

Euclid and ET Synergies

Andrea Cimatti

University of Bologna
Department of Physics and Astronomy *Augusto Righi*

On behalf of the Euclid Consortium

XV Einstein Telescope Symposium - 28/05/2025

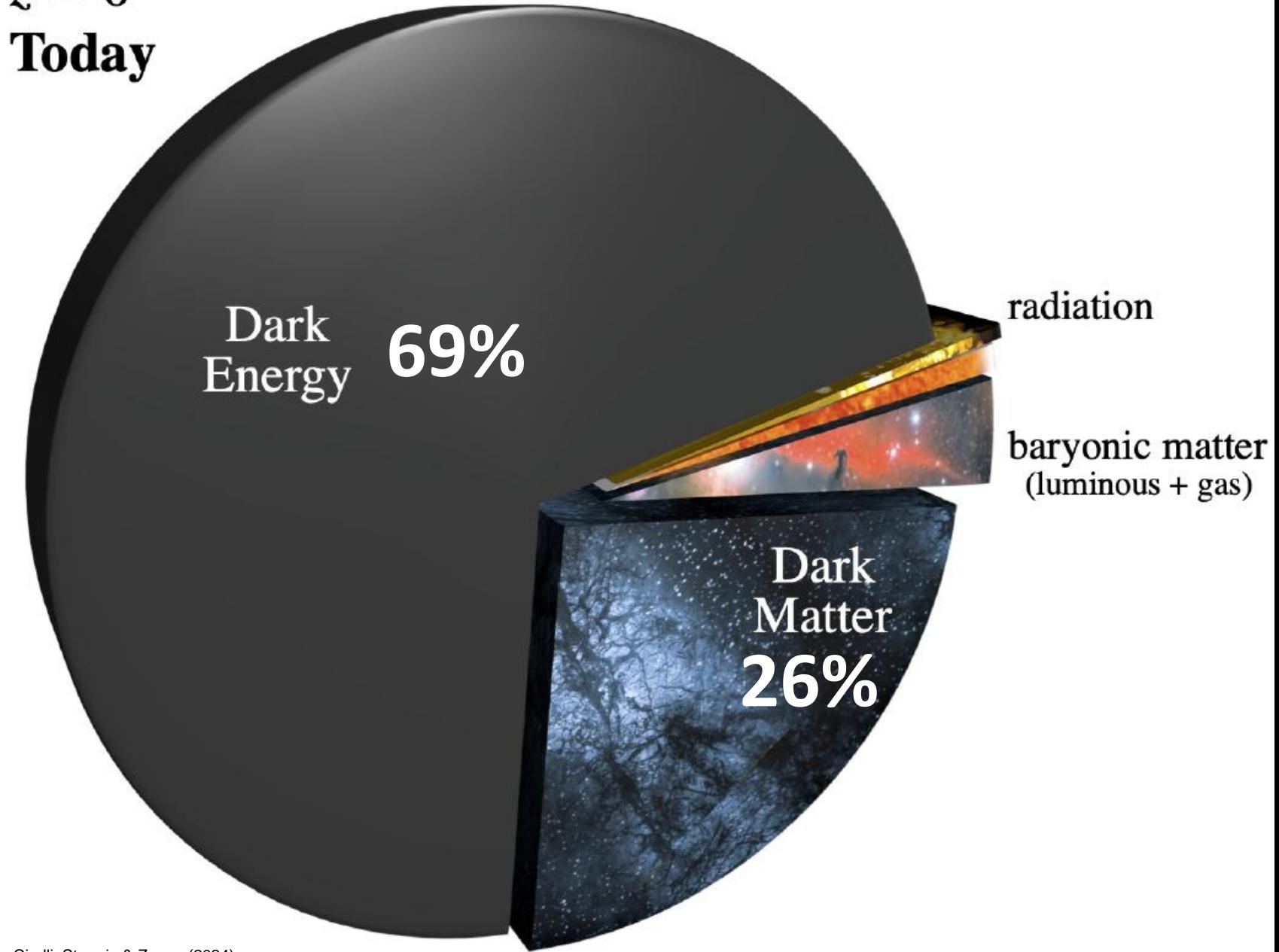


Part I

The ESA Euclid Mission

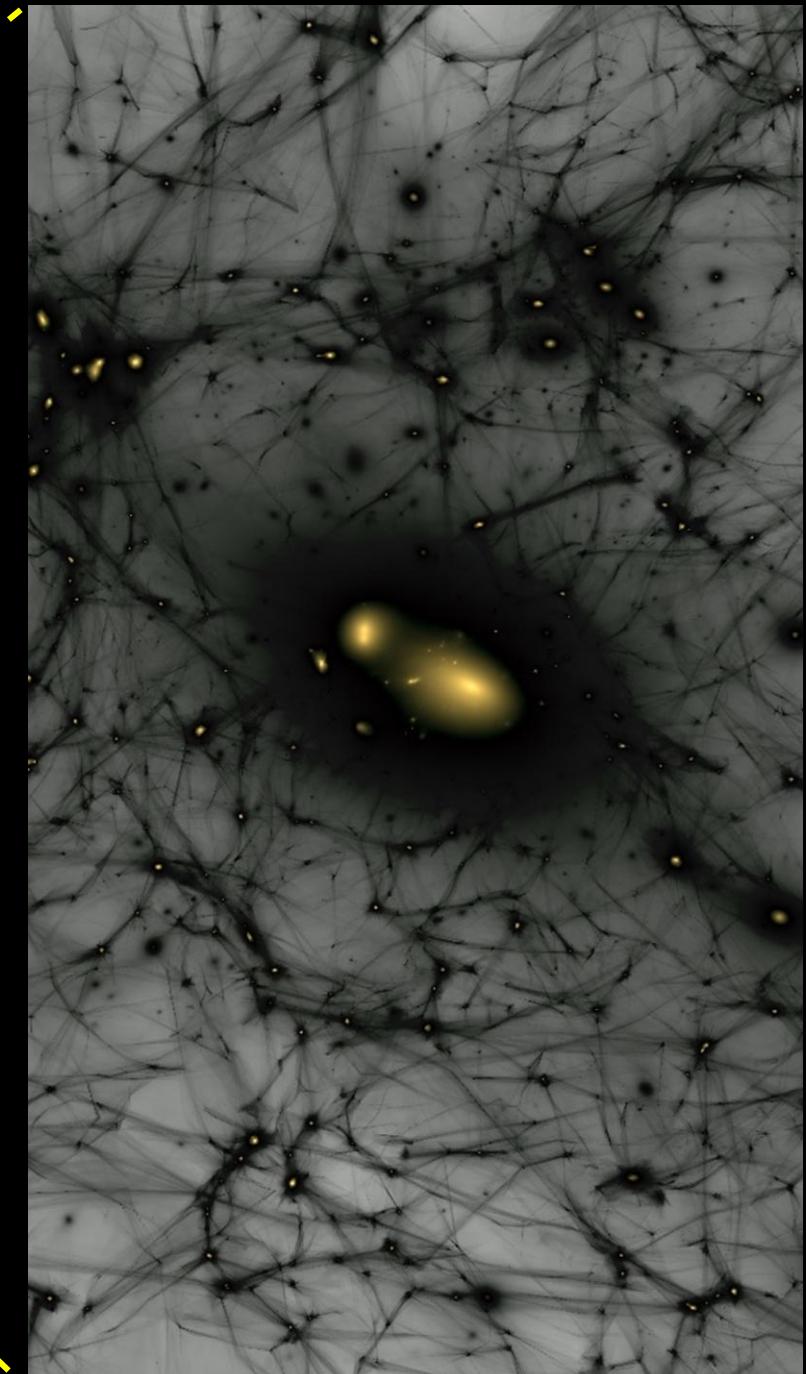
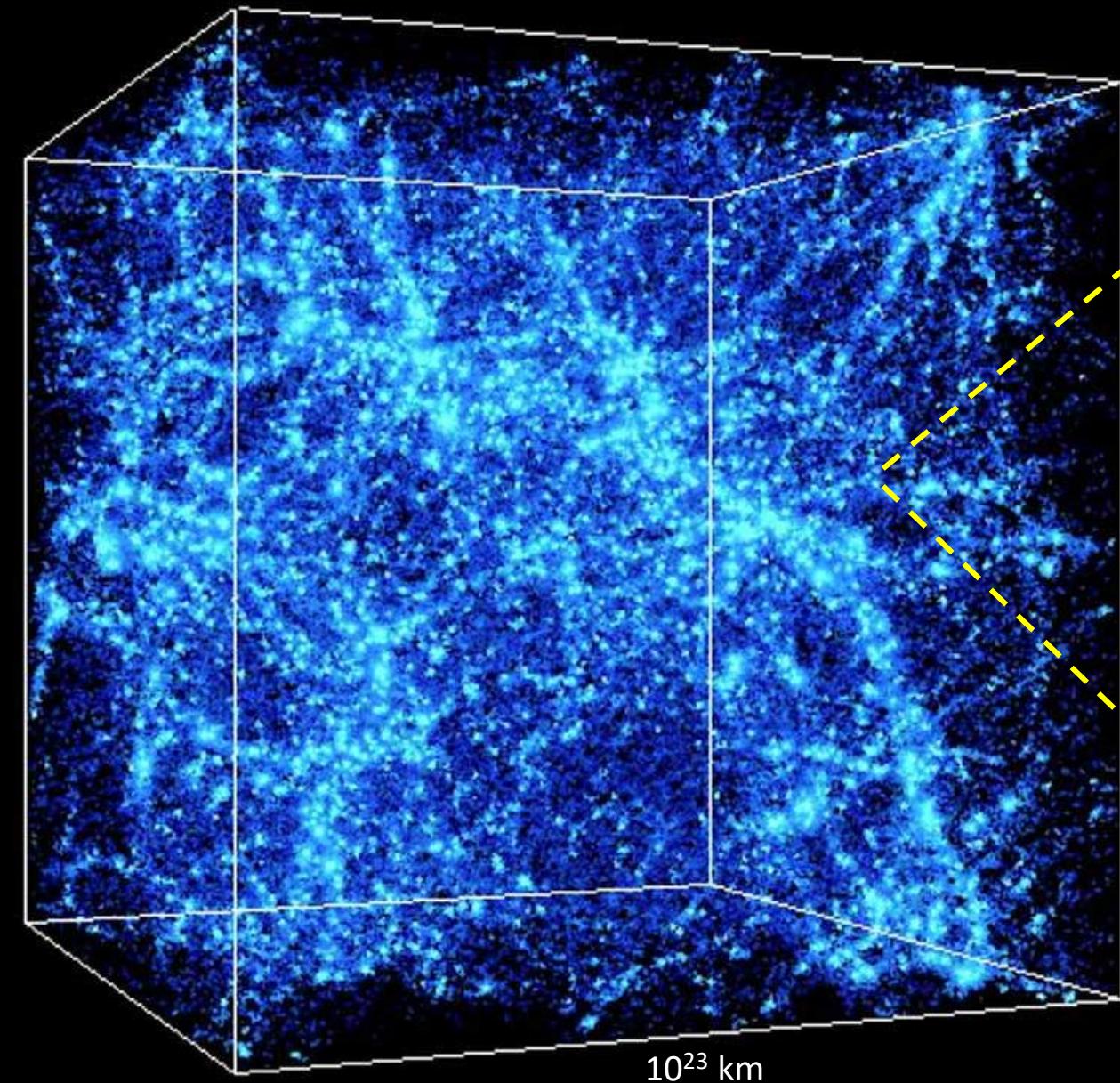
$z = 0$

Today

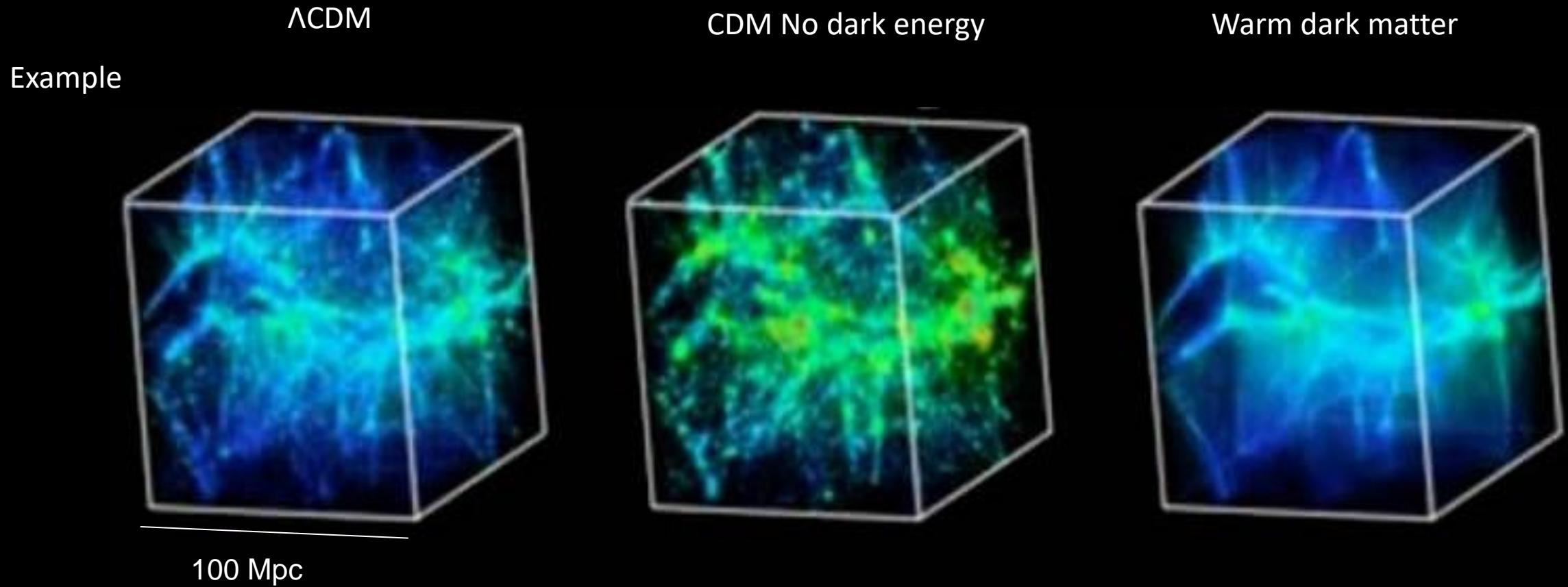


We live in a
95% **dark** universe

The Cosmic Web



The cosmic web: a gold mine for cosmology and fundamental physics

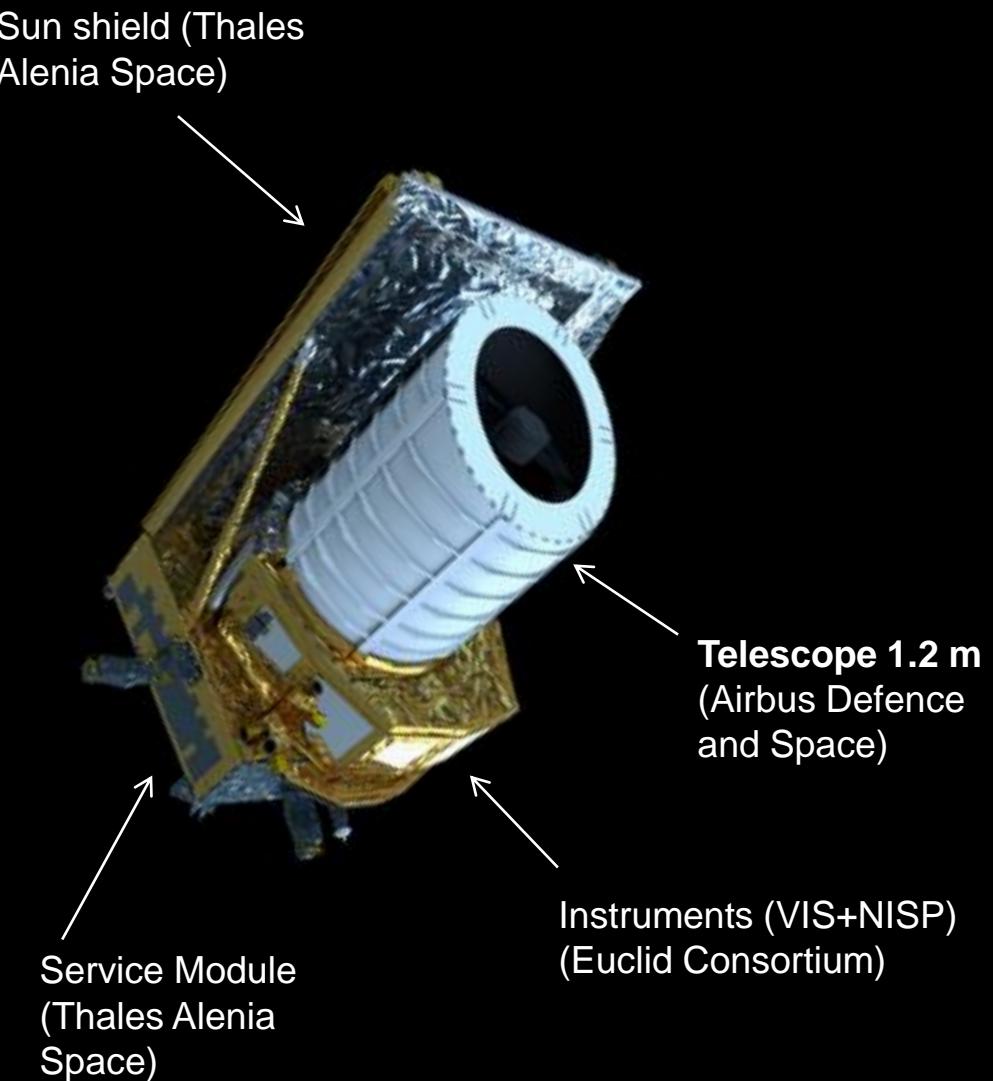




Euclid

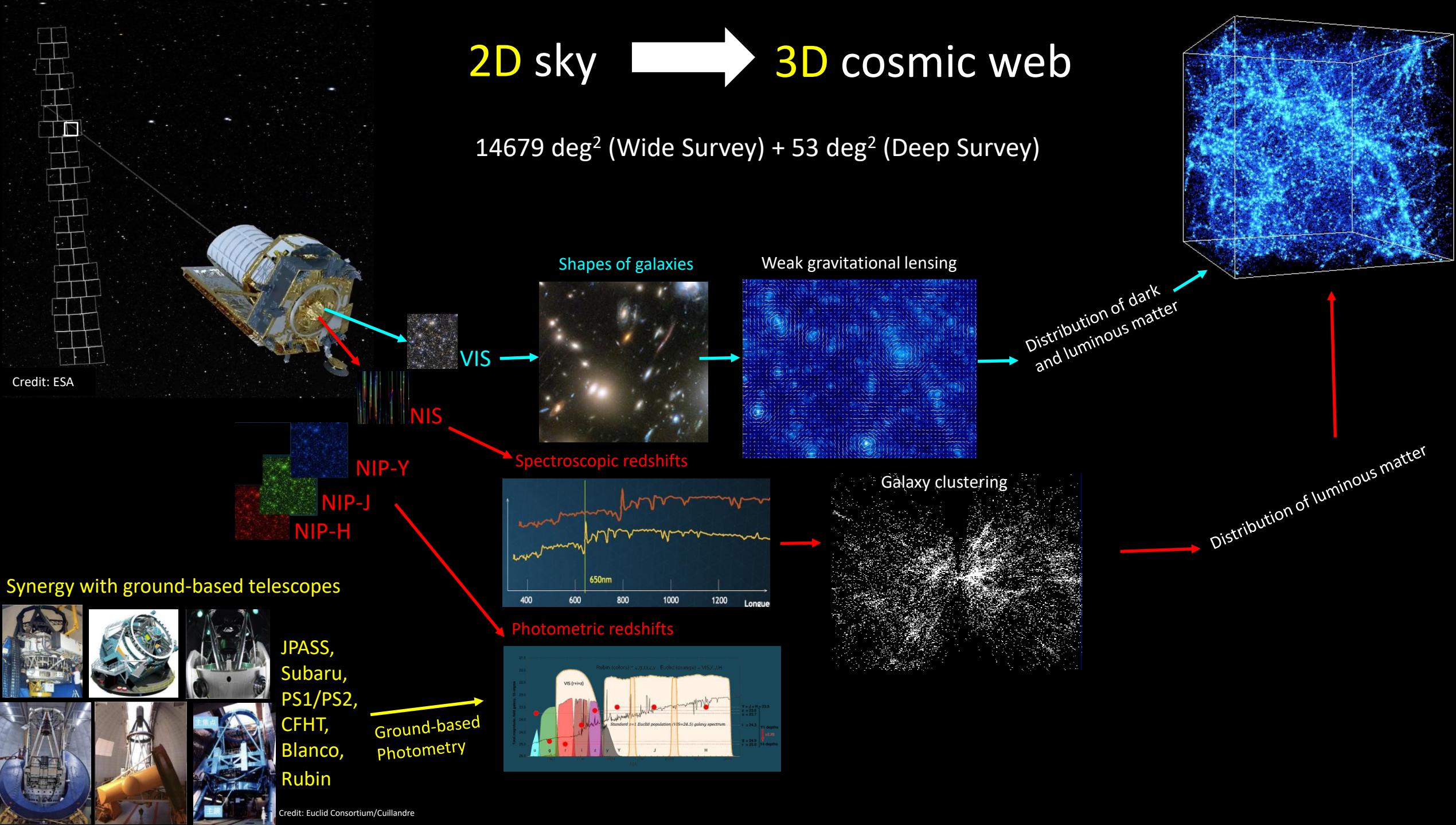
SPACE (PI A. Cimatti) + **DUNE** (PI A. Refregier)
mission concepts - ESA *Cosmic Vision* call (2007)

- 2008 – 2009: Assessment Phase
- 2010 – 2011: Definition Phase
- 2012: Adoption
- 2015: PDR → construction
- 2018: CDR
- 2023: launch (July 1st; L2 orbit)
- Duration: ≥ 6 years
- ESA + Euclid Consortium + NASA + CSA + Japan + Industries



2D sky → 3D cosmic web

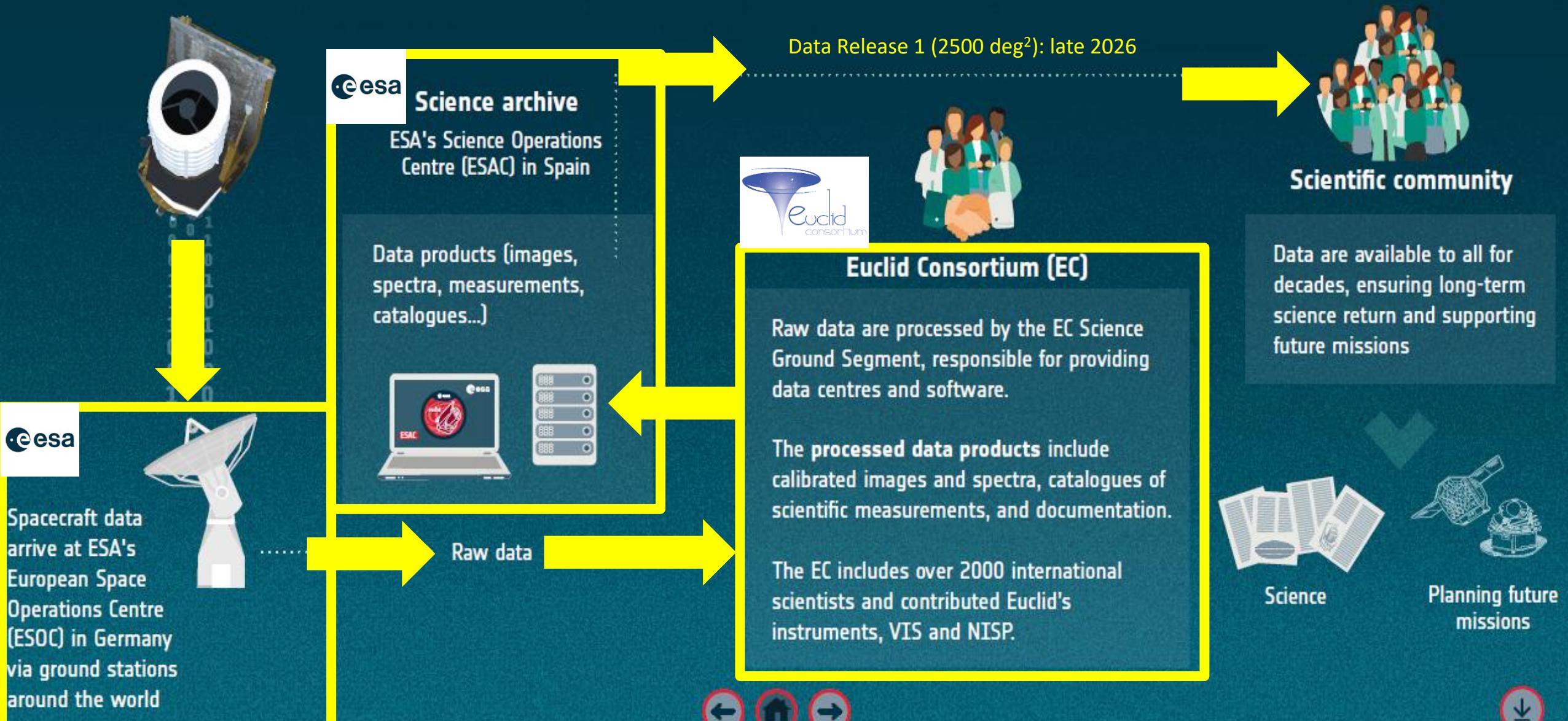
14679 deg² (Wide Survey) + 53 deg² (Deep Survey)

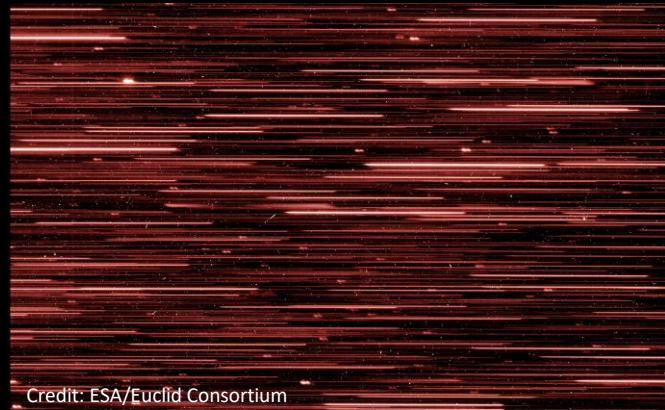


Euclid Ground Segment

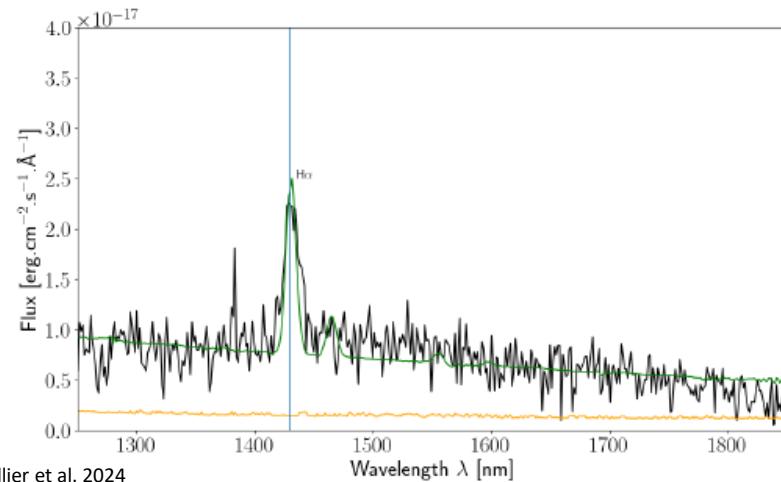
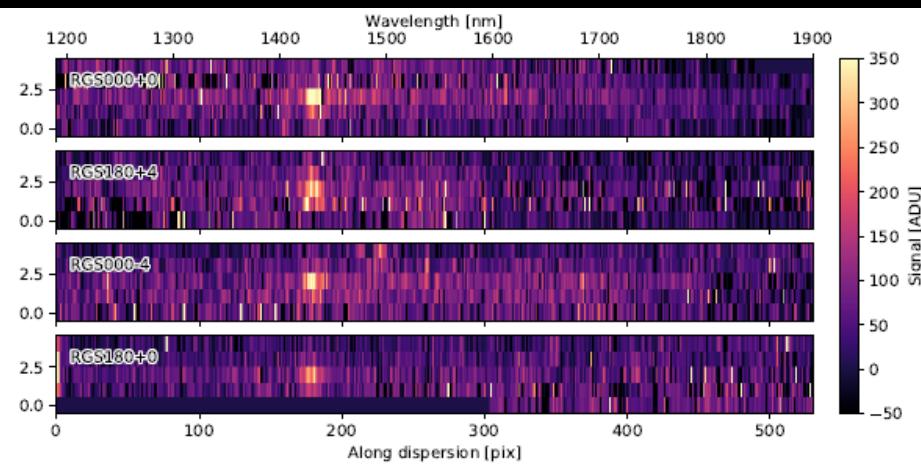


ASTRONOMY SCIENCE ARCHIVE: MAXIMISING SCIENCE FROM OUR MISSIONS

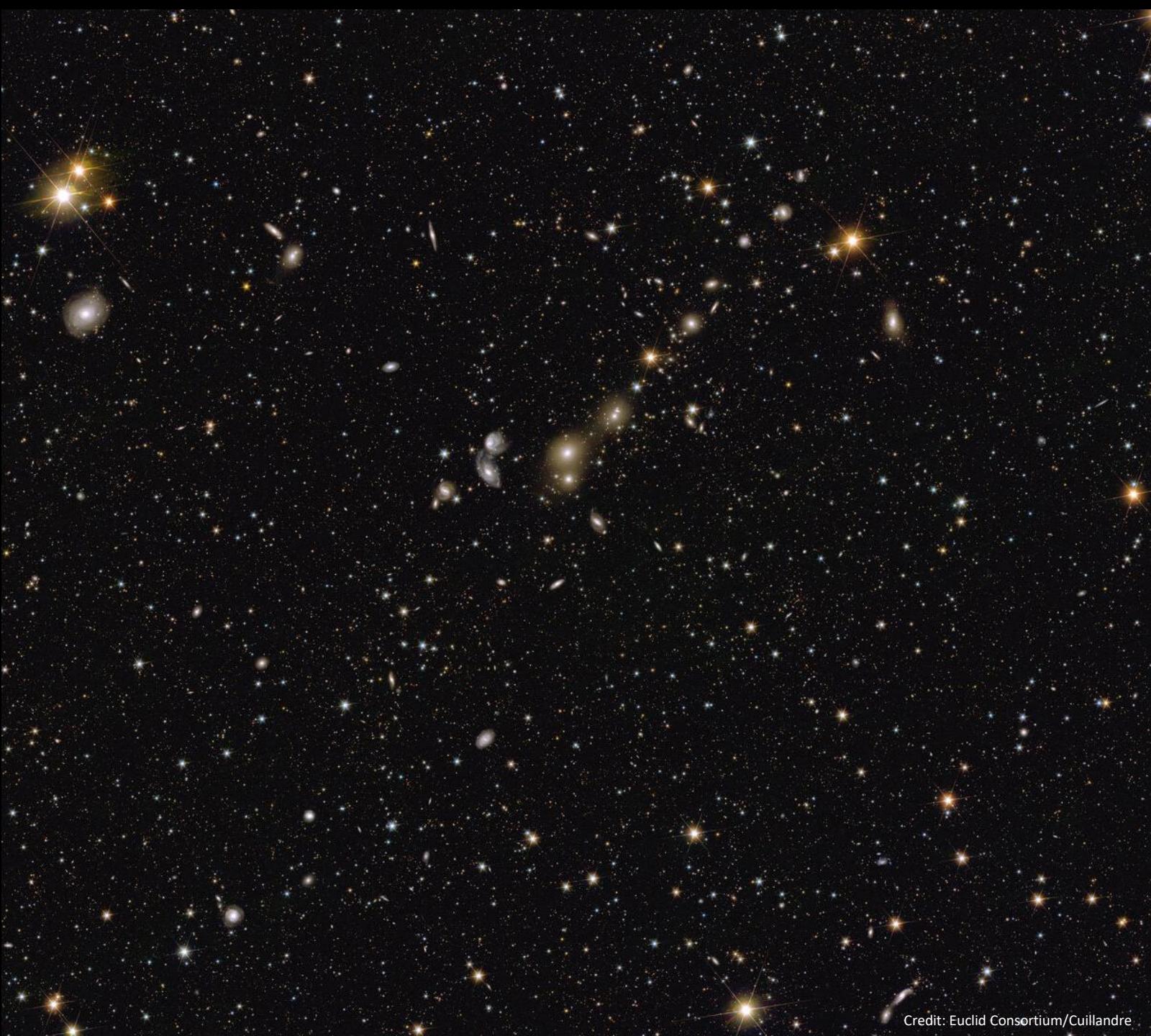




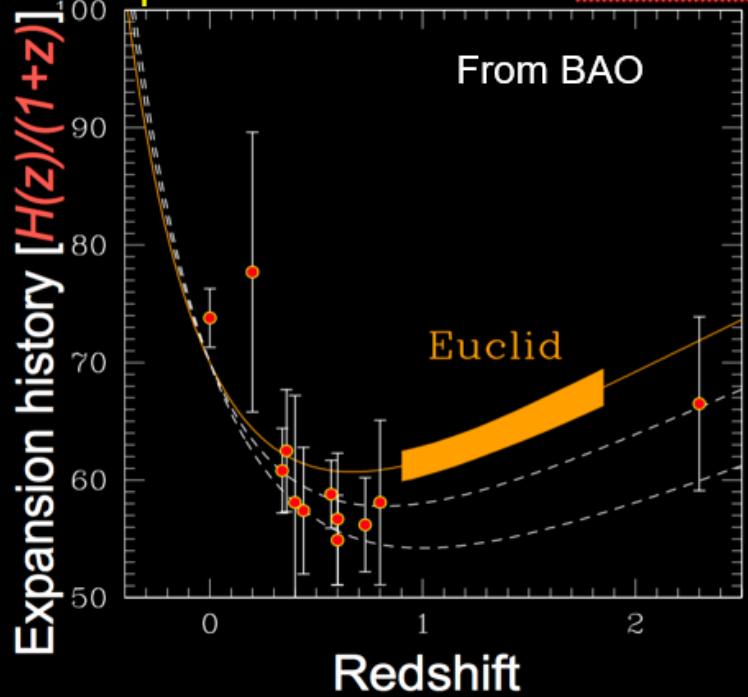
First Data



Mellier et al. 2024



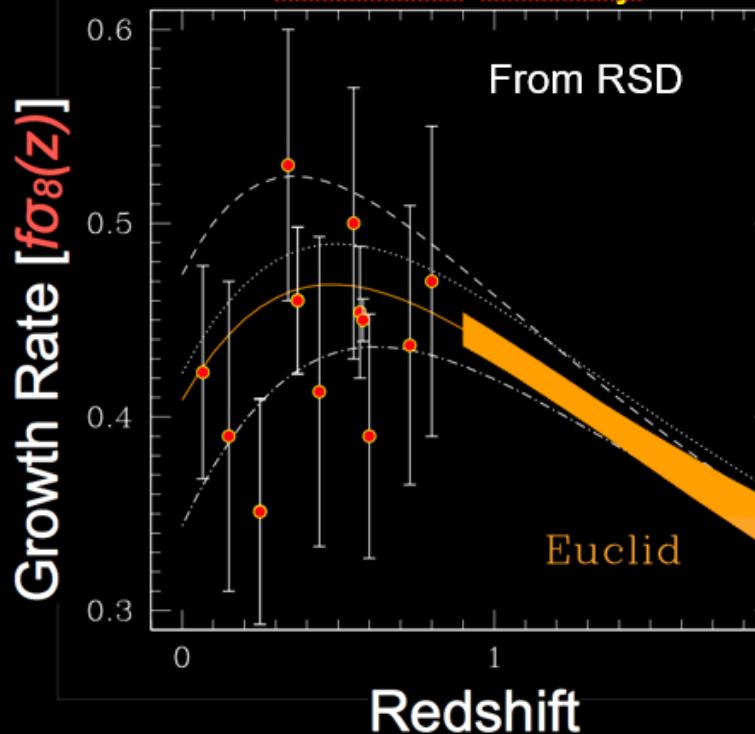
Expansion Rate of the Universe



The Power of Euclid

- Multiple cosmological probes
 - Weak lensing
 - Galaxy clustering
 - CMB cross-correlations
 - Clusters of galaxies
 - Strong lensing
- Mitigation of systematics
- Mitigation of degeneracies
- Improvement of 1-2 dex
- Immense legacy value

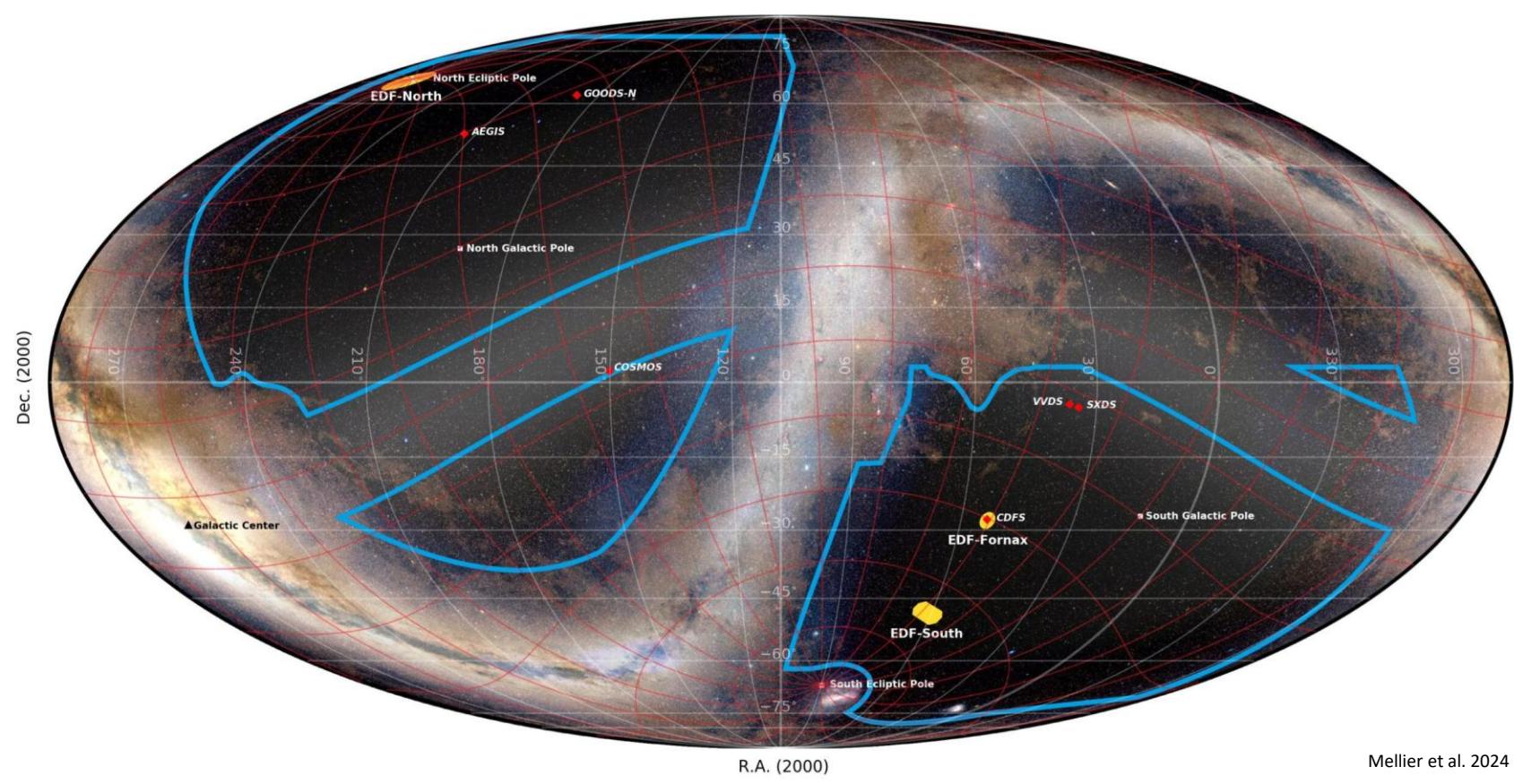
Modified Gravity



Part II

Euclid and ET Synergies

Euclid Surveys and Data Products



Wide Survey (extragalactic sky): **14679 deg²**

- 12 billion sources with $m_{AB} < 24.4$ (NISP) and < 26.7 (VIS) at 5σ
- Photometric SEDs: $ugriz$ (ground) + YJH (NISP)
- > 30 million spectroscopic redshifts mostly at $z < 1.9$ with H α ($z < 2.7$ with [O III]5007)
- Weak gravitational lensing shear maps
- Galaxy clusters: $> 10^5$, $M > 10^{14} M_\odot$, up to $z \sim 2$
- Morphologies of galaxies (VIS imaging)

Deep Survey: 53 deg²

- 3 sky fields
- 2 magnitudes (6.3x) deeper
- Spectroscopic redshifts up to $z \sim 7$

The role of the Euclid dataset

- Mission completion by 2030 → good timing for ET
- Space-based data (deep, homogeneous, stable)
- Extended to high redshifts (spectro-z + photo-z)
- **Selected in the near-IR:** mass-selected, all galaxy types, less affected by dust extinction, less biased
- Galaxy physical properties and morphologies
- Galaxy clusters
- Nothing comparable before ET

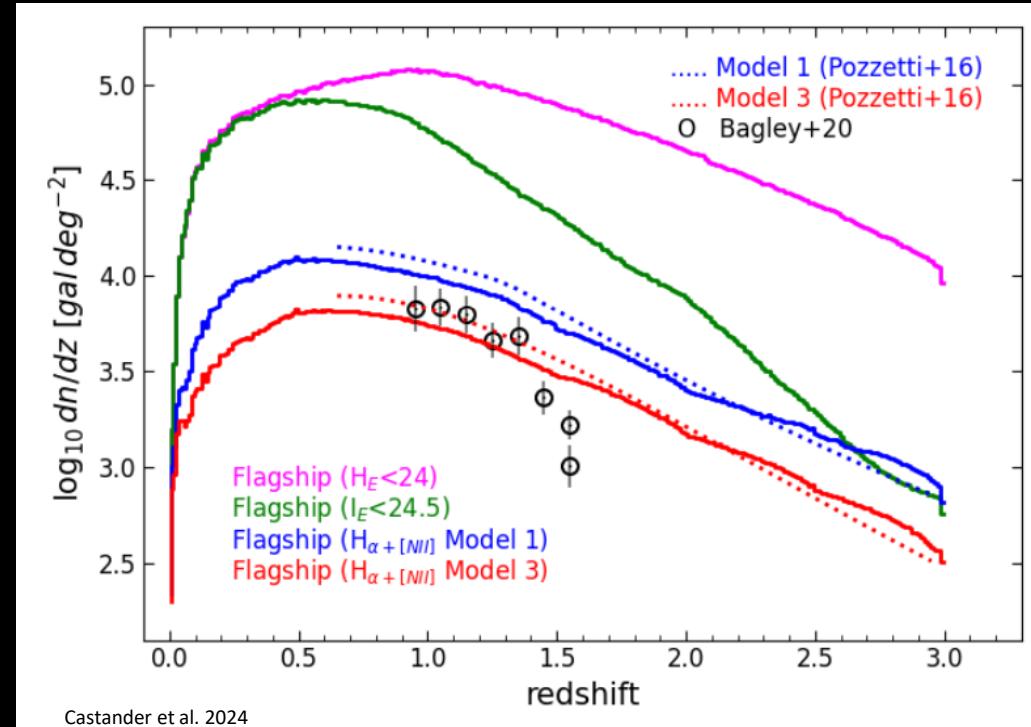


Fig. 36. Expected number densities for photometric and spectroscopic emission line flux selected samples as a function of redshift in the *Euclid* Wide survey. Different colours show various selection cuts for the EWS (in I_E , H_E and in the 2 model calibrations for the line fluxes of $f_{H\alpha+[NII]} > 2 \times 10^{-16} \text{ erg s}^{-1} \text{ cm}^{-2}$). The empirical model of Pozzetti et al. (2016) and the data based on slitless HST spectroscopy (Bagley et al. 2020) are also shown.

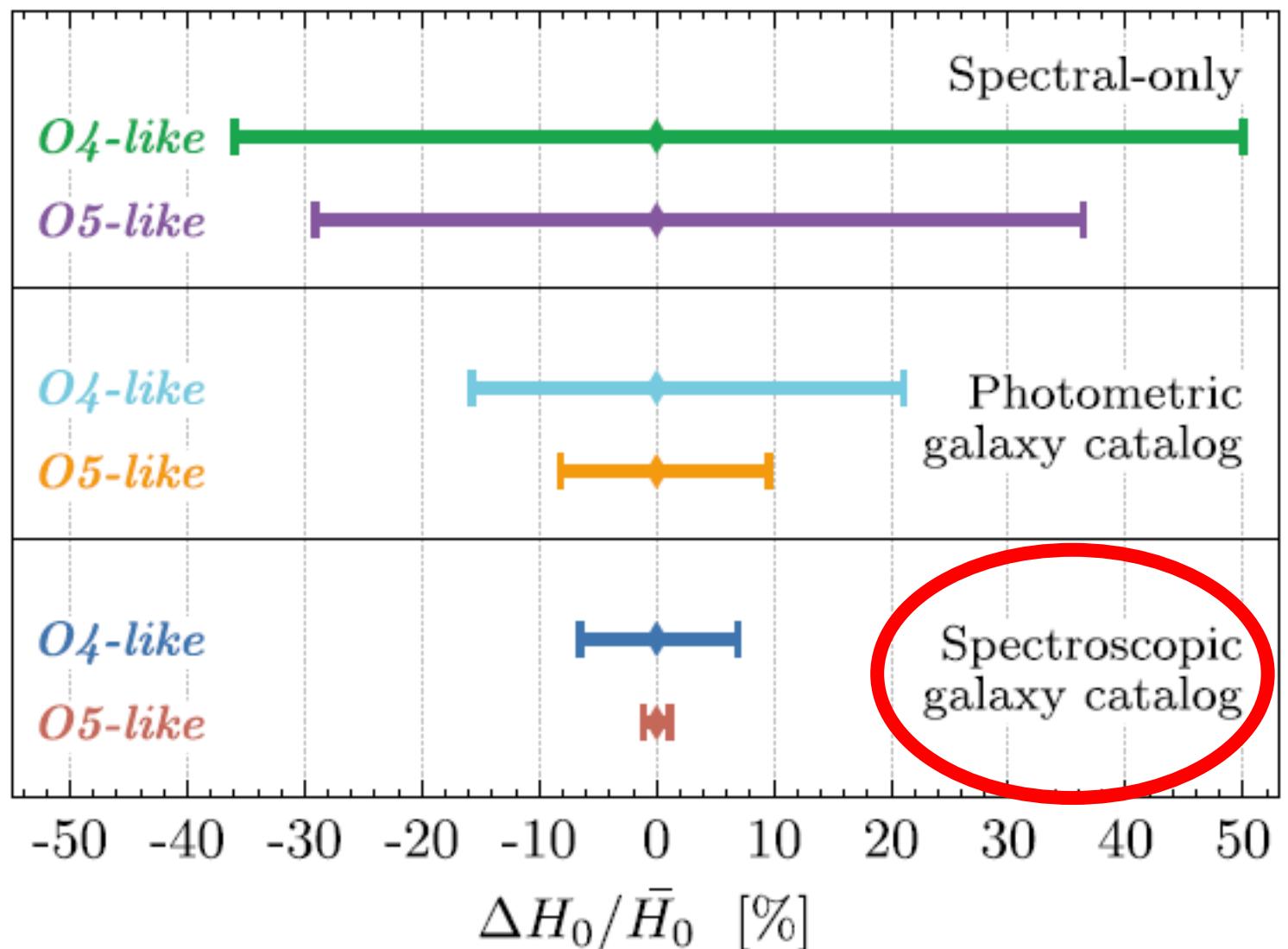
Some Euclid-ET Synergies

- **GWs with EM counterpart**: host galaxies in a wide redshift range over 1/3 of the sky
- **GWs without EM counterpart**: cosmology with dark sirens and galaxy catalogs
- **Resolved GW**
 - cross-correlations with galaxy clustering
 - cross-correlations with weak gravitational lensing maps
- **Unresolved GW**: cross-correlations of Euclid data with unresolved GW background

Euclid Consortium new **Science Working Group on Euclid and Gravitational Waves**

Co-leads: Michele Moresco (Uni. Bologna - DIFA), Miguel Zumalacarregui (MPI for Gravitational Physics)

Example



Hubble constant
with LVK O4-O5-like

Figure 5. Relative uncertainty on H_0 obtained from spectral-only and full standard sirens analysis of 100 BBHs in the O4- and O5-like network configurations, including a complete galaxy catalog with photometric or spectroscopic redshift uncertainties.

Borghi et al. 2024

A Schematic Outline of the Cosmic History

