen2 5 Gbaud Per Lane Full Duplex
  - Ethernet MAC Subsystem Two SGMII Ports w/ 10/100/1000 Mbps operation
  - 64-Bit DDR3 Interface (DDR3-1600)

### Computing power of a high-end GPU but extremely energy efficient and specifically designed for hard real-time applications

	Platform	Effective Time to complete 1024 complex to complex FFT (single precision) μs	Power (Watts)	Energy per FFT (µJ)
GPU	nVidia Tesla C2070	0.16	225	36
GPU	nVidia Tesla C1060	0.3	188	56.4
GPP	Intel Xeon Core Duo @ 3 GHz	1.8	95	171
GPP	Intel Nehalam Quad Core @ 3.2 GHz	1.2	130	156
DSP	TI C6678 @ 1.2 GHz	0.86	10	8.6



# AdVirgo Superattenuator SA control system hardware • A total of 14 boards, each one equipped with an 8-core TMS320C6678 DSP, are connected to each

long suspension:



#### AdVirgo Superattenuator Control system software

#### SA control is an extremely complex system:

- 131 DSP boards are installed on BPC, BS, IB, MC, PR, NI, NE, WI, WE, SR, OB
- 185 control code files are running at the same time on the DSP cores at 10 kHz (IP, F7, LC controls), 40 kHz
- (Global signals oversampling) and 320 kHz (Digital demodulation of sensors)

#### **O2 SOFTWARE MAP**

SA	BOARD IP	CONNECTED DEVICES	SOFTWARE RUNNING (Core4, 10 kHz)	SOFTWARE RUNNING (Core1, 320 kHz)
BPC	172.16.2.104	PSD	/virgoDev/Sa/DSPCode Adv/BPC/BPC PSD	
BPC	172.16.2.141	PIEZO	/virgoDev/Sa/DSPCode_Adv/BPC/BPC_CD	
Sa_BS	172.16.2.62	BR LVDTs	/virgoDev/Sa/DSPCode_Adv/BS/LVDT/BS_MASTER	/virgoDev/Sa/DSPCode_Adv/BS/LVDT/BS_BR_LVDT_Demod
Sa BS	172.16.2.53	F0 LVDTs	/virgoDev/Sa/DSPCode_Adv/BS/LVDT/BS_LVDT_HG_SRIO	/virgoDev/Sa/DSPCode_Adv/BS/LVDT/BS_LVDT_HG2
Sa BS	172.16.2.32	F0 VAccs	/virgoDev/Sa/DSPCode_Adv/BS/Accs/BS_vAcc_LQG	/virgoDev/Sa/DSPCode_Adv/BS/Accs/BS_vAcc_Demod
Sa BS	172.16.2.33	F0 Coils	/virgoDev/Sa/DSPCode_Adv/BS/InertialDamping/BS_ID_Diag	
Sa_BS	172.16.2.133	F0 HAccs	/virgoDev/Sa/DSPCode_Adv/BS/Accs/BS_Acc_LQG	/virgoDev/Sa/DSPCode_Adv/BS/Accs/BS_Acc_Demod
Sa_BS	172.16.2.52	F1-F7 VLVDTs	/virgoDev/Sa/DSPCode_Adv/BS/LVDT/BS_VLVDT_SRIO	/virgoDev/Sa/DSPCode_Adv/BS/LVDT/BS_VLVDT
Sc_BS	172.16.2.80	PSD	/virgoDev/Sa/DSPCode_Adv/BS/LC/BS_PSDf	
Sc BS	172.16.2.108	PSD	/virgoDev/Sa/DSPCode_Adv/BS/LC/BS_PSDm	
SC_BS	172.16.2.110	PSD	/virgoDev/Sa/DSPCode_Adv/BS/LC/BS_PSDt	
Sc BS	172.16.2.84	PSD	/virgoDev/Sa/DSPCode_Adv/BS/LC/BS_PSDi	
SC_BS	172.16.2.181	MIR, MAR Coils	/virgoDev/Sa/DSPCode_Adv/BS/LC/BS_Mir	
SC BS	172.16.2.179	MAR Coils	/virgoDev/Sa/DSPCode_Adv/BS/LC/BS_Mar	
Sc BS	172.16.2.139	F7 LVDT	/virgoDev/Sa/DSPCode_Adv/BS/LVDT/BS_F7_LVDT	/virgoDev/Sa/DSPCode_Adv/BS/LVDT/BS_F7_LVDT_Demod
Sc_BS	172.16.2.120	F7 Coils	/virgoDev/Sa/DSPCode_Adv/BS/F7/BS_F7_CD	
Sa IB	172.16.2.28	BR LVDTs	/virgoDev/Sa/DSPCode Adv/IB/LVDT/IB MASTER	/virgoDev/Sa/DSPCode Adv/IB/LVDT/IB BR LVDT Demod
SaliB	172.16.2.130	F0, F4, F7 LVDTs	/virgoDev/Sa/DSPCode_Adv/IB/LVDT/IB_LVDT	/virgoDev/Sa/DSPCode_Adv/IB/LVDT/IB_LVDT_Demod
SaliB	172.16.2.9	F0 VAccs	/vireoDev/Sa/DSPCode_Adv/IB/Accs/IB_vAcc_LOG	/vireoDev/Sa/DSPCode_Adv/IB/Accs/IB_vAcc_Demod
Sa IB	172.16.2.121	F0 Coils	/virgoDev/Sa/DSPCode Adv/IB/InertialDamping/IB ID Diag	
SallB	172.16.2.23	F0 HAccs	/virgoDev/Sa/DSPCode Adv/IB/Accs/IB Acc LQG	/virgoDev/Sa/DSPCode Adv/IB/Accs/IB Acc Demod
Sc IB	172.16.2.118	PSD	/virgoDev/Sa/DSPCode Adv/IB/LC/IB PSDr	
Sc IB	172.16.2.86	PSD	/virgoDev/Sa/DSPCode_Adv/IB/LC/IB_PSDi	
Sc IB	172.16.2.107	PSD	/virgoDev/Sa/DSPCode_Adv/IB/LC/IB_PSDt	
SC IB	172.16.2.173	MAR Coils	/virgoDev/Sa/DSPCode Adv/IB/LC/IB Mar1	
Sc_IB	172.16.2.174	MAR Coils	/virgoDev/Sa/DSPCode_Adv/IB/LC/IB_Mar2	
Sa_MC	172.16.2.128	BR LVDTs	/virgoDev/Sa/DSPCode_Adv/MC/LVDT/MC_MASTER	/virgoDev/Sa/DSPCode_Adv/MC/LVDT/MC_BR_LVDT_Demod
Sa_MC	172.16.2.51	F0, F4, F7 LVDTs	/virgoDev/Sa/DSPCode_Adv/MC/LVDT/MC_LVDT	/virgoDev/Sa/DSPCode_Adv/MC/LVDT/MC_LVDT_Demod
Sa_MC	172.16.2.158	F0 VAccs	/virgoDev/Sa/DSPCode_Adv/MC/Accs/MC_vAcc_LQG	/virgoDev/Sa/DSPCode_Adv/MC/Accs/MC_vAcc_Demod
Sa_MC	172.16.2.103	F0 Coils	/virgoDev/Sa/DSPCode_Adv/MC/InertialDamping/MC_ID_Diag	
Sa_MC	172.16.2.14	F0 HAccs	/virgoDev/Sa/DSPCode_Adv/MC/Accs/MC_Acc_LQG	/virgoDev/Sa/DSPCode_Adv/MC/Accs/MC_Acc_Demod
Sa_MC	172.16.2.150	PIEZO	/virgoDev/Sa/DSPCode_Adv/MC/Tilt/Piezo_Test	
SC_MC	172.16.2.101	PSD	/virgoDev/Sa/DSPCode_Adv/MC/LC/MC_PSDf	
Sc_MC	172.16.2.168	PSD	/virgoDev/Sa/DSPCode_Adv/MC/LC/MC_PSDi	
Sc_MC	172.16.2.88	PSD	/virgoDev/Sa/DSPCode_Adv/MC/LC/MC_PSDTf	
SC_MC	172.16.2.109	PSD	/virgoDev/Sa/DSPCode_Adv/MC/LC/MC_PSDTi	
SC_MC	172.16.2.171	MAR Coils	/virgoDev/Sa/DSPCode_Adv/MC/LC/MC_Mar1	
SC_MC	172.16.2.172	MAR Coils	/virgoDev/Sa/DSPCode_Adv/MC/LC/MC_Mar2	
Sc_MC	172.16.2.176	MIR Coils	/virgoDev/Sa/DSPCode_Adv/MC/LC/MC_Mir	
Sa_NE	172.16.2.37	BR LVDTs	/virgoDev/Sa/DSPCode_Adv/NE/LVDT/NE_MASTER	/virgoDev/Sa/DSPCode_Adv/NE/LVDT/NE_BR_LVDT_Demod
Sa_NE	172.16.2.40	F0 LVDTs	/virgoDev/Sa/DSPCode_Adv/NE/LVDT/NE_LVDT	/virgoDev/Sa/DSPCode_Adv/NE/LVDT/NE_LVDT_Demod

## **AdVirgo Mirrors**

- FP Mirrors made of SiO2, 350 mm of diameter, 200 mm wide, surface roughness < 10<sup>-8</sup> m.
- Very strong requirement on the amount of light lost per round trip: < 0.008 %
- Monolithic suspensions: made with extremely thin SiO2 fibers (400 µm of diameter) to suspend mirrors of about 42 kg.



## The 3 ingredients of a GW Mirror



**Credit: Jerome Degallaix** 

## Coating

- Picture of the Ion Beam Sputtering (IBS) custom machine in the LMA clean room, the largest in the world.
- All the coatings for the test masses of the GW detectors have been made in LMA IBS (Virgo – LIGO - Kagra)



## **Future AdVirgo Mirrors**

- To reduce thermal noise, in AdVirgo+ phase II, scheduled to start after the end of O4, mirrors of 55 cm (+ 60%) in diameter and about 105 kg in weight (x 2.5) will be used on the end towers of the Fabry-Perot cavities.
- The larger mirrors will induce many modifications on the coating process (cleaning system, larger uniform coatings, annealing), on the manipulation devices and also on the metrology benches (stronger and larger sample-holders).



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