# Interferometer and Data Preparation ET EIB workshop

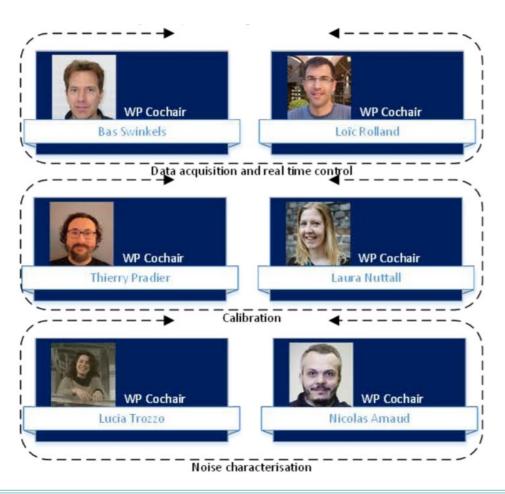
### October 2023

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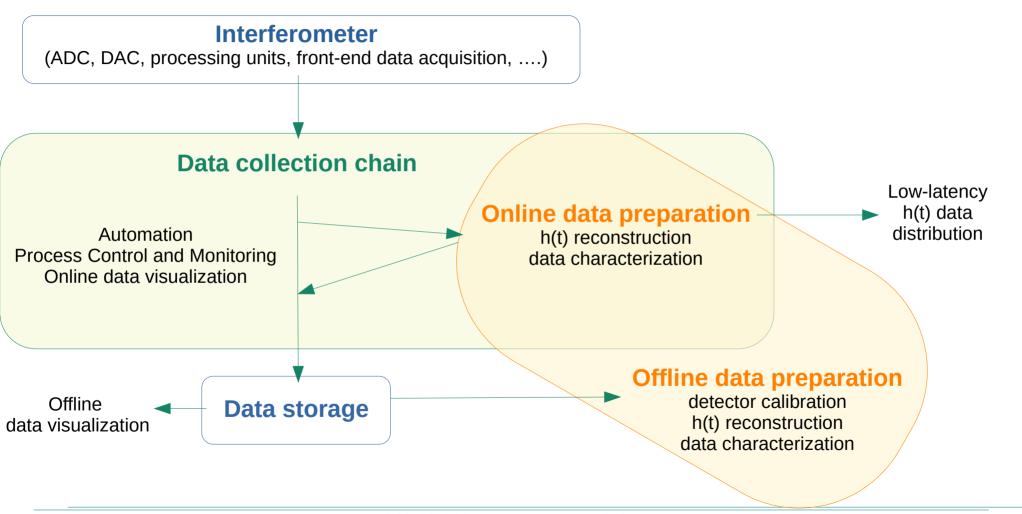
## ISB / Interferometer and Data Preparation WPs



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## Data collection and Data Preparation in Virgo



## LIGO-Virgo-KAGRA data format: gwf (frame format)

### Specification of a Common Data Frame Format for Interferometric Gravitational Wave Detectors (IGWD)

https://dcc.ligo.org/T970130/public

Collaborative effort

Standard definition of the frame format to be used by projects wishing to adhere to a common representation of data produced by GW detectors

Useful for collaborative analyses of data taken by the different projects

Main type of data: time series data of arbitrary duration: v(t)

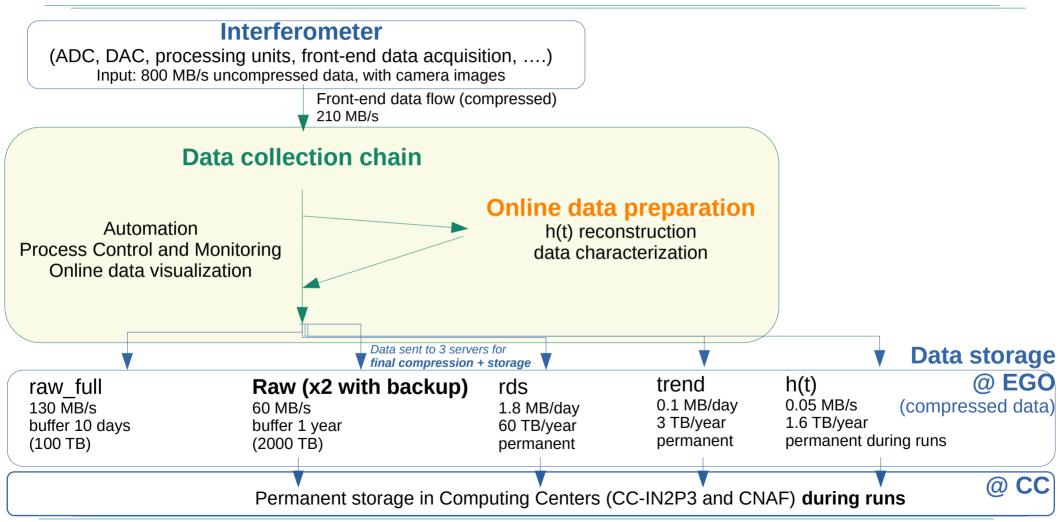
Can encapsulate also other types of data: spectra v(f,t), lists, vectors, arrays v(x,y,t), events, ...

→ write (raw) data into frame structures

Frame Format: capture the informational content of real-time data acquisition systems associated with GW detectors to efficiently archie this information

- → definition of frame header and trailer, data types, composition of files and frames written to media
- → data compression schemes
- → conventions for detector names, channel names, file names

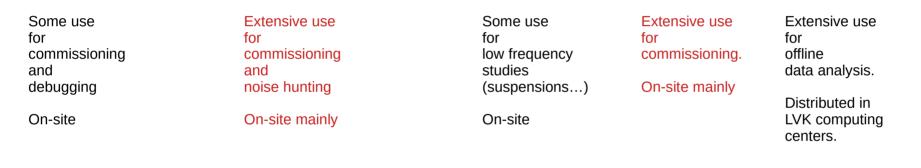
## Virgo data collection: main data flows



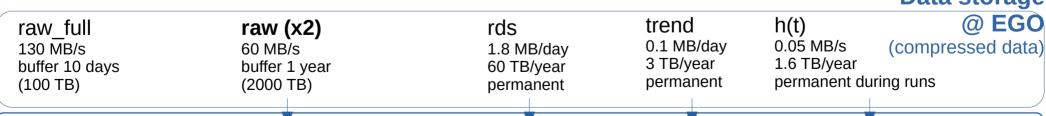
## Virgo data collection: offline usage of main data flows

### Main offline usage for commissioning and noise hunting:

- data visualisation (same tool as for online visualisation)
- data analysis with other tools



### Data storage



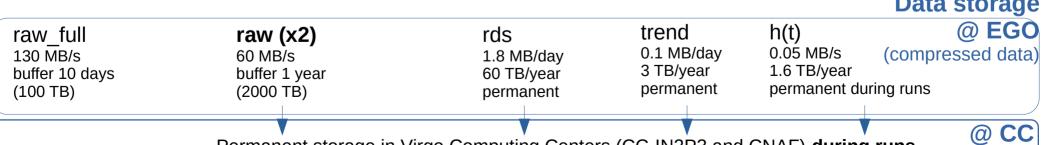
Permanent storage in Virgo Computing Centers (CC-IN2P3 and CNAF) during runs

## Virgo data collection: some remarks on detector data storage

Raw, rds, trend data flows DO NOT scale with the number of GW events depend on the ITF complexity and frequency of the digital control loops

h(t) data flow DOES NOT scale with the number of GW events depend on the sampling frequency of the time series (16384 Hz for current LVK data analysis)

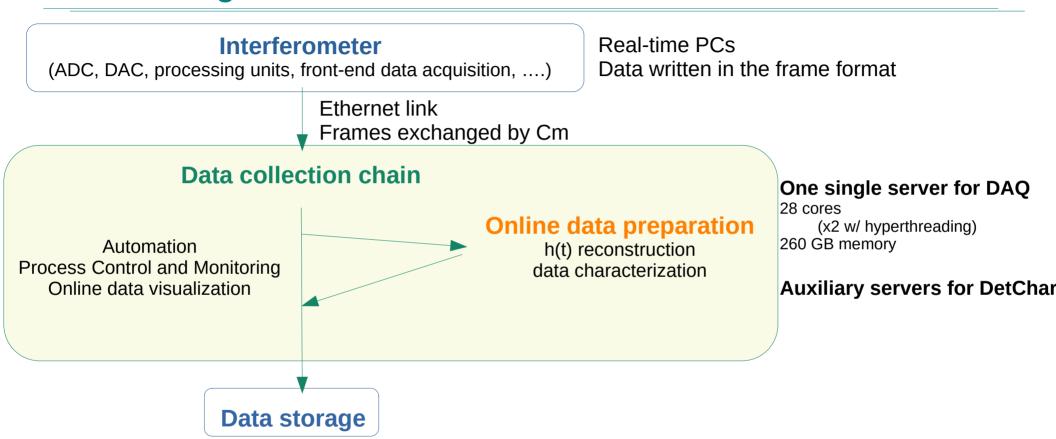
### Data storage



Permanent storage in Virgo Computing Centers (CC-IN2P3 and CNAF) during runs

26/10/2023 – ET-PP EIB workshop -

## Virgo data collection: some more informations



Exchange of messages and data via Cm Exchange of data via Cm or shared memory (gwf files)

## Virgo slow monitoring and automation

#### **Automation**

running on the DAQ server 1 Hz loop

### Detector slow monitoring: some additional machines at EGO

Cm nameserver Virgo Process Monitoring server (VPM) Detector Monitoring System server (DMS)

GUIs Web interfaces for users

### **Camera image visualization in control room screens**

2 vid machines, each with 16 CPU and 50 GB memory

4 screens for display

home-made software for display configuration







## Virgo data collection: some constraints to be kept for E.T.?

### Flexibility of the data collection system (and real-time digital system)

evolution of channels vs time (add/remove channels online, change sampling frequency online, ...) GW detectors are permanently evolving following commissioning and noise hunting

Same data access for both online and offline use

#### Same visualization tools for online and offline data

mainly for commissioning, noise hunting, detector characterization

#### **Data access**

for commissioning for data analysis

Same data acces from on-site and from remote

Same data access for on-site data and for data stored in computing centers

Tools for operation/commissioning/noise hunting from both on-site and remote

## Virgo data preparation: calibration and reconstruction

### Online data preparation

h(t) reconstruction data characterization

### Offline data preparation

detector calibration h(t) reconstruction data characterization

### Calibration

offline data analysis: ~500 MB/year (ROOT and png files)
ROOT and python scripts

### Offline h(t) reconstruction

reprocessed only if online h(t) not good enough distributed to computing centers final h(t) with estimation of h(t) uncertainties vs time: <2 TB/year

#### h(t) data flow DOES NOT scale with the number of GW events

depend on the sampling frequency of the h(t) time series (16384 Hz for current LVK data analysis)

## Virgo data preparation: detector characterization

### Online data preparation

h(t) reconstruction data characterization

### Offline data preparation

detector calibration h(t) reconstruction data characterization some online info stored in the raw data stream

+ low-latency/offline information stored in database, files, ... ??

#### **Main consumers and producers**

- \* Omicron
  - running continuously on one server: 16 cores, 50 GB memory
  - ~ 3 TB/year for Omicron online output (run on-site) if needed: reprocessing of Omicron, run at CC-IN2P3
- \* SpectroMoni running continuously on one server ~2 TB/vear
- \* VIM
  regular update of plots on web pages
  ~ 1 TB/vear
- \* VDQR (Virgo Data Quality Report)

  1 GB/alert
- 998.6 Online area
  338.7 Total
  419.0
  259.1
  419.0
  259.1
  119.7
  299.3
  0.0
  2023/09/01 2023/06/01 2023/07/01 2023/09/01 2023/10/01
- $\rightarrow$  scale with detection rate
- \* Others: Bruco (~300 GB/year), NoeMI

## **Detector commissioning**

Virtual machines at EGO used for online and offline data access for commissioning and noise hunting

presently: 38 ctrl machines, each with 8 CPU and 16 GB memory

connexion from Virgo control room and on-site laboratories connexion from external labs for remote operation and studies





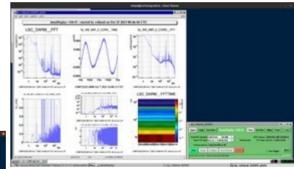
### Some examples of uses

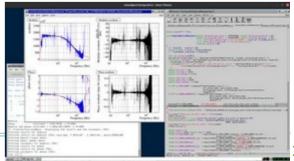
data visualization, online and offline GUIs (and command-lines) for detector control and monitoring edition of the configurations of the control processes for the detector sub-systems

data analysis with user's scripts: python, ROOT, Matlab, ... testing noise budget

. . . .







### **Detector simulations**

### For design and then detector commissioning

→ understand the interferometers and understand noise sources

#### **Current situation**

Currently: 31 different codes (for optics, electronics, mechanics, suspensions, controls, noise budget)

https://indico.ego-gw.it/event/562/contributions/4999/attachments/2748/4846/20230507\_ETSymposium\_WPModelSim.pdf (sildes from XIII ET Symposium)

https://wiki.et-gw.eu/ISB/Interferometer/ModelDesignTools/Swlist

Current simulations are mostly running on laptops, no particular need in term of CPU power

#### Some needs (already for Virgo)

Need for personpower for code development, test and maintenance

Need for infrastructure for testing the different codes, store the configurations and associated outputs....

#### What about "Digital twins"?

→ stronger needs (long-term personpower, combining various fonctionalities, ...)

## Preliminary estimations and questions for ET

### \* Very preliminary assumptions for E.T. data flows

- assumption: 6 "independent" ITFs
  - + new sub-systems to monitor and control (cryogeny, more suspended mirrors/benches, newtonian noise cancellation, ....)
  - + frequencies of the numerous digital loops for E.T. control? (10 kHz in Virgo for longitudinal controls)
  - + need for faster control loop (a few loops running at 400 kHz in Virgo)
  - + sampling frequencies of the ADC channels?
  - + need for channels sampled at ~1 MHz/400 MHz (possible in Virgo since a few years)
- environmental sensors partly shared but probably more monitoring than in Virgo
- → data flows multiplied by ~10 to 20 compared to Virgo?

### \* Storage architecture?

- independent streams stored for every detectors + tools to access multiple streams
- or store merge streams for every detectors

### \* Data format

- same format for online data collection, data storage and offline usage as in Virgo?
   or different formats?
- which format(s)? gwf, hdf5, other formats, ...

### \* Running detector simulations for commissioning and/or monitoring the detector controls?

## **CPU**

## Ganglia info for olserver52 (Virgo DAQ server)



CPU Count CPU Speed Memory Total

Swap Space Total

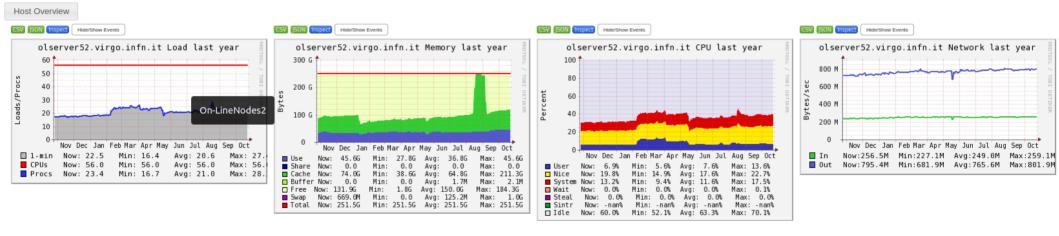
56 CPUs

3300 MHz

263742464 KB

4194300 KB

#### EGO Computing Grid > On-LineNodes2 > olserver52.virgo.infn.it



## Ganglia info for olserver52 (Virgo DAQ server)



## Ganglia info for olserver53 ("spare "Virgo DAQ server)

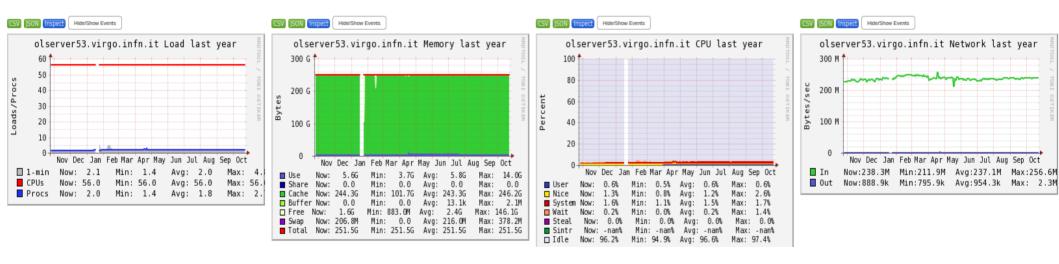


 CPU Count
 56 CPUs

 CPU Speed
 3300 MHz

 Memory Total
 263742464 KB

 Swap Space Total
 4194300 KB



Used for running DetChar/Spectro process

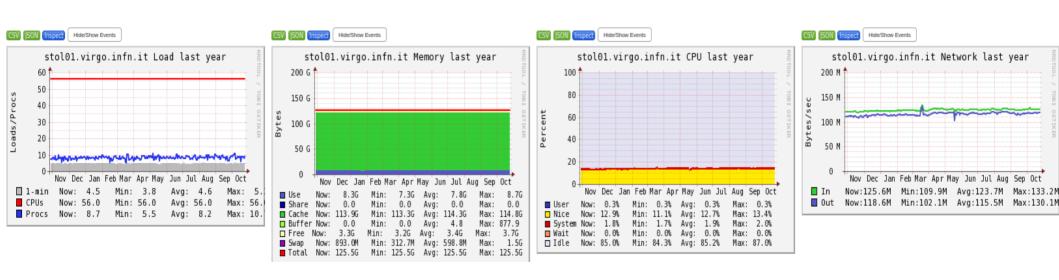
## Ganglia info for olserver53 ("spare" Virgo DAQ server)



Used for running DetChar/Spectro process

## Ganglia info for stol01 server (raw data compression and storage)



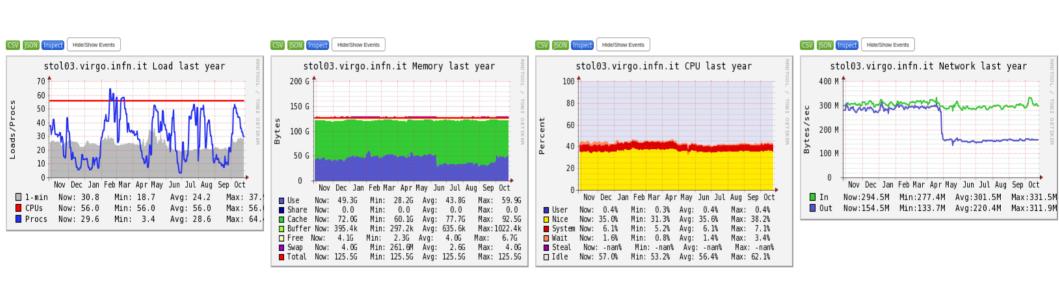


## Ganglia info for stol01 server (raw data compression and storage)

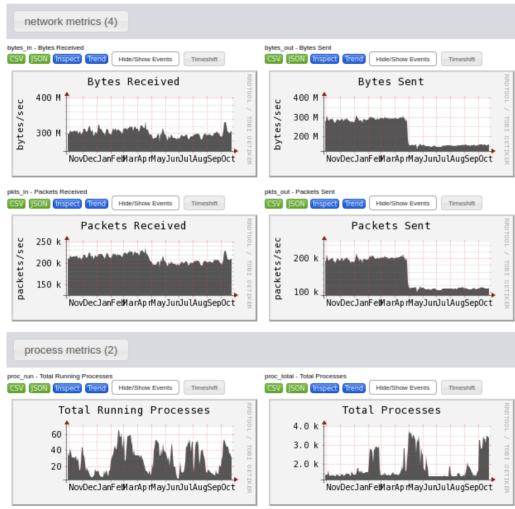


## Ganglia info for stol03 server (other data compression and storage)





## Ganglia info for stol03 server (other data compression and storage)

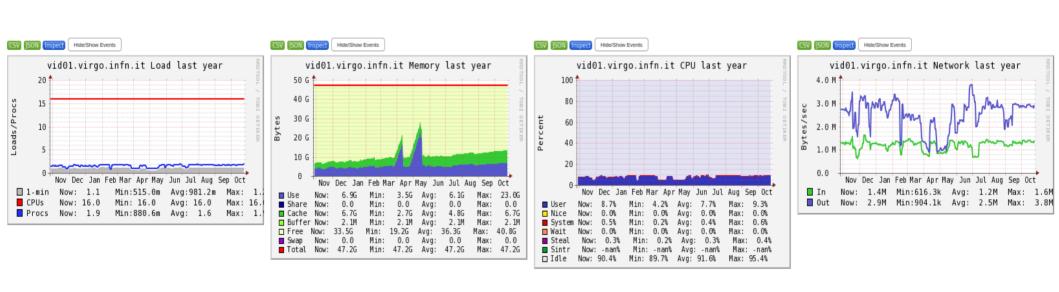


## Ganglia info for vid01 server (camera visualisation in control room)

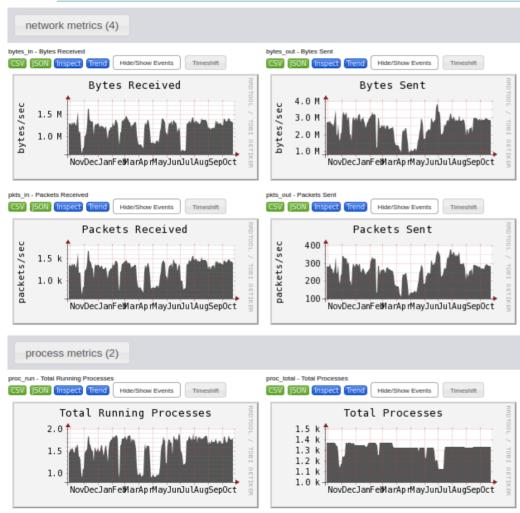


CPU Count CPU Speed Memory Total Swap Space Total 16 CPUs 2394 MHz 49444328 KB

3145724 KB



## Ganglia info for vid01 server (camera visualisation in control room)



## Ganglia info for olserver116 (Virgo Omicron server for DetChar)

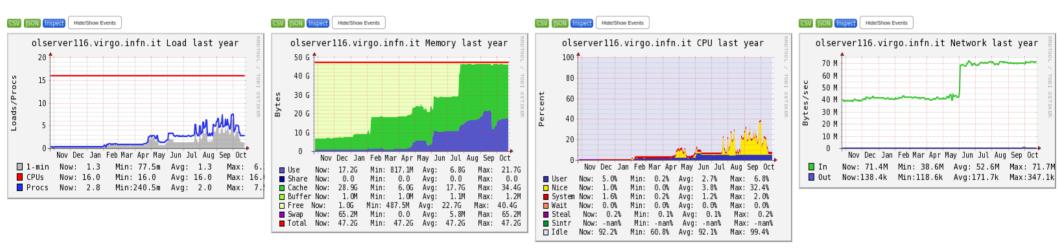


 CPU Count
 16 CPUs

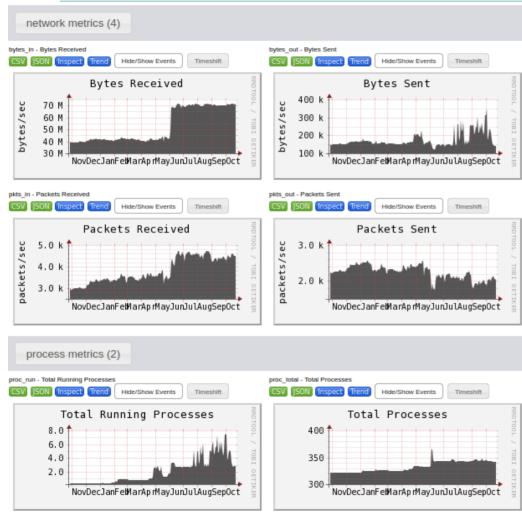
 CPU Speed
 2394 MHz

 Memory Total
 49444328 KB

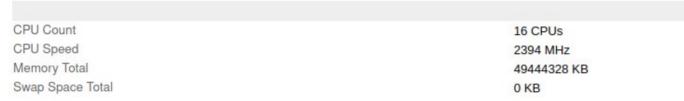
 Swap Space Total
 3145724 KB



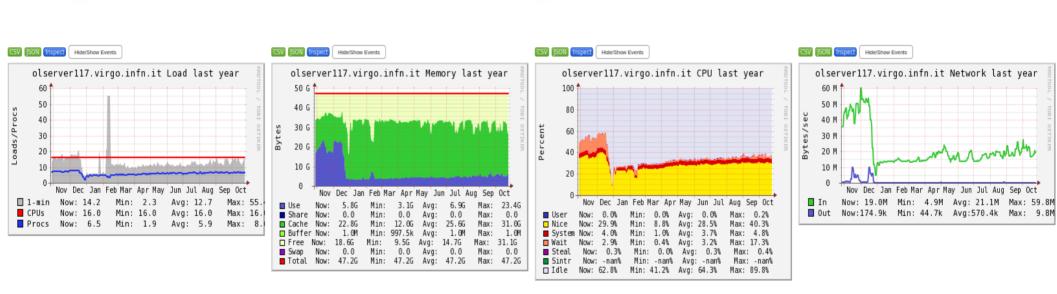
## Ganglia info for olserver116 (Virgo Omicron server for DetChar)



## Ganglia info for olserver117 (VIM, Bruco for DetChar)



Constant Metrics



## Ganglia info for olserver117 (VIM, Bruco for DetChar)



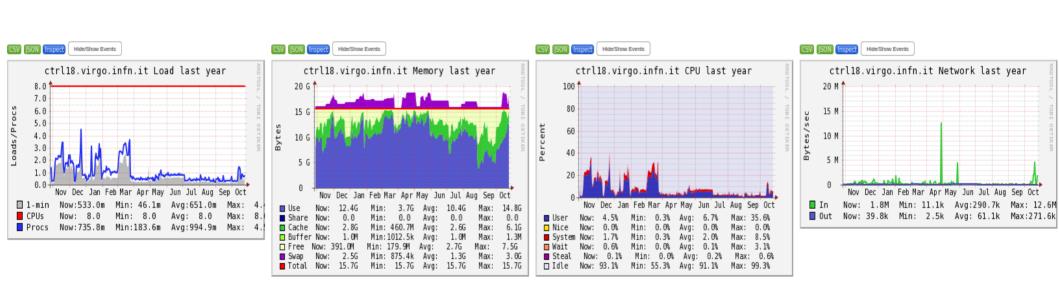
## Ganglia info for ctrl18 server (control workstation example)



CPU Count
CPU Speed
Memory Total
Swap Space Total

2600 MHz 16414200 KB 3145724 KB

8 CPUs



## Ganglia info for ctrl18 server (control workstation example)

