

Learning to detect continuous gravitational waves

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Machine Learning in GW search: G2net next challenges



Universitat
de les Illes Balears

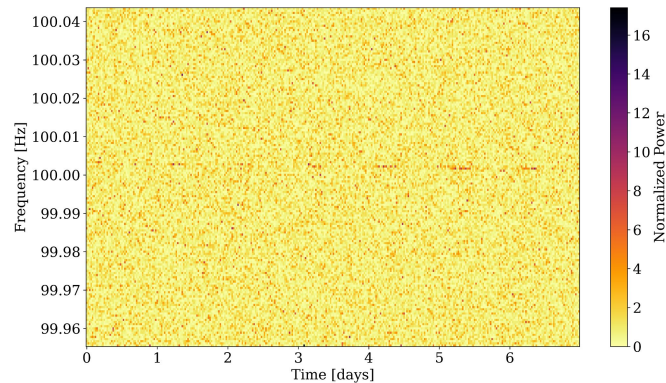
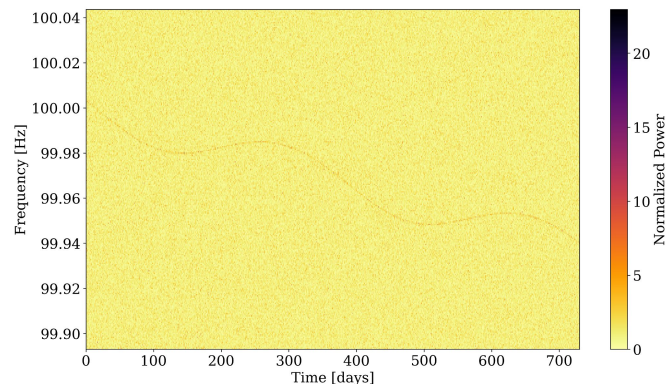


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Introducing continuous gravitational waves (CWs)

Continuous gravitational waves (CWs):

- Long-standing gravitational waves (months - years).
- No detection to date!
- Expected sources: rapidly-spinning neutron stars (NS).
 - Lots of interesting physics to learn from them.
 - Probe the EM-quiet population of NS.
- Typical CW amplitudes: 2-3 orders of magnitude smaller than that of a CBC signal.
 - Peak strain of GW150914: $O(10^{-21})$
 - CW amplitude upper limit on O3: $O(10^{-25})$
- Simple signal model:
 - Amplitude is constant for the duration of the run
 - Quasi-monochromatic signal.
 - Doppler modulation induced by the Earth's motion.

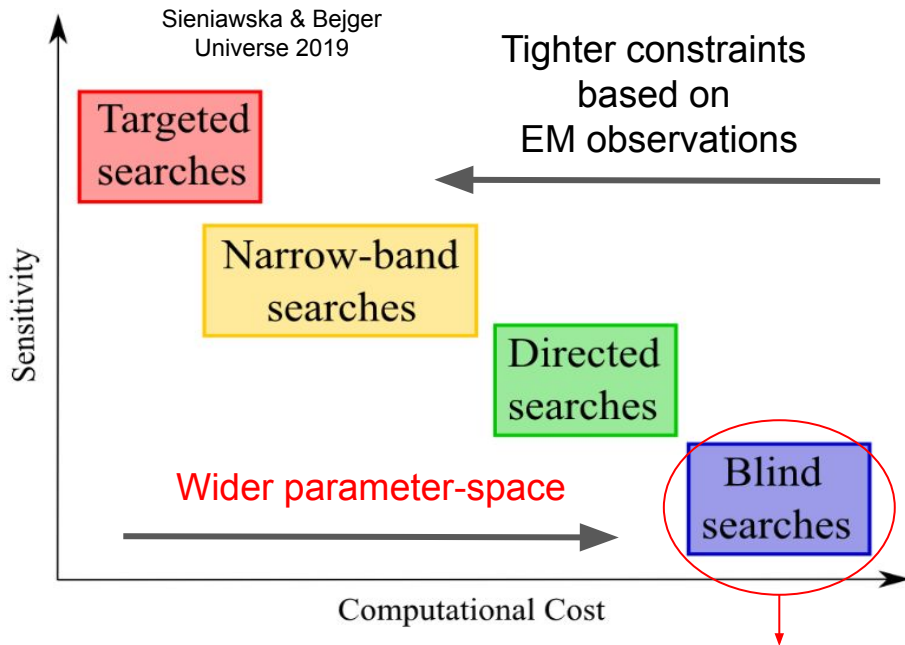


Searching for CWs

CW searches are classified according to what information goes into them:

- Targeted:
Focus on a specific pulsar, phase-lock GW radiation to EM emission.
- Narrow-band / directed:
Focus on a specific sky position, GW emission independent from EM.
- Blind:
Look for signals from unknown sources with a specific Doppler modulation.

Most non-targeted searches are source agnostic: no specific property is required from a source insofar they emit a quasi-monochromatic signal.



First trade-off:

We'll focus on these ones!

- The less we know about our sources, the wider the parameter-space to cover.
- Relying on EM observations may drive us away from the GW-loud NS population.

Searching for CWs

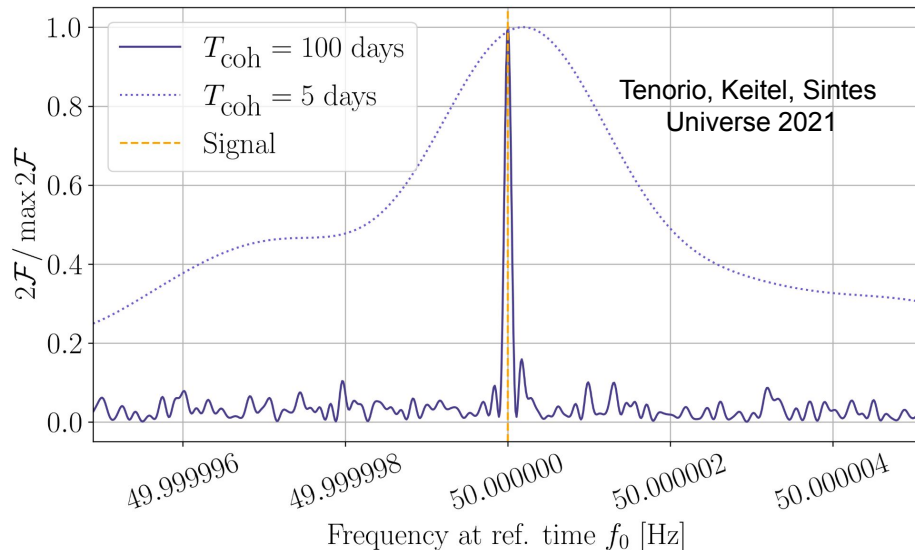
CW signals are significantly longer than CBC signals (months – years vs. seconds).

Matched-filtering (standard tool) is **unfeasible** for wide parameter-space searches.

- Long signal → Huge template banks.

“Solution”: Hierarchical approach

- Use sub-optimal methods to sweep wide parameter-space regions.
 - Impose a looser signal model, reducing the required parameter-space resolution.
- Follow-up interesting candidates with more sensitive methods.



Second trade-off:

- Non-optimal methods are less sensitive than matched-filtering for a given model.
- Non-optimal methods perform better than matched-filtering at a fixed computing cost.

State of the art of CW searches

Less candidates



Lower computing cost

Most searches are structured as follows:



Main search stage

Evaluate a detection statistic over the parameter-space. Accounts for most of the computing cost!

Post-processing

Reduce the number of candidates to take care of using some form of clustering.

Application of vetoes

Discard obvious candidates produced by instrumental artifacts.

Follow-up

Re-analyze surviving candidates using a more sensitive method.

Typical searches run in **a few million CPU hours** on a supercomputer [Abbott+ PRD 100, 024004 (2019)].

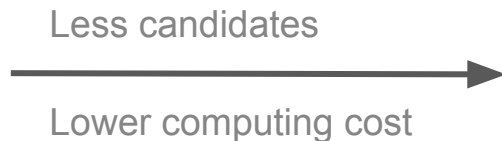
GPU acceleration allows for deeper searches and more diverse signals [Wette+ PRD 103, 083020 (2021)]

[Covas & Sintes PRD 99, 124019 (2019),
Keitel & Ashton CQG. 35 (2018), 205003].

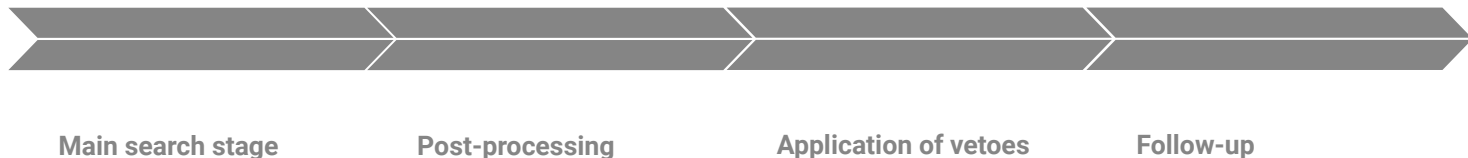
Broadest and deepest searches to date run on the volunteer-distributed computing platform **Einstein@Home**, consuming **~100 M CPU hours** [Steltner+ ApJ 909 79 (2021)].

Extensive manual tuning, however, tends to be **required for a good performance**.

State of the art of CW searches



Most searches are structured as follows:



Can we use machine learning to simplify this kerfuffle?

Typical searches run in a few million CPU hours on a supercomputer [Abbott+ PRD 100, 024004 (2019)].

GPU acceleration allows for deeper searches [Wette+ PRD 103, 083020 (2021)] and more diverse signals

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Machine learning for the search of CWs

Most popular ML application: post-processing the results of a main search.

*Applied to real-data searches!

- *Identify signal tracks on a spectrogram using CNNs [Bayley+ PRD (2020)].
- Identify excess power using CNNs [Yamamoto+ PRD (2021), arXiv:2206.00882].
- Classification of candidate distributions using CNNs [Morawski+ Mach. Learn.: Sci. Technol. (2020)].
- Image-segmentation-based clustering using CNNs [Beheshtipour+ PRD (2020, 2021)].
- *Unsupervised clustering [Singh+ PRD (2017), Tenorio+ PRD (2021), Steltner+ arXiv:2207.14286].

To date, no “machine-learning method” has been used as a main stage in a CW search.

- Only proposal in this front: CNNs on Fourier transforms of the data [Dreissigacker+ PRD (2019, 2020)].

Can we design a machine learning pipeline that...

1. Does not require too much data pre-processing?
2. Is able to compete with current wide parameter-space searches?
3. Does not require 100 M CPU hours to run?

Introducing Kaggle competitions

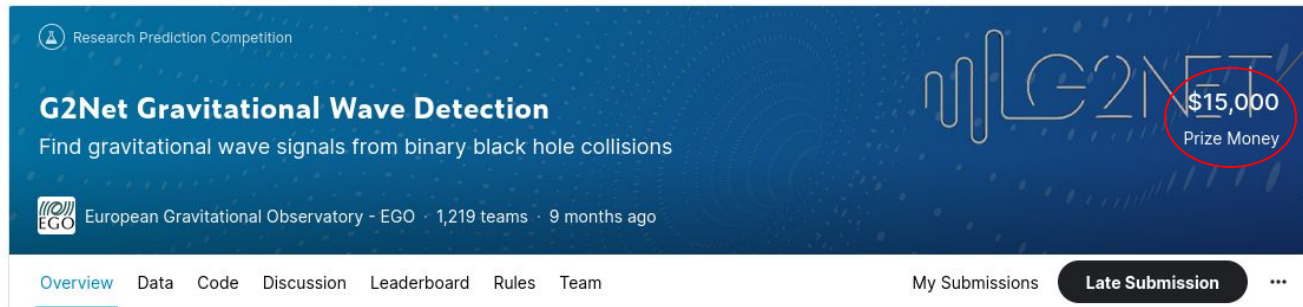


Kaggle: Online platform of data-analysis competitions.

- Host proposes a problem based on a dataset → participants propose ML-based solutions.
- Broad and diverse community of participants with different levels of expertise.
- Competitive nature:
 - Competitions tend to last for a few months.
 - Scores are made public in a global leaderboard.
 - Participants are encouraged to continuously improve their solution.

Upon completion of the challenge, winners are required to deliver their winning solutions in such a way that they can be fully reproduced.

Previous GW data-analysis competitions



Monetary prize
for the winners!

Sponsored by Google.

People involved:
Michael Williams
Chris Messenger
Elena Cuoco
Chris Zerafa

Use ML to detect GW signals from binary black hole collisions.

- Over 1200 teams took part during the 3 months of competition.
- Diverse pool of expertise amongst participants, both at academic and industry levels.

Competition overview:

- Finite set of samples:
 - 500k training samples.
 - 220k testing samples.
- Solutions were **ranking** using the **Area Under the ROC** curve.
- No signal generation instructions were shared by the hosts.

Towards a Kaggle competition to detect CWs

We are about to release a Kaggle competition aimed at detecting CWs.

- Profit from the momentum generated by the previous competition.
- CW are longer and weaker than CBC signals: More difficult challenge!
- CBC searches are close to optimal sensitivity; CW are limited by computing cost!
 - Plenty of room for improvement, considering the lack of ready-to-use techniques.

Lessons learnt from the previous challenge:

- Competitors benefit enormously from having as much data as possible.
 - Use PyFstat → New tutorials available!
- Knowledge is shared amongst competitors.
 - Extensive explanations are not required in order to have a successful competition.

Summary

- Continuous gravitational waves are currently undetected forms of gravitational radiation.
- Typical searches for this kind of signals tend to be computationally limited.
- Little effort has been put into developing an end-to-end machine-learning search.
- We will use Kaggle as a mean to spark the development of ML-based CW searches, benefiting from the wide range of expertise amongst the Kaggle community.

Acknowledgments

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