

# Silicon mirrors – an asymmetric test mass configuration

Iain Martin<sup>1</sup> and Jessica Steinlechner<sup>2,3</sup>

<sup>1</sup>University of Glasgow

<sup>2</sup>Maastricht University

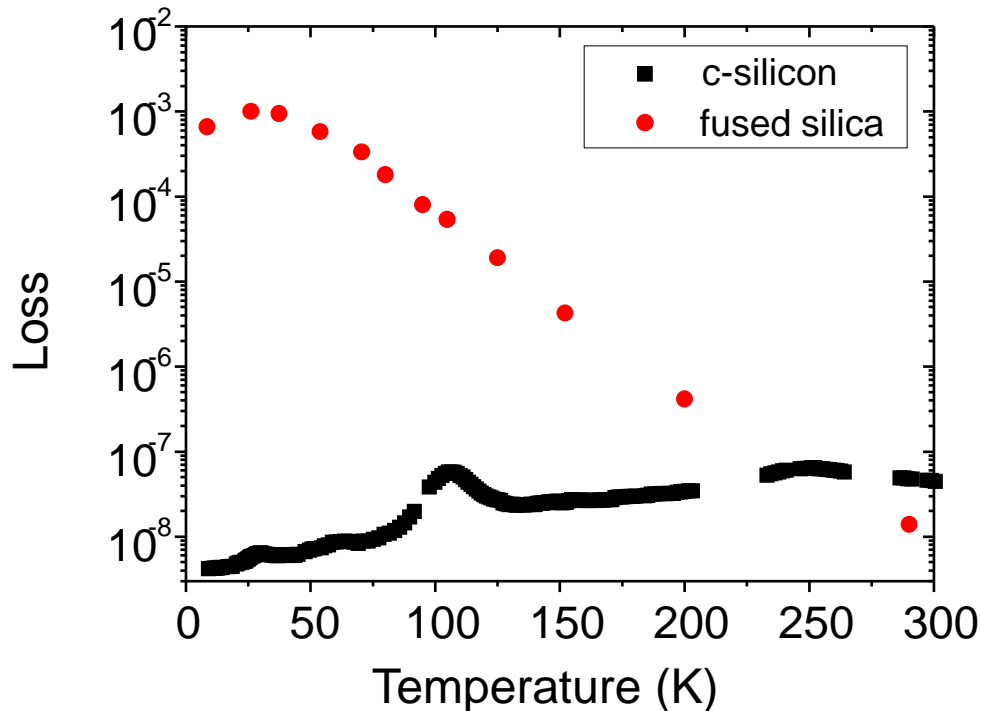
<sup>3</sup>Nikhef

# Overview

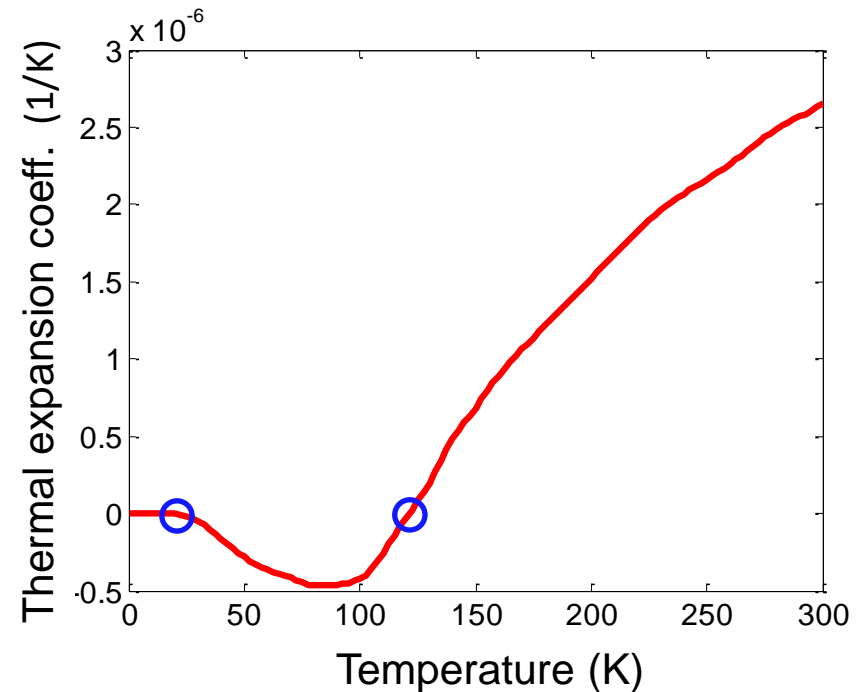
- Silicon as a test mass material
- Silicon – types, sizes and optical absorption
- Absorption tolerances of ITM and ETM
- Asymmetric test-mass configuration concept
- Implications for coating thermal noise

# Silicon mirrors for GW detectors

- Silicon attractive due to low mechanical loss (low thermal noise), zeros in thermal expansion coefficient (low thermoelastic noise)



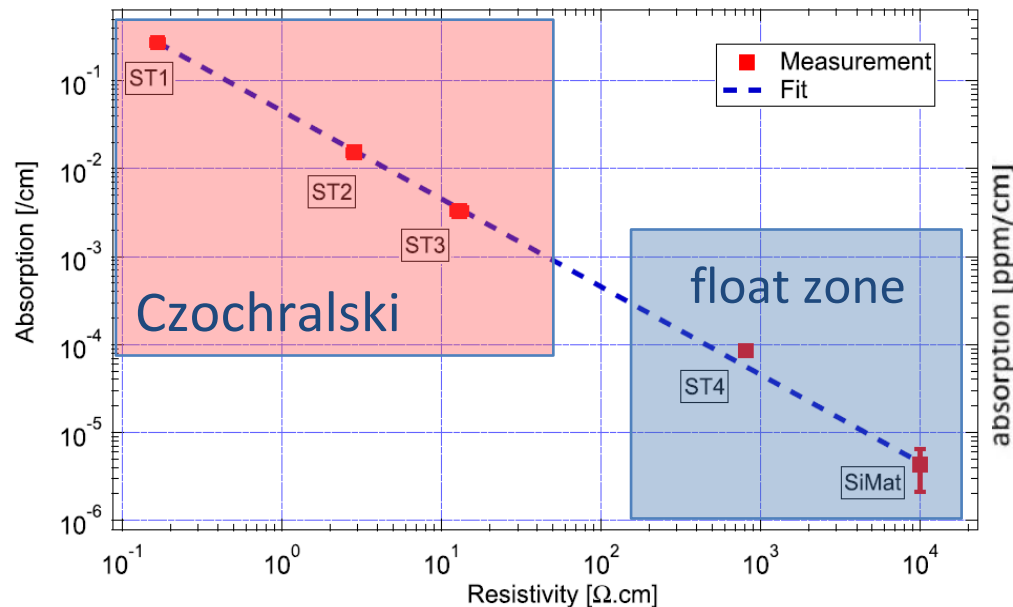
[This data – Schwarz et al., 2009]



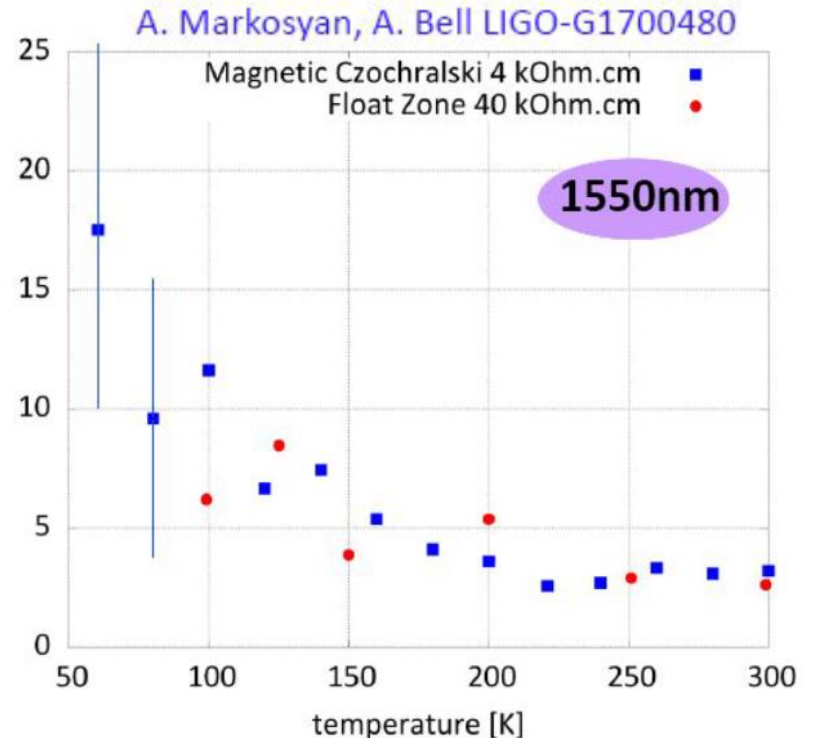
[Touloukian 1970]

# Silicon mirrors for GW detectors

- **Float zone** silicon with low levels of impurities can have low optical absorption – size (currently?) limited to  $\approx 20$  cm diameter
- **Czochralski** silicon available in larger sizes, but tends to have higher absorption
- **Magnetically purified Czochralski** growth – can produce similar absorption to good float zone (but indications of large spatial variation), in principle can be made in sizes up to  $\approx 40$ -45cm



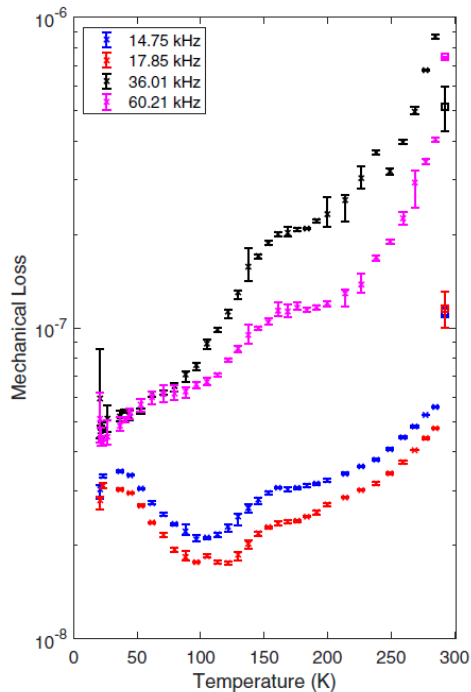
Degallaix et al., Optics Letters 38 (2013)



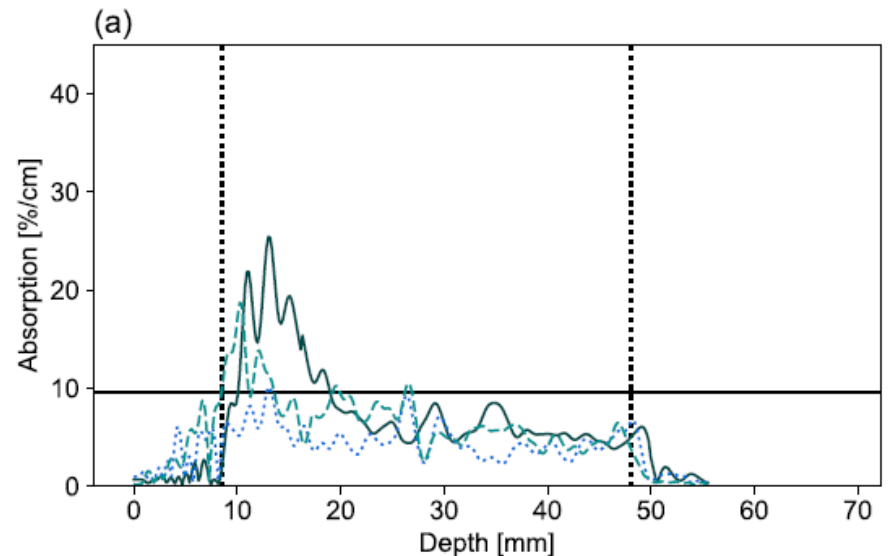
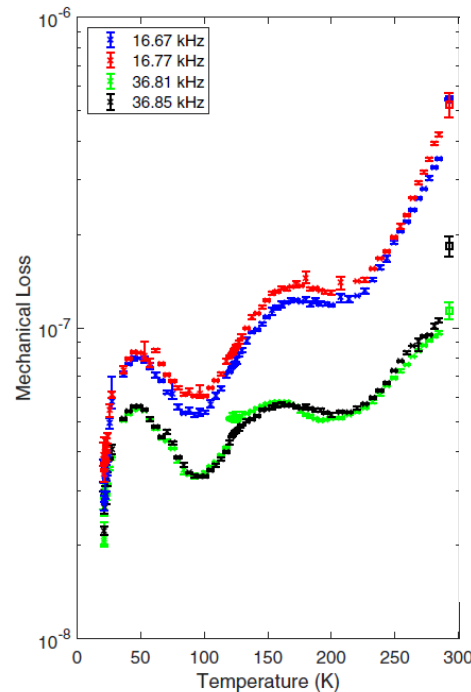
# Silicon mirrors for GW detectors

- **Quasi-monocrystalline** silicon grown by directional solidification
  - Can be produced in large sizes e.g. squares > 1 m
  - Cryogenic mechanical loss comparable to float zone
  - Test sample had high absorption, however doped feedstock used
  - Investigations into improving absorption required e.g. pure feedstock, alterations to growth process to reduce impurities
  - Growth likely to need optimised to ensure a large enough area is monocrystalline

(a) Quasi-monocrystalline silicon



(b) Float-zone silicon



[Bruckner et al. PRR 4 (2022)]

# Silicon mirrors for GW detectors

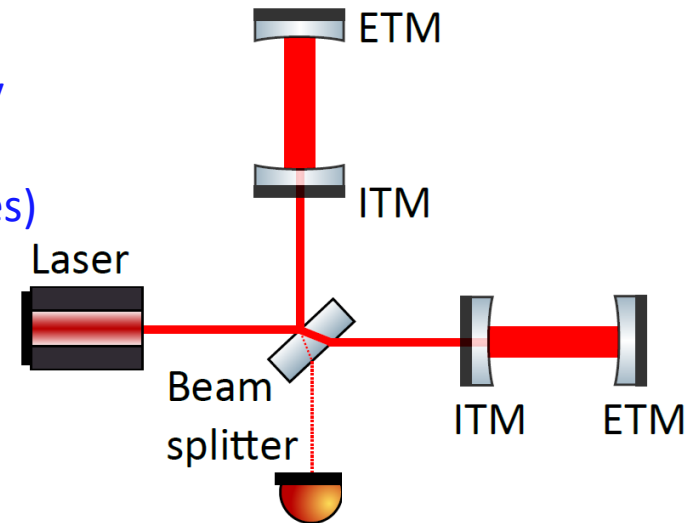
- Broadly speaking: we can have large silicon, or we can have silicon with low optical absorption, but challenging (so far) to get both at once
- Option 1 – find a technique for making low optical absorption silicon in larger sizes that is economically viable for GW detectors (magnetically purified Czochralski shows some promise in principle)
- Option 2 – construct composite test masses by bonding together pieces of low optical absorption silicon
  - Potential for excess thermal noise from bonds close to the laser beam
  - Potential for optical effects (scattering / distortion of wave fronts / absorption) due to transmitting beam through bonds
- Option 3– can we exploit different requirements of ITMs and ETMs and use an ‘asymmetric test-mass’ design?

# Requirements for ITMs and ETMs

- Different absorption tolerances on the ITM and ETM, as the ITM transmits more laser power - e.g. for ET-LF

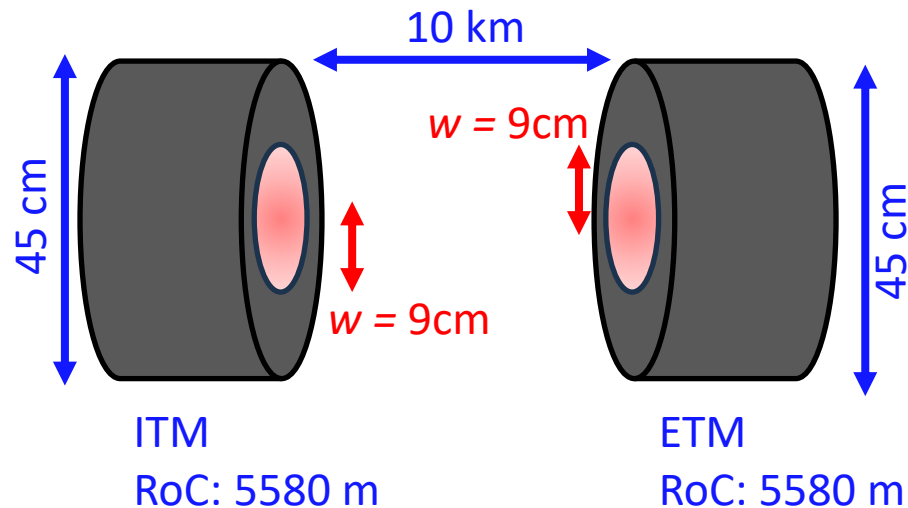
- ITM transmits  $\approx 30$  W
  - ETM transmits  $\approx 90$  mW
- (Assuming 18 kW circulating arm cavity power and standard ITM/ETM reflectivities)

- So the ETM material can have a factor of  $30$  W/ $90$  mW  $\approx 300$  higher optical absorption than ITM material, and absorb the same absolute laser power
- Only need low absorption material for the ITM



# Asymmetric test-mass configuration

- Only need low absorption for ITM – can we use a small float zone mirror for ITM and only use a large mirror for the ETM?

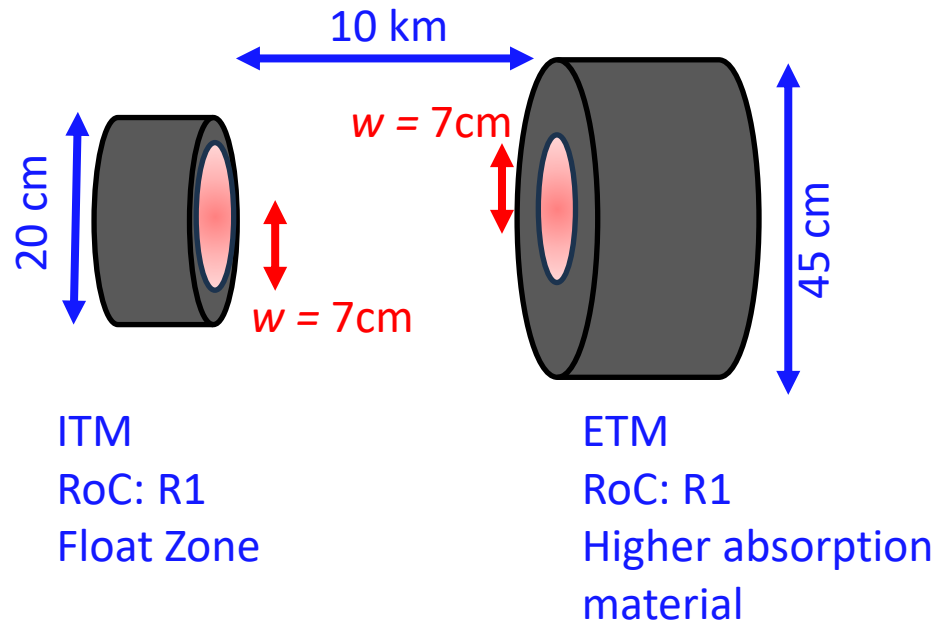


ET arm configuration from Design Study



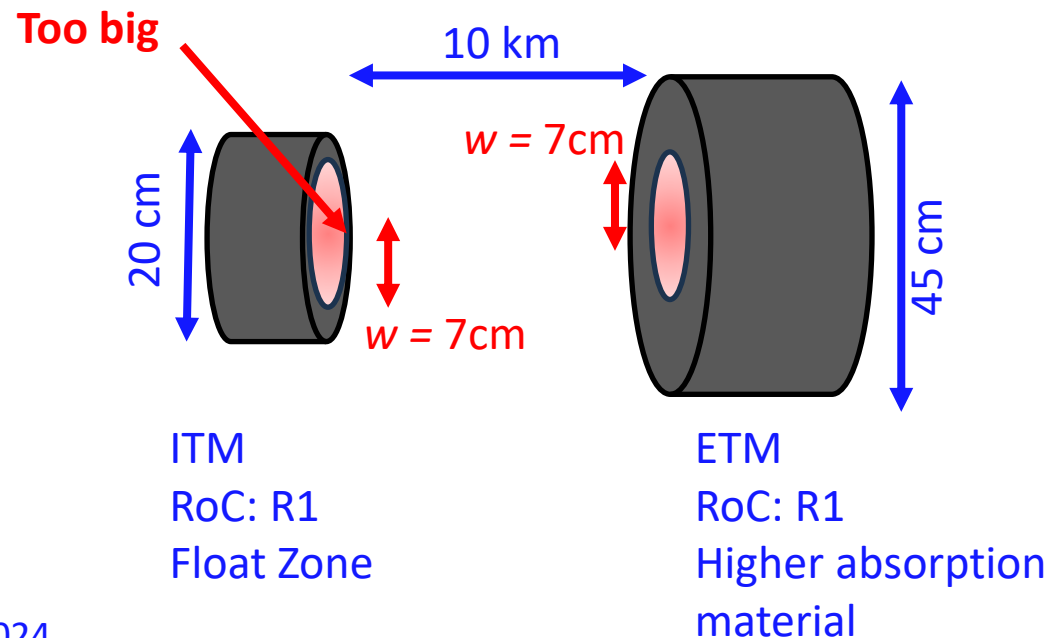
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- Beam size is limited by mirror size (scattering losses at edge) – want mirror diameter  $\approx 2.5 \times$  beam diameter i.e. beam radius 4 cm for 20 cm mirror



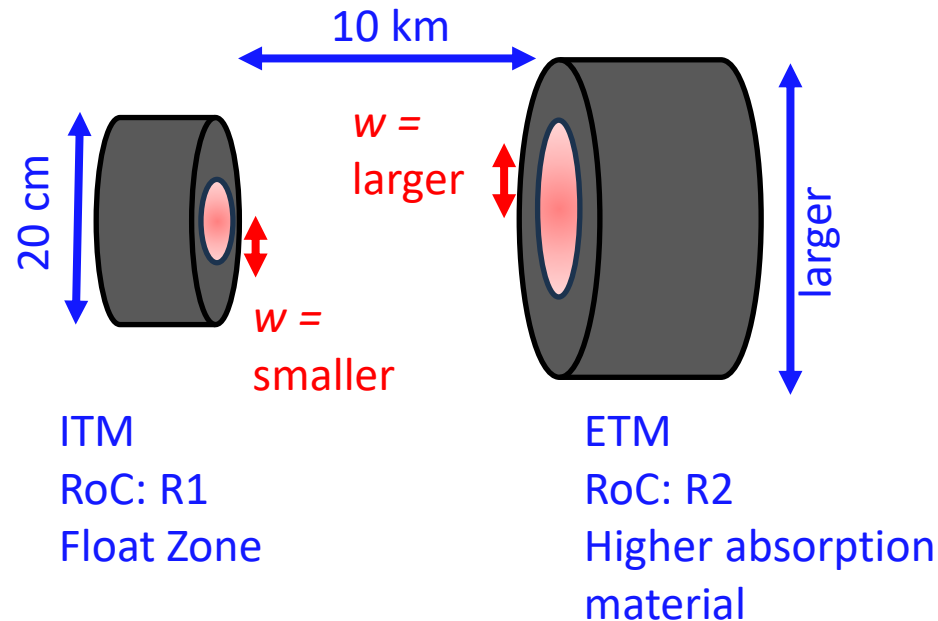
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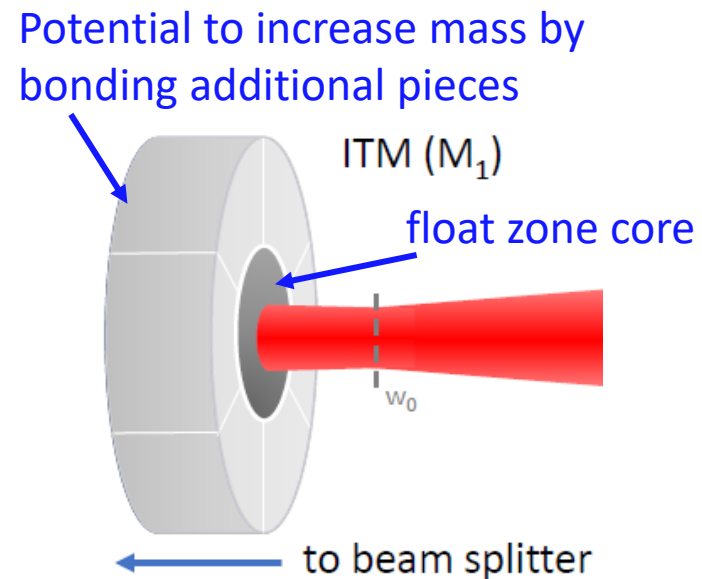
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- But we can make beam smaller on the ITM – resulting in larger beam on ETM



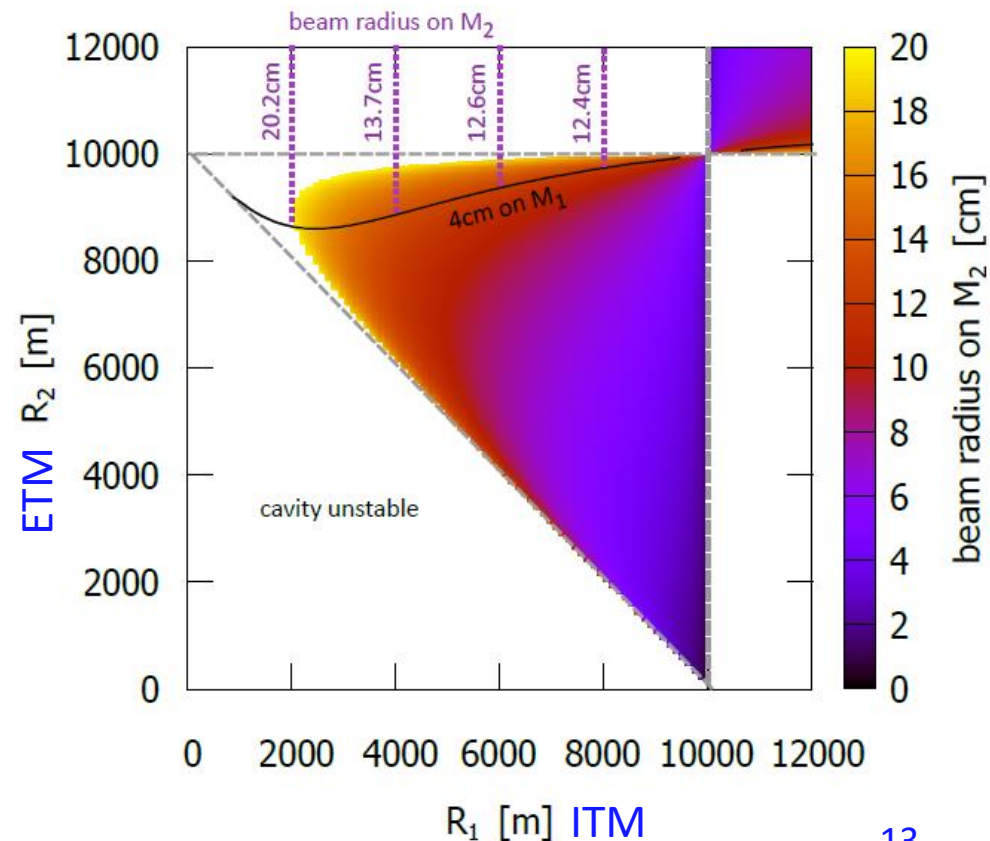
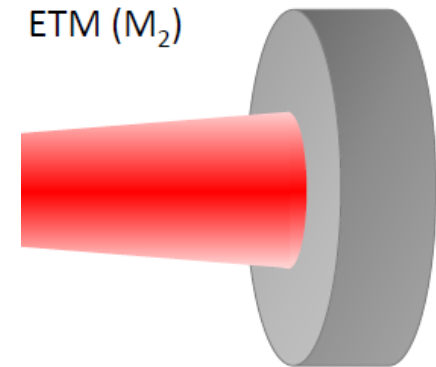
# Input test masses

- ITM - use low absorbing **float zone silicon**, available (currently?) in diameters of  $\approx 20$  cm, rather than the  $\approx 45$  cm Design Study value
- Requires **smaller beam** to avoid excess scattering at edges of mirror
  - e.g. beam radius of  $\approx 4$  cm to maintain factor of 2.5 between beam diameter and mirror diameters
- Reduction in ITM mass not ideal (e.g. for radiation pressure noise)
  - Possibly solve by **bonding** pieces around 20 cm float zone core
  - Reducing potential issues with the bonds transmitting the laser beam in the “classical” composite mass case.
- Smaller beam will increase coating thermal noise – but more later....



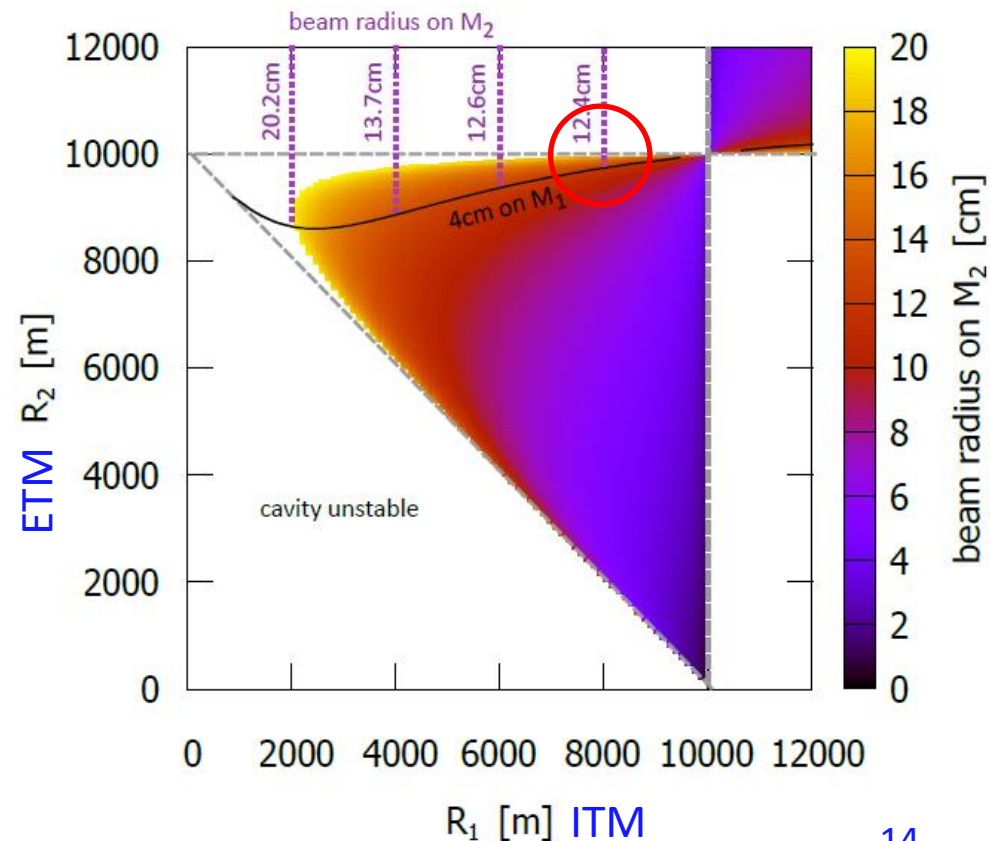
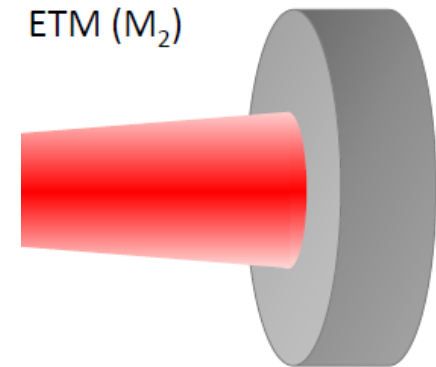
# End test masses

- ETM – can use **larger diameter silicon** due to relaxed absorption requirement e.g. Czochralski / quasi-monocrystalline silicon
- Larger mirror allows larger laser beam – reducing the ETM coating thermal noise
- For a stable cavity with a 4 cm beam radius on the ITM (allowing use of float zone Si):
  - the minimum beam radius on the ETM is  $\approx 12.5$  cm
  - resulting in an ETM mirror diameter of  $\approx 62.5$  cm being required



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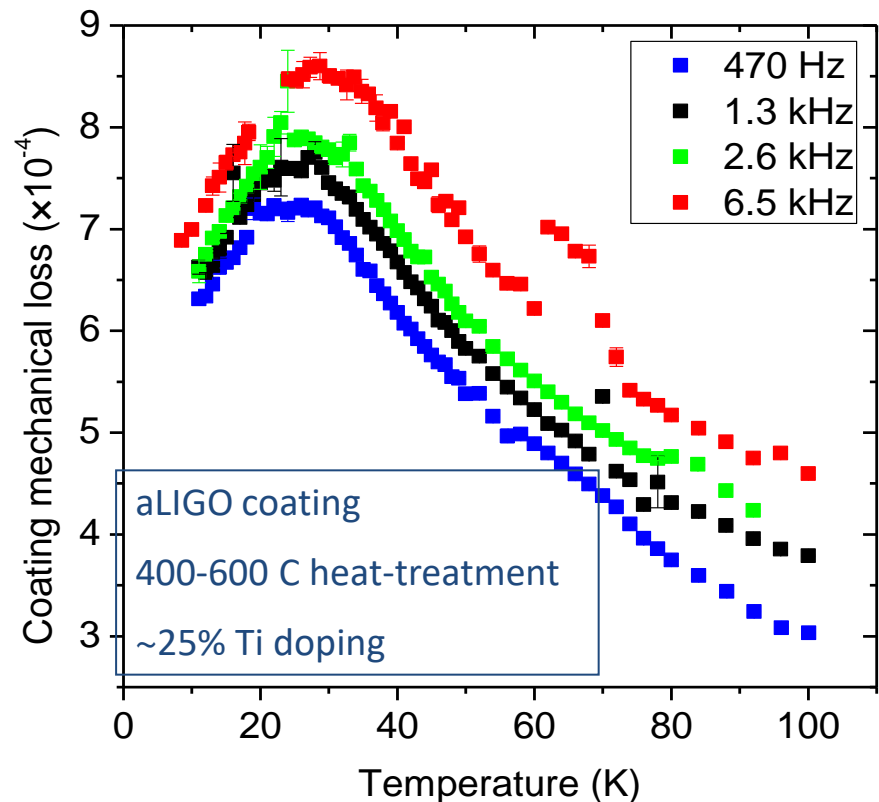
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# Coating thermal noise

- Coating thermal noise (CTN) reduces when cooling to cryogenic temperature
- For current aLIGO/Adv Virgo coating materials ( $\text{SiO}_2/\text{Ti}:\text{Ta}_2\text{O}_5$ ) reduction partially offset by increase in cryogenic mechanical loss
- ET Design Study CTN target
  - $\lesssim 3.6 \times 10^{-21} \text{ m/Hz}^{-1/2}$  at 10 Hz
  - target is 2x lower than likely to be achievable using  $\text{SiO}_2/\text{Ti}:\text{Ta}_2\text{O}_5$  coatings at 10 K (3x at 20 K)
  - assumed 9 cm (radius) beams on ITM and ETM

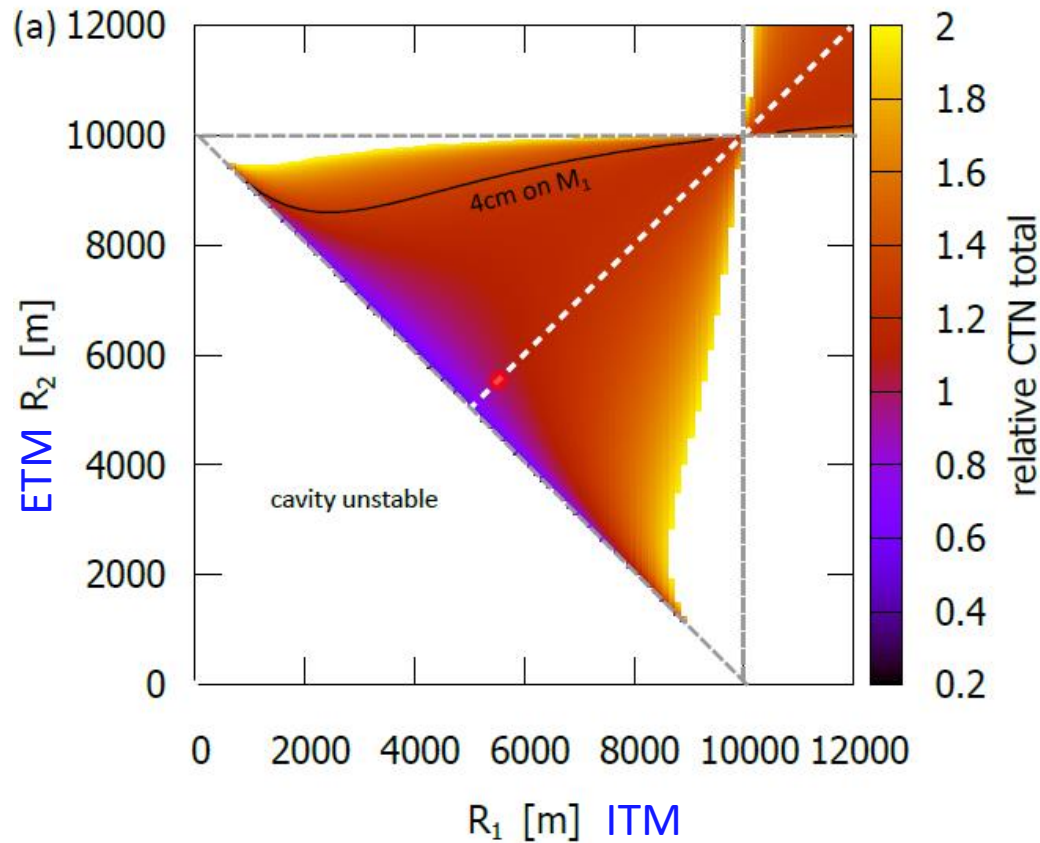
$$x(f) = \sqrt{\frac{2k_B T}{\pi^2 f} \frac{d}{w^2} \phi \left( \frac{Y_{\text{coat}}}{Y_{\text{sub}}^2} + \frac{1}{Y_{\text{coat}}} \right)}.$$



M Granata et al Optics Letters 38 (2013)

# Coating thermal noise optimisation 1

- ITM coatings less reflective  $\rightarrow$  thinner  $\rightarrow$  lower thermal noise
- CTN goes as  $1/w^2$
- Red dot: ET Design Study case – 9 cm beam on both mirrors (5580 m RoC)
- White dashed line shows equal RoC on both mirrors
  - due to thicker ETM coating, slightly larger beam on ETM and smaller on ITM is favourable
  - aLIGO uses slightly asymmetric beams (smaller ITM, larger ETM); Virgo also considered/planning to use much larger beam on ETM



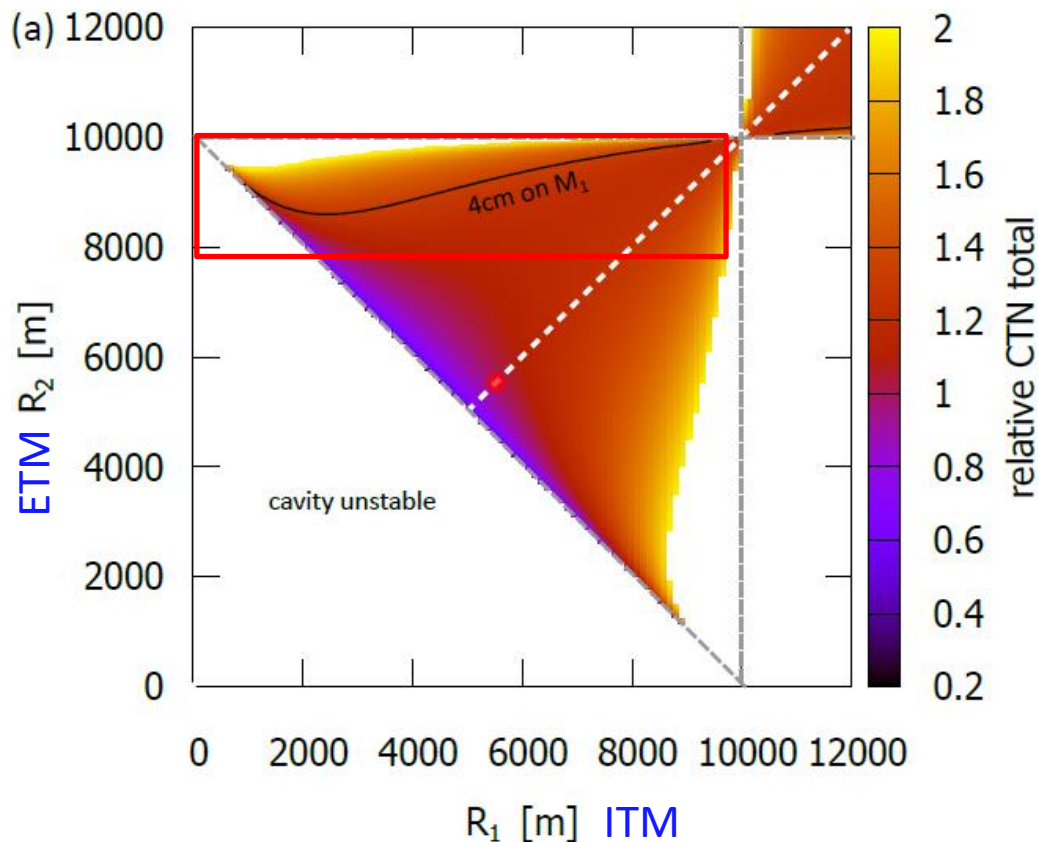
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$$x_{\text{total}} = \sqrt{2 \times x_{\text{ETM}}^2 + 2 \times x_{\text{ITM}}^2}$$



# Coating thermal noise optimisation 2

- Suggested configuration of 4 cm beam on 20 cm float zone ITM, while keeping ETM as small as possible, results in overall increase in CTN by  $\approx 30\%$
- No solution with a 4 cm ITM beam that reduces CTN
- Increase in CTN may be acceptable if this configuration is the only viable way to solve the silicon size problem

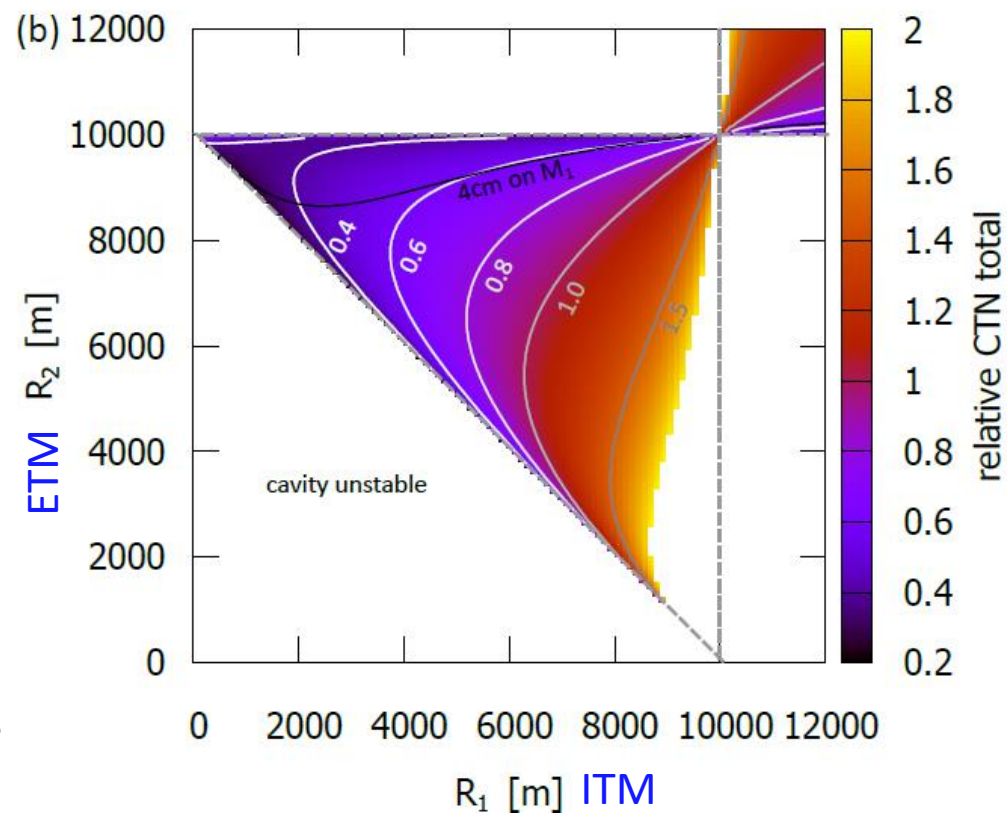


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# Coating thermal noise optimisation 3

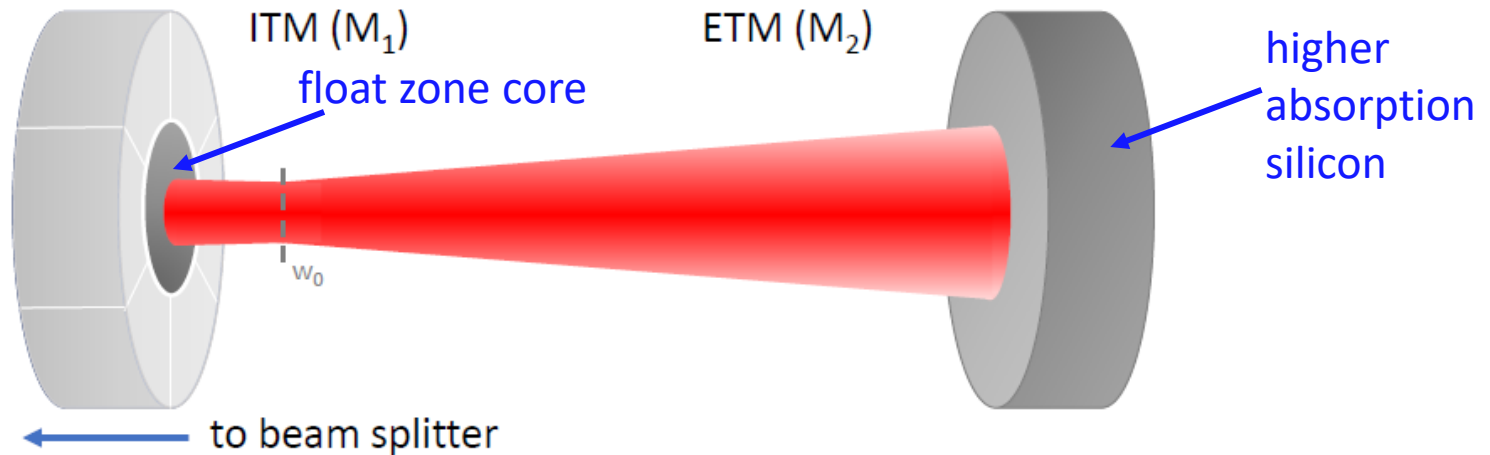
- However, 20 cm ITM allows possibility of using AlGaAs coatings
  - Single crystalline MBE coatings
  - very low cryogenic loss (up to 100x lower than SiO<sub>2</sub>/Ta<sub>2</sub>O<sub>5</sub>)
  - very low absorption
  - currently limited in size to 20 cm diameter due to availability of GaAs wafers as growth substrates
- Use of AlGaAs on ITM and a 4-material amorphous coating on ETM (Craig et al. PRL 122, 2019) gives significant CTN reduction with our proposed configuration (4 cm beam on 20 cm ITM) – lower than the ET Design Study CTN requirement



$$x(f) = \sqrt{\frac{2k_B T}{\pi^2 f} \frac{d}{w^2} \phi \left( \frac{Y_{\text{coat}}}{Y_{\text{sub}}^2} + \frac{1}{Y_{\text{coat}}} \right)}$$

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# Summary



- Investigated using **different types of silicon** and **very different beam sizes** for the ITMs and ETMs
- Based on 20 cm diameter float zone, a stable configuration that minimises the size of the ETM is:
  - 20 cm diameter float zone Si ITM with 4 cm beam radius
  - 62.5 cm diameter ETM, higher absorbing Si, 12.5 cm beam radius
  - Possibility to increase mass of ITM via bonding a composite mass, but with the beam 'entirely' contained on the 20 cm float zone core
- Small ITM beam would **allow the use of 20 cm AlGaAs ITM coatings**, potentially allowing the **ET-LF CTN goal to be surpassed**

# Challenges and open questions

- How close to the cavity stability criterion can we go in practice?
- Bonding to increase mass of ITM – what bonding configurations are possible? What is the thermal noise effect of bonds located outside the nominal laser beam radius?
- AlGaAs coatings - in principle can work up to 20 cm but (to my current knowledge) 20 cm AlGaAs substrate transfer to silicon has not yet been demonstrated
- Can quasi-monocrystalline silicon absorption be lowered enough for use as an ETM? If not, can we get big enough ETM, with low enough absorption using some other method?