# ETO Design Task Force update

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on behalf of the ETO Task Force on detector layout

ET Symposium - Bologna, May 29, 2025



## Scope of Task Force

#### Background

- 2020 ESFRI proposal
- 2022 CDR update
- 2024 optical layout update (ETC-ISB)
  - Triangle first, then 2L
- 2024 detector layout (ETO-ED)
  - Triangle first, then 2L
- Informal feedback by civil engineering experts from local teams: infrastructure to host updated detector layout would be significantly more expensive than for the ESFRI proposal

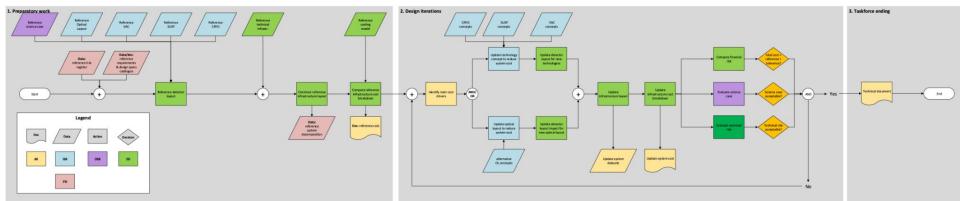
#### Task Force mandate

- Adapt detector layouts of ET towards an acceptable preliminary costing for the civil infrastructure (which is expected to dominate the total cost of ET)
  - for both Triangle and 2L geometries equally and independently
  - while maintaining ET's scientific performance



# Task Force study logic

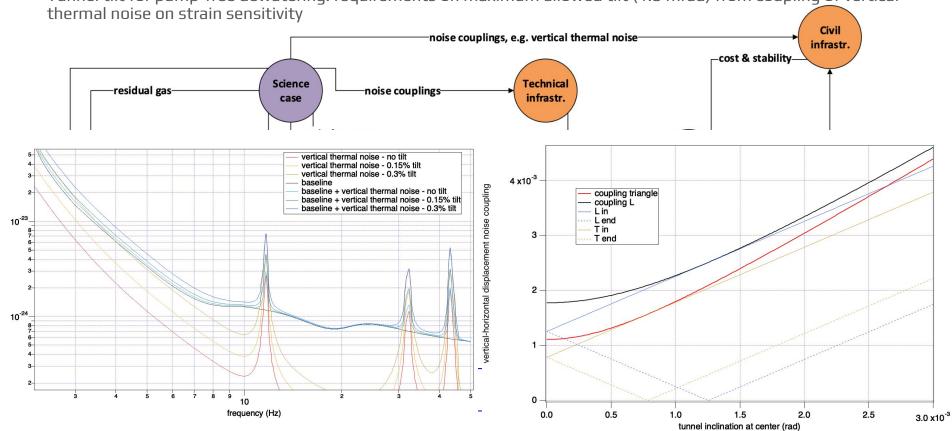
- Build a simplified picture of the ET detector limited to subsystems conveying major volume claims
- Develop dedicated tools for science case, detector layout, civil engineering
- Identify detector <-> infrastructure interfaces
- Identify main cost drivers for civil infrastructure
- Identify viable configurations to preserve performance and reduce cost
  - optical layout
  - critical technologies
  - independently for Triangle and 2L geometry
- comparative analysis of technical risk and infrastructure cost between identified configurations and 2024 reference



Basic decomposition, interfaces & expertise Civil -noise couplings, e.g. vertical thermal noiseinfrastr. cost & stability Science **Technical** residual gasnoise couplings infrastr. case optical response--space claims residual gas Detector size of caverns and tunnels lavout **Optical** inter-optics distance & beam sizetowers location layout Missing: arm pipes experts (CERN) Vacuum pipes length and diameter pipes tower size cryopumps Integrated Science (OSB, optics co-locationtowers ISB) thermal noise, residual gas-**Optics (ISB)** mirror size-Vacuum -vessel size and access-**Technologies** tanks Cryogenic (ISB) -crvostat size payload Engineering (ETO, local Seismic -seismic noisetank height isolation teams, CERN

## Example of engineering requirements

Tunnel tilt for pump-free dewatering: requirements on maximum allowed tilt (1.5 mrad) from coupling of vertical thermal noise on strain sensitivity



## Example - volume claims vs HFI tower height

Einstein Telescope Organisation

#### 1. Updated L Layout

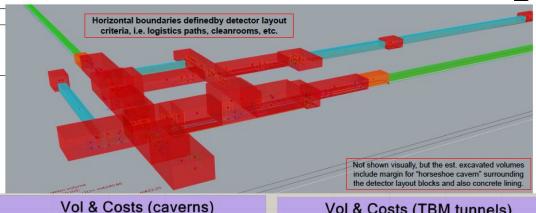


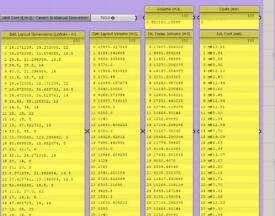
Est. total excavation volume
of caverns = 321150 m3;
Est. total cost of caverns (1L)

Est. total excavation volume
of conventional tunnels = 29662 m3;
Est. total cost of conventional tunnels (1L)=

Est. total excavation volume
of TBM tunnels = 1137241 m3;
Est. total cost of TBM tunnels (1L)=

Not perfect, but close enough to previous manual work (i.e. revit block model → excel tables)

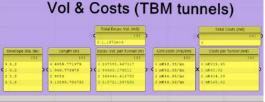


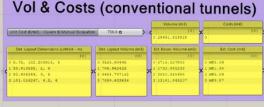


29 12412.487168

29 859.31

9 7090,960524





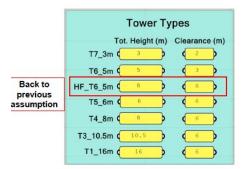


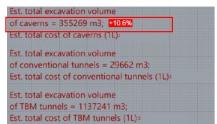


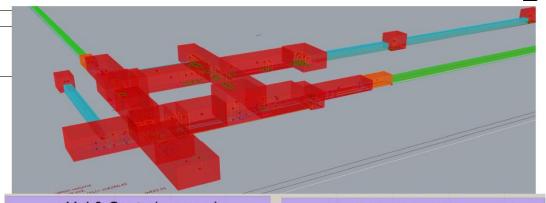
## Example - volume claims vs HFI tower height

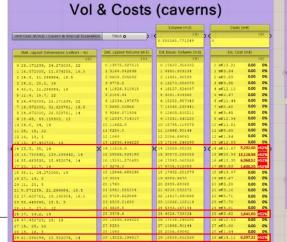
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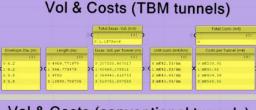
#### 3. Reverted HF towers

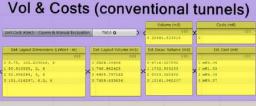










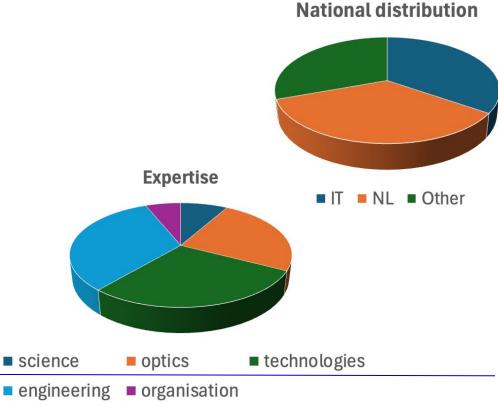






## Task Force composition

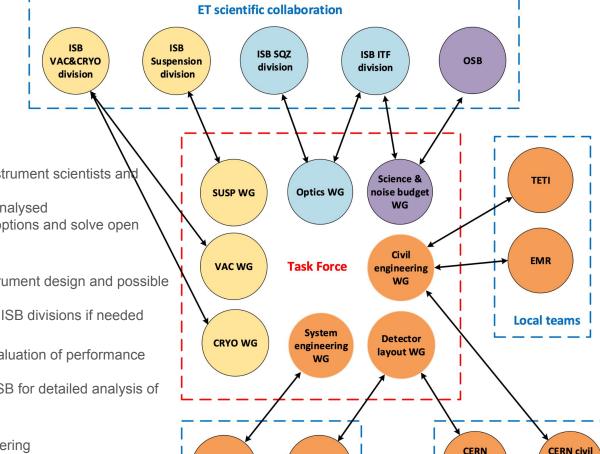
- 24 core team members
- 25 consultants
- National distribution
  - 35% from Italy
  - 35% from Netherland & Belgium
  - 30% from other countries
- Expertise distribution
  - 8% science case & noise budget (ETC ISB & OSB)
  - 24% optics (ETC ISB)
  - 29% instrument technologies (ETC ISB)
  - 33% engineering (ETO PD, ETO PO, local teams, CERN)
  - 6% organisation & management (ETO)





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# Working groups and external interactions



Engineering

Department

**ETO** 

**Project Office** 

beampipe

team

engineering

team

**CERN** 

- Plenary Task Force meetings bring together instrument scientists and engineers;
  - outcome include design options to be analysed
  - asynchronous work to analyse design options and solve open questions for next iteration
- Task Force members from ISB
  - provide information about baseline instrument design and possible alternative options
  - o share part of the open questions within ISB divisions if needed
- OSB liaisons
  - Provide prompt feedback for coarse evaluation of performance risk during configuration search
  - Carry out extensive analysis work to OSB for detailed analysis of science case on selected options
- Local teams liaisons
  - Share and refine criteria for civil engineering
  - Identify and discuss most relevant engineering requirements on infrastructure

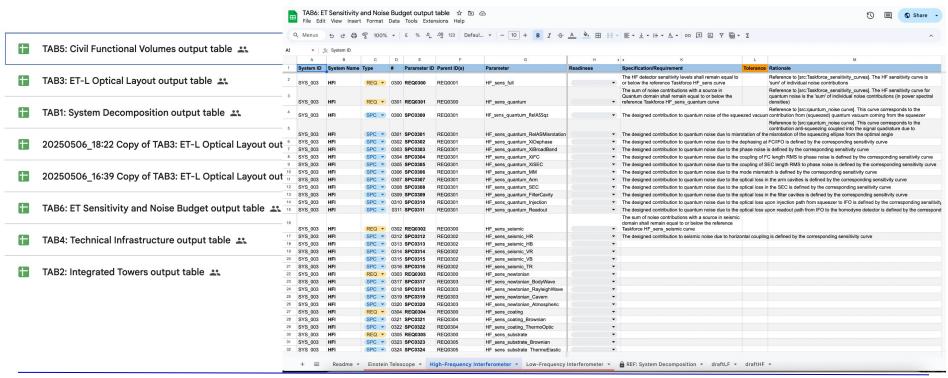
## Overview of Task Force work

- Started on January 2025
- Weekly plenary meetings minutes on ET Wiki
- Aperiodic meetings with subgroups of experts (~2/week)
- In-person workshops minutes on ET Wiki
  - 1st workshop methods consolidation & 2L layout update Pisa, February 18÷20
  - o 2nd workshop 2L layout consolidation Amsterdam, March 18÷20
  - o 3rd workshop Triangle layout update CERN May 5÷7
- Output delivered to external review team on 23/05
  - o available on Gitlab to ETC
  - final version after review ~end of June



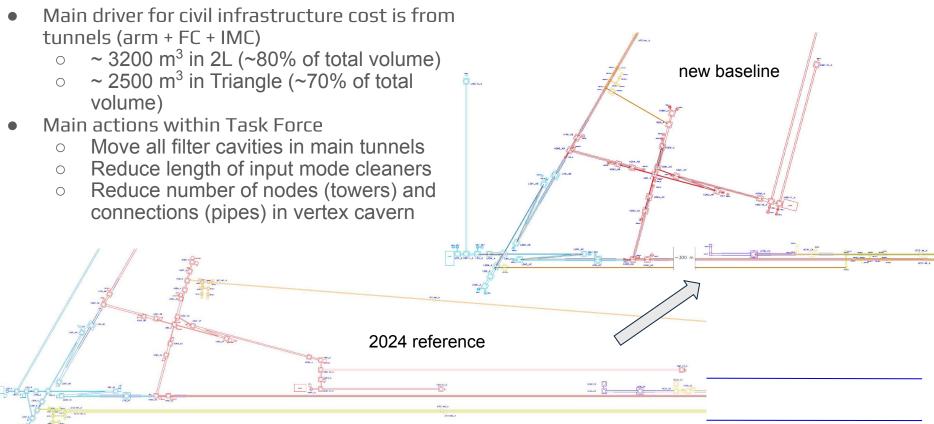


## System decomposition - output tables





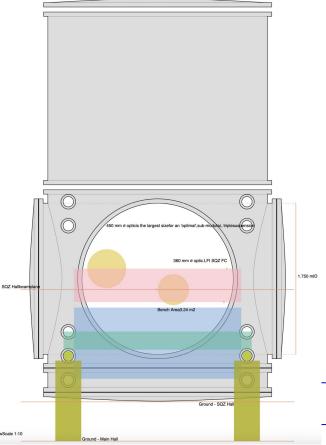
# Optical layout update



Integrated towers

- Identified main interfaces between design elements determining height and footprint of detector layout nodes
  - Seismic isolation and payload
  - Vacuum tank and access type
  - Cryostat and cryogenic payload for LF\_TM
- Seismic isolation:
  - Improved categorisation
  - reduction of benches footprint and tower height wherever convenient
    - 29% reduction in overall caverns volume
- Tank access:
  - Identification of main constraints from optical layout, suspension design, cryogenic system
  - Identification of interface with technical infrastructure (clean rooms)
- Cryogenic systems:
  - Identification of available design options
  - Assessment of impact on detector layout and on civil infrastructure





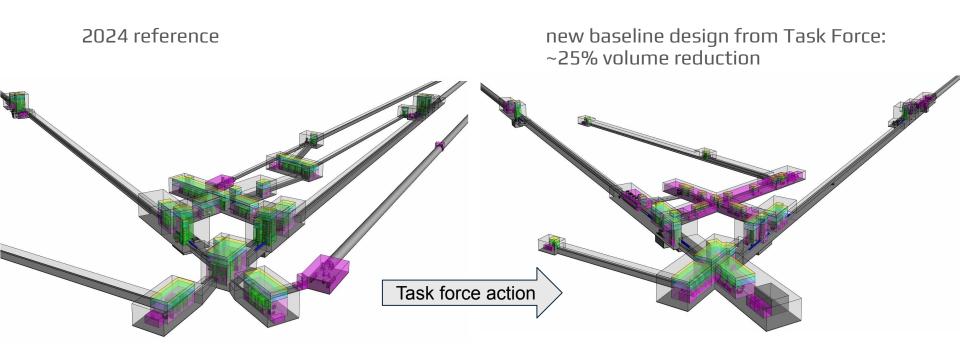
## Flexibility envelope & flexibility demands

• Upgrade over 2024 reference optical layout: extended flexibility envelope for optical layout, included basic demands from optical layout, introduced flexibility envelope for detector layout

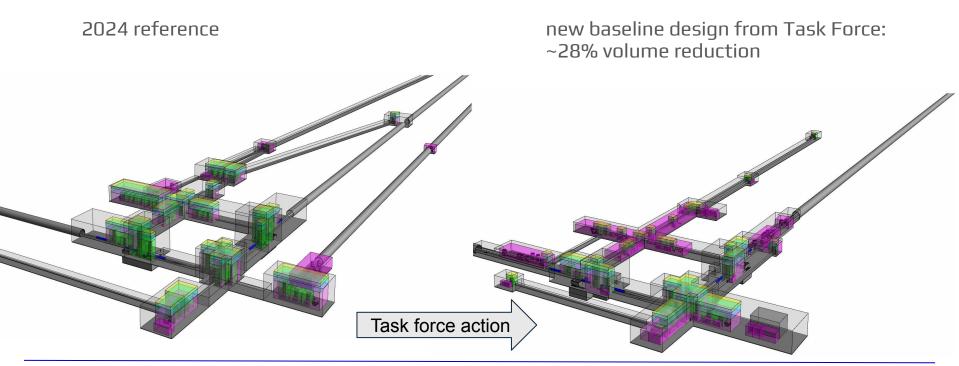
3A: general - 3B: layout - 3C: optic specs - 3D: flexibility demands - old HF-SQZ - old LF-SQZ - 🔒 REF: System Decomposition -

| System | Location |    | Optical Element Type | OLD I | *Parameter (auto-generated)    | Parameter Type    | Primary optical<br>surface or group of<br>optics | Secondary optical<br>surface or group (if<br>required) | Value | unit | Tolerance | Readiness         | Flexibility Envelope   | OLD Specification/Requirement (by-hand)  |     |
|--------|----------|----|----------------------|-------|--------------------------------|-------------------|--|--|-------|------|-----------|-------------------|------------------------|--|-----|
| HF ▼   | ARMY     | •  | Single optic         | HF-ET | HF ETMY-HR orientation         | orientation -     | ETMY-HR  | ITMY-HR  | 0     | deg  |           | experimental v •  | none 🔻                 | HF-ETMY shall have a normal AOI with relation to HF-ITMY   |     |
| HF ▼   | RCY      | •  | Single optic ▼       | HF-Z) | HF ZY3-HR orientation          | orientation -     | ZY3-HR   | ZY2-HR   | 45    | deg  |           | analytical or n ▼ | major redesign 💌       | The AOI at HF-ZY3 to HF-ZY2 shall be 45 degrees  |     |
| HF ▼   | RCY      | *) | Single optic         | HF-Z) | HF ZY1-HR orientation          | orientation -     | ZY1-HR   | ZY2-HR   | 2.5   | deg  |           | analytical or n   | major redesign 💌       | The AOI at HF-ZY2 to HF-ZY1 shall be 2.5 degrees   |     |
| HF ▼   | ARMX     | *  | Single optic         | HF-E7 | HF ETMX-HR orientation         | orientation       | ETMX-HR  | ITMX-HR  | 0     | deg  |           | experimental v    | none                   | HF-ETMX shall have a normal AOI with relation to HF-ITMX   |     |
| HF ▼   | RCX      | •  | Single optic ▼       | HF-Z) | HF ZX2-HR orientation          | orientation •     | ZX2-HR   | ZX3-HR   | 45    | deg  |           | analytical or n   | major redesign •       | The AOI from HF-ZX3 to HF-ZX2 shabe<br>45 degrees  | all |
| HF ▼   | RCX      | *  | Single optic ▼       | HF-Z) | HF ZX1-HR orientation          | orientation       | ZX1-HR   | ZX2-HR   | 2.5   | deg  |           | analytical or n ▼ | major redesign 💌       | The AOI from HF-ZX2 to HF-ZX1 sha<br>be 2.5 degrees  | al  |
| HF ▼   | RCS      | •  | Single optic ▼       | HF-ZS | HF ZS1-HR orientation          | orientation       | ZS1-HR   | ZS2-HR   | 1.56  | deg  |           | analytical or n   | major redesign 💌       | The AOI from HF-ZS1 to HF-ZS2 shabe<br>1.56 degrees  | al  |
| HF ▼   | RCS      | •  | Single optic ▼       | HF-SF | HF SEM-HR orientation          | orientation       | SEM-HR   | ZS1-HR   | 1.56  | deg  |           | analytical or n   | major redesign 🔻       | The AOI from HF-SEM to HF-ZS1 shall be 1.56 degrees  |     |
| HF ▼   | RCP      | •  | Single optic ▼       | HF-ZF | HF ZP1-HR orientation          | orientation       | ZP1-HR   | ZP2-HR   | 2.44  | deg  |           | analytical or n   | major redesign 🔻       | The AOI from HF-ZP1 to HF-ZP2 shabe<br>2.44 degrees  | a   |
| HF ▼   | RCP      | •  | Single optic ▼       | HF-PF | HF PRM-HR orientation          | orientation •     | PRM-HR   | ZP1-HR   | 2.44  | deg  |           | analytical or n ▼ | major redesign 💌       | The AOI from HF-PRM to HF-ZP1 shall be 2.44 degrees  |     |
| HF ▼   | global   | •  | Group of optics ▼    | HFI_c | HF unresolved - requires input | unresolved - r 🔻  |  |  |       | ?    |           | experimental v 🔻  | none 🔻                 | HFI X- and Y-cavities shall have<br>symmetrical optical paths  |     |
| HF ▼   | RCS      | ~  | Single optic ▼       | HF-ZS | HF unresolved - requires input | unresolved - r    |  |  |       | ?    |           | analytical or n   | major redesign 🔻       | HF-ZS2 shall be along the same axis as HF-BS and HF-ZY1  | *** |
| HF ▼   | RCP      | *  | Single optic         | HF-ZF | HF unresolved - requires input | unresolved - r    |  |  |       | ?    |           | analytical or n ▼ | major redesign 💌       | HF-ZP2 shall be along the same axis as HF-BS and HF-ZX1  | -   |
| HF ▼   | IN       | •  | Group of optics      | HF-PS | HF PSL room footprint          | footprint (shape) | PSL room   |  | 8x10  | m    |           | assumption        | freedom on orientation | HF-PSL room surface area shall be a least 6x10m, with additional 2x10m along the side for hosting laser- & input/output electronic racks | N   |
| HF ▼   | OUT      | •  |                      |       | HF unresolved - requires input | unresolved - r    |  |  |       | ?    |           | assumption        | •                      | HF-SQI shall be in-line with HF-SEM and HF-ZS1   |     |
| HF ▼   | ARMY     | •  | Single optic         | HF-ET |                                | position          | ETMY-HR  | ITMY-HR  | 15000 | m    |           | analytical or n   | major redesign 💌       | The distance from HF-ETMY to HF-ITMY shall be 15000m   |     |
|        |          |    |                      |       |                                |                   |  |  |       |      |           |                   |                        | The distance along the optical path  |     |

## Detector layout - Triangle



## Detector layout - 2L





## Main outcomes from Task Force

- The ETO task force produced a new **baseline detector layout** for both triangle and 2L geometries
- Main features in comparison with 2024 reference detector layouts
  - Basic system decomposition interfaces and requirements flow between instrument & infrastructure
  - Updated optical layouts
    - Filter cavities in arm tunnels, reduced length for IMC tunnels, simplified vertex design
    - Update of flexibility envelope, assessment of flexibility demands
  - Full classification of integrated towers, with reduction of benches footprint and tower height
  - Detector layout update
    - Better definition of technical infrastructure (cryogenics, clean rooms, noisy rooms)
    - Assessment of flexibility envelope
  - Interface with civil engineering
    - Identification of main cost drivers, estimate relative cost changes vs detector layout configuration
    - Identification of main engineering requirements on technical and civil infrastructure
  - Risk and flexibility analysis on design choices and alternative configurations
  - Noise budget, comparison with the official ET science case, scientific requirements on design parameters
  - Volume claims and infrastructure cost reduced by ~25% from 2024 reference layouts
- The new baseline detector layout corresponds to a **cost-effective** baseline design for the ET instrument
  - Further minor reduction of infrastructure cost may be achieved at the cost of substantial technical risk
  - Substantial reduction of infrastructure cost may only be achieved by descoping of the ET science case



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## Output documentation

#### Main document (74 pages): baseline detector layout

date: May 24, 2025

**Extended supporting document** Technical annexes (33 additional files): tables, (197 pages): details and study logic 2D&3D drawings, technical specifications

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The ET Baseline Detector Layout

| 0   | Supporting Document for The ET Baseline Detector Layout date: May 24, 2025 page: 1 of 197            |  |  |  |  |  |  |  |  |  |
|-----|--|--|--|--|--|--|--|--|--|--|
| 0   | ontents  |  |  |  |  |  |  |  |  |  |
| C   | ontents  |  |  |  |  |  |  |  |  |  |
| 1   | Study logic and workflow   |  |  |  |  |  |  |  |  |  |
| 2   | Background documentation for the ETO Design Task Force External Review Committee                     |  |  |  |  |  |  |  |  |  |
| 3   | Optical layout   |  |  |  |  |  |  |  |  |  |
|     | 3.1 Methodology of the Optical Layout  |  |  |  |  |  |  |  |  |  |
|     | 3.2 Core Optical Layout (DRFPMi)   |  |  |  |  |  |  |  |  |  |
|     | 3.3 Squeezing subsystem  |  |  |  |  |  |  |  |  |  |
|     | 3.4 Input and Output Optical Systems   |  |  |  |  |  |  |  |  |  |
|     | 3.5 Auxiliary Systems  |  |  |  |  |  |  |  |  |  |
|     | 3.6 Flexibility Demands of the Optical Layout  |  |  |  |  |  |  |  |  |  |
|     | 3.7 Guide to the Optical Layout Output Tables  |  |  |  |  |  |  |  |  |  |
| 4   | Integrated towers  |  |  |  |  |  |  |  |  |  |
|     | 4.1 Context and extended summary of Tower Categorization   |  |  |  |  |  |  |  |  |  |
| - 8 | 4.2 Tower categorization outcomes - Triangle and 2L geometry   |  |  |  |  |  |  |  |  |  |
|     | 4.3 Main design options for seismic isolation  |  |  |  |  |  |  |  |  |  |
|     | 4.4 Main design options for cryogenics   |  |  |  |  |  |  |  |  |  |
|     | 4.5 Main options for tower access  |  |  |  |  |  |  |  |  |  |
| 5   | Vacuum pipes   |  |  |  |  |  |  |  |  |  |
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| 0   | Detector Layout  |  |  |  |  |  |  |  |  |  |
| U   | 6.1 Explanation of major space claims  |  |  |  |  |  |  |  |  |  |
|     | 6.2 Optional detector layouts  |  |  |  |  |  |  |  |  |  |
|     | 0.2 Optional desector injours  |  |  |  |  |  |  |  |  |  |
| 7   | Civil engineering 1  |  |  |  |  |  |  |  |  |  |
|     | 7.1 Parametric tool for determining relative cost of civil infrastructure vs detector layout changes |  |  |  |  |  |  |  |  |  |
|     | 7.2 Criteria for determining relative cost of civil infrastructure vs detector layout changes        |  |  |  |  |  |  |  |  |  |
|     | 7.3 Cost Estimation Methodology  |  |  |  |  |  |  |  |  |  |
|     | 7.4 Element Composition  |  |  |  |  |  |  |  |  |  |
|     | 7.5 Technical Requirements   |  |  |  |  |  |  |  |  |  |
| 8   | Risk and flexibility   |  |  |  |  |  |  |  |  |  |
|     | 8.1 Extended explanation of risk Analysis 2L Geometry: Alternative options not included in the       |  |  |  |  |  |  |  |  |  |
|     | Baseline 2025 Task Force   |  |  |  |  |  |  |  |  |  |
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| 10  | 10.1 Technical drawings  |  |  |  |  |  |  |  |  |  |
|     | 10.1 Technical drawings<br>10.2 Plots  |  |  |  |  |  |  |  |  |  |
|     | 10.2 Piots<br>10.3 Tables  |  |  |  |  |  |  |  |  |  |
|     | 10.3 Tables 10.4 Other external documents  |  |  |  |  |  |  |  |  |  |
|     | 10.4 Other execute quemients   |  |  |  |  |  |  |  |  |  |





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## Conclusions

- Civil & technical infrastructure are essential components of the ET project
  - instrument and infrastructure design are closely interconnected
- In 4.5 months the Task Force produced a first baseline layout of the ET detector accounting for interfaces with civil and technical infrastructure
  - o civil infrastructure would be ~25% cheaper than with 2024 reference detecto layout
- Extraordinary effort by team members
  - final outcome is larger and better than my expectations
  - o joint work of instrument scientists and engineers played a key role
- Material will be useful for engineering studies by local teams
- Methods can be useful for next steps to generate a full ET TDR
  - o but emergency operation should **not** be taken as an example

Spare slides

Task Force composition - core team

NL

IT

other

| Name                | institution           | main expertise                       |  |  |  |  |
|---------------------|-----------------------|--------------------------------------|--|--|--|--|
| Anna Green          | Nikhef                | optics                               |  |  |  |  |
| Antonio Perreca     | Trento University     | optics                               |  |  |  |  |
| Marco Vardaro       | Maastricht University | squeezing                            |  |  |  |  |
| Nathan Holland      | Nikhef                | Seismic isolation                    |  |  |  |  |
| Leonardo Lucchesi   | INFN Pisa             | Seismic isolation                    |  |  |  |  |
| Antonino Chiummo    | EGO                   | Optics, injection                    |  |  |  |  |
| Francesca Spada     | INFN Plsa             | Seismic isolation, suspended benches |  |  |  |  |
| Paolo Ruggi         | EGO                   | Seismic isolation                    |  |  |  |  |
| Julien Gargiulo     | EGO                   | Vacuum, cryogenics                   |  |  |  |  |
| Henk Jan Bulten     | Nikhef                | Cryogenics, seismic isolation        |  |  |  |  |
| Fulvio Ricci        | Roma 1 University     | Cryogenics                           |  |  |  |  |
| Angelo Cruciani     | INFN Roma 1           | Cryogenics                           |  |  |  |  |
| Jonathan Bratanata  | Nikhef                | Civil engineering                    |  |  |  |  |
| Max Majoor          | Nikhef                | Technical engineering                |  |  |  |  |
| Mikhail Korobko     | Hamburg University    | Optics, squeezing, noise budget      |  |  |  |  |
| Elena Licciardello  | INFN-LNS              | Civil engineering, TETI liaison      |  |  |  |  |
| Romano Meijer       | Nikhef                | System engineering                   |  |  |  |  |
| Ghada Mahmoud       | APC                   | Risk management                      |  |  |  |  |
| Benoit Tuybens      | Nikhef                | Organisation and documentation       |  |  |  |  |
| Fiodor Sorrentino   | INFN                  | Coordination                         |  |  |  |  |
| Ulyana Dupletsa     | GSSI                  | OSB liaison                          |  |  |  |  |
| Francesco lacovelli | Geneve University     | OSB liaison                          |  |  |  |  |
| Patricia Lamas      | Amberg Engineering    | EMR liaison                          |  |  |  |  |
| Tamara Alice Bud    | CERN                  | Civil engineering                    |  |  |  |  |

science

optics

technologies

engineering

organisation

## Task Force composition - consultants

| Name                      | institution           | main expertise     |  |  |  |
|---------------------------|-----------------------|--------------------|--|--|--|
| Tommaso Napolitano        | INFN                  | ETO ED             |  |  |  |
| Paolo Martella            | INFN                  | TETI Liaison       |  |  |  |
| Jan Vesely                | EMR                   | EMR liaison        |  |  |  |
| Tom Hundertmark           | EMR                   | EMR liaison        |  |  |  |
| Jerome Degallaix          | IN2P3                 | optics             |  |  |  |
| Daniel Brown              | Adelaide University   | optics             |  |  |  |
| Giacomo Ciani             | Trento University     | squeezing, optics  |  |  |  |
| Julia Casanueva           | EGO                   | controls           |  |  |  |
| Sebastian<br>Steinlechner | Maastricht University | Optics, general    |  |  |  |
| Conor Mow-Lowry           | Nikhef                | Seismic isolation  |  |  |  |
| Antonio Pasqualetti       | EGO                   | Vacuum, cryogenics |  |  |  |
| Steffen Grohman           | KIT                   | Cryogenics         |  |  |  |

| Name                | institution       | main expertise                  |
|---------------------|-------------------|---------------------------------|
| Ettore Majorana     | Roma 1 University | Cryogenics, payload             |
| Piero Rapagnani     | Roma 1 University | Cryogenics, payload             |
| Wissam Wahbeh       | Roma 1 University | Civil engineering               |
| Maria Marsella      | Roma 1 University | Civil engineering               |
| Patrick Werneke     | Nikhef            | Technical engineering           |
| Andreas Freise      | Nikhef            | Broad instrument expertise      |
| Riccardo de Salvo   | n.a.              | Broad instrument expertise      |
| Marco Galimberti    | EGO               | Optics, large infrastructures   |
| Lucia Lilli         | INFN Pisa         | Organisation and documentation  |
| Archisman Ghosh     | Gent University   | OSB liaison                     |
| Joseph Ickmans      | EMR               | EMR liaison                     |
| John Andrew Osborne | CERN              | Civil engineering               |
| Valeria Sequino     | INFN              | Optics, squeezing, noise budget |

## Configurations for Triangle and 2L

- New baseline: main changes from 2024 reference
  - LF Filter cavities in X arm with periscope
  - HF filter cavity in Y arm with periscope
  - 2-mirror FC -> reduced pipe diameter
  - Reduced length of LF IMC
  - Merging HF IMCs in same tunnel
  - Route BHD through BS
  - Other reshuffling in central area
  - Tower access constrained on LF\_TM (bottom), SQZ (lateral), and few other
    - in flexibility envelope otherwise
  - o Reduced LF TM susp. height to 13 m
  - Reduced tower height for other HFI optics
  - Reduce footprint of CAT1 benches

Alternative configurations

- Double cavern
- 2. No periscope for LF\_FC
- 3. Alternative routing for SQZ beam
- 4. Bow-tie IMC
- 5. Reduced tower height for HF TM
- 6. Reduced tower height for LFI optics
- 7. Reduced cryostat size

# Configuration changes for Triangle & 2L

Ratio of (cost-volume reduction in CI)/(risk & flexibility cost)

High

Medium

Low

- Optical layout
  - LF Filter cavities in X arm
    - with periscope
      - two alternative routings
    - without periscope
  - HF filter cavity in Y arm
  - 2-mirror FC -> reduced pipe diameter
  - Reduced length of LF IMC
  - Merging HF IMCs in same tunnel
  - Route BHD through BS
  - Other reshuffling in central area
  - Bow-tie IMC

- Integrated tower
  - Tower access constrained on LF\_TM (bottom), SQZ (lateral), and few other
    - in flexibility envelope otherwise
  - Reduced LF TM susp. height to 12 m
  - Reduced tower height for HF TM
  - Reduced tower height for other HFI optics
  - Reduced tower height for LFI optics
  - Reduce footprint of CAT1 benches
  - Reduced cryostat size
- Detector layout
  - Double cavern

# Analysis on configurations for Triangle & 2L

- Compare cost-volume reduction on CI and risk/flexibility cost of each option
- Validate preferred configuration in detail
  - combine all compatible green options
  - merge equivalent options into flexibility envelope
- Analysis on global configurations for main document
  - Compare 2024 reference with chosen configuration (new baseline)
    - separately for Triangle and 2L
  - consider an individual configuration for each considered change (green in the list)
  - for each configuration estimate change in
    - civil infrastructure cost
    - technical risk
    - Flexibility
- For extended document repeat on yellow and red configuration changes

# Flexibility envelope & flexibility demands Requirement: a required property of an element or a part of a subsystem

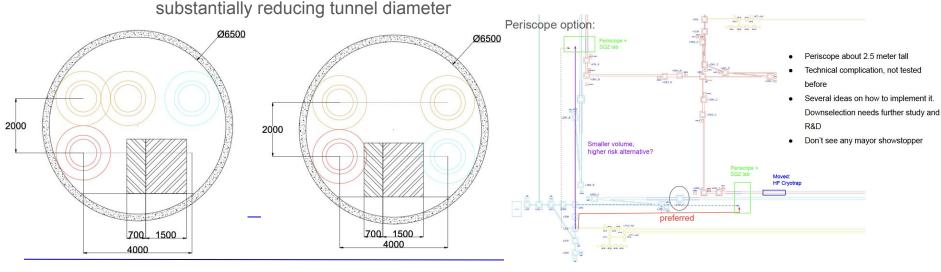
- - e.g. distance between 2 optical elements)
- **Tolerance**: the deviation that can be accepted for the element or subsystem to still fulfill its performance when constructed
  - e.g. deviation of the realised distance of the two optical elements compared to the design/nominal distance;
- **Flexibility Demand**: flexibility that is requested to be kept in the design at the current stage of the design process
  - e.g. we want to keep the possibility in the vacuum system / cavern design to be able to shift an optical element by e.g. 1 m sideways off the optical axis);
- **Flexibility envelope**: range of possible values that is given to the engineering team for optimisation
  - i.e. the optical design is based on a distance between two optical elements of e.g. 20 m (requirement) ±0.1 m (tolerance), but the optical design can be tuned to cope also if the distance is in the range from e.g. 15-25 m
- Upgrade over 2024 reference optical layout: extended flexibility envelope for optical layout, included basic **demands** from **optical** layout, introduced flexibility **envelope** for **detector** layout



## Filter cavity positioning

- Configurations for 2L geometry
  - LF filter cavities in one arm tunnel (+ HF filter cavity in the other arm tunnel)
    - Possible options:
      - FCs on different plane from HFI and LFI
      - FCs on same plane as LFI, while HFI on different plane

■ Third option to separate LF FC into separate tunnels would increase vertex complexity without



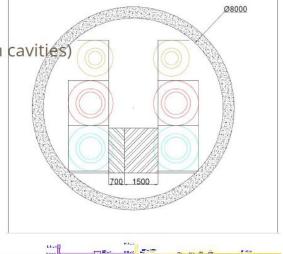
## Filter cavity positioning - Triangle

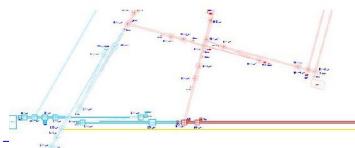
#### In summary:

- 4.0m periscope
- 550m path from IFO to SQZ-lab

 Beam diameter should be around 4-5 cm thus beam reducing telescope by a factor 12 is needed between OFI and SEM vessels

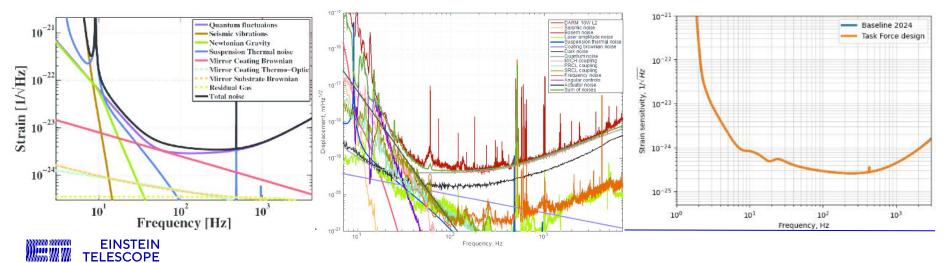
- Flexibility on positioning of periscope (close to IFO or to SQZ-lab)
- 3.4m horizontal separation between cavities (3.8m for main ITF arm cavities)
- Preliminary tunnel layout results in 8.0m tunnel envelope diameter





## Integrated towers - seismic isolation

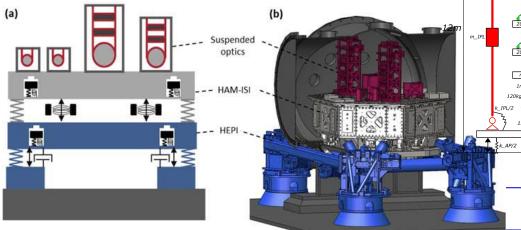
- aLIGO vs ET sensitivity:
  - o 10x @30 Hz
  - o 100x @10 Hz
  - o ~10^6 @ 2÷3 Hz
- Direct transmission of seismic noise only relevant for ET\_LF
- Control noise to be reduced by ~10÷100 in ET\_HF, and by > 100 in ET LF

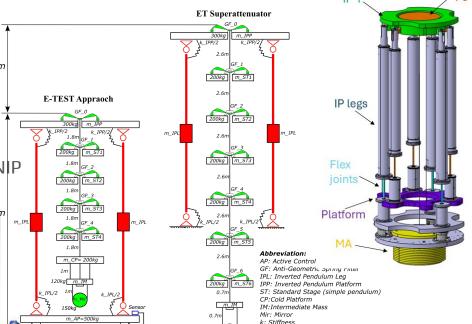


Integrated towers - seismic isolation

- Design concepts from existing detectors
- Soft suspension (Virgo)
  - highest filtering of direct seismic noise
  - Compatible with both access options
- Active platform (LIGO)
  - Compatible with co-located optics
  - Lateral access for top-loaded benches
  - Lower height

Hybrid systems, e.g. E-TEST, or new concepts, e.g. NIP

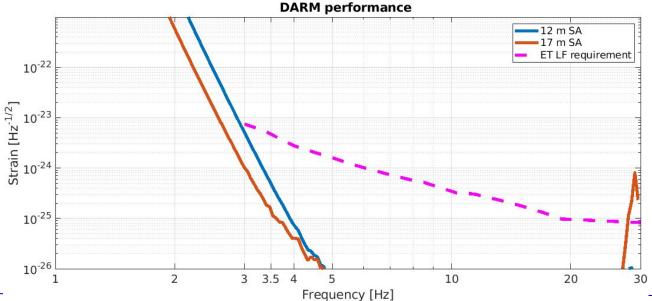




m· Mass

## Integrated towers - seismic isolation

- For ET HF, new baseline has short suspensions (except TMs)
  - Compatible with moderate seismic noise requirements
- For ET LF, test mass tower reduction from 17 m to 12 m
  - compatible with science case

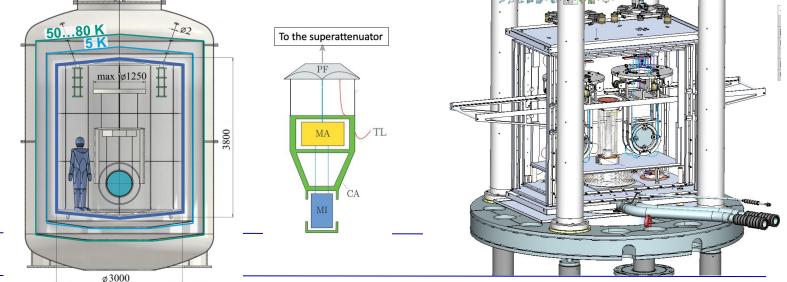




# Integrated towers - cryogenic payload Baseline conceptual design from Instrument Science Board of ET collaboration

- - Bottom access, tank footprint diameter ~5 m
- Alternative concept from ET Pathfinder to reduce size of cryostat
  - lateral access, tank footprint diameter ~4÷5 m
  - inverted pendulum base underneath cryostat
- No impact on tunnel diameter for 2L, marginal impact for Triangle

Large technical risk in both options, no need to constrain design from detector layout volumes



## Integrated towers - tank access

- Two types of access: lateral and bottom access
  - pros and cons of access options identified
- Assessment of sizes for different cleanroom classes from cleanliness requirements

Assessment on constraints for tower access from optica/detector layout or from suspension design

input to flexibility envelope of detector layout

80 K.

19 m

10 K

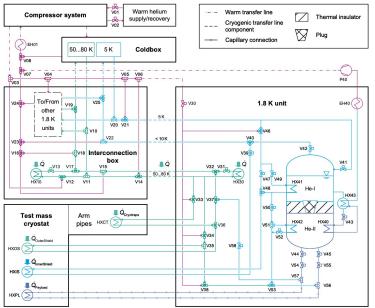
2 K

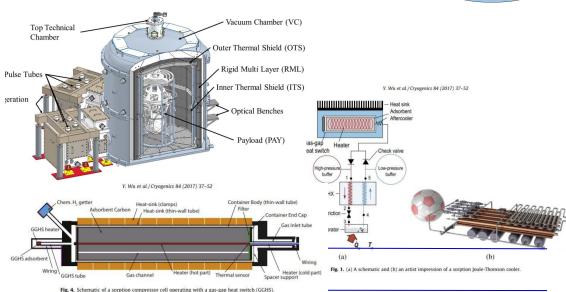
10-

20 K

Detector layout - cryogenic system

- - 10 K cryogenic payload and surrounding cryostat for thermal noise mitigation on LF TM
  - cryotraps around LF TM for thermal shield
  - other cryotraps for UHV in both LFI and HFI
- Baseline design: liquid He distribution for all tasks (Instrument Scienc Board in ET collaboration)
- Alternative concepts:
  - battery of pulse tubes (KAGRA): issues with underground power dissipation, noise
  - liquid N + sorption coolers (ET Pathfinder): issues with safety, efficiency of sorption coolers





- Double cavern concept
  Advantage: risk mitigation by removing interface between cryostat and inverted pendulum
- Concept: reduce main cavern height, move inverted pendulum to upper cavern (like KAGRA), all mechanical filters outside of interconnecting shaft (unlike KAGRA)
- Definition of inter-cavern separation & access tunnels would require civil engineering studies

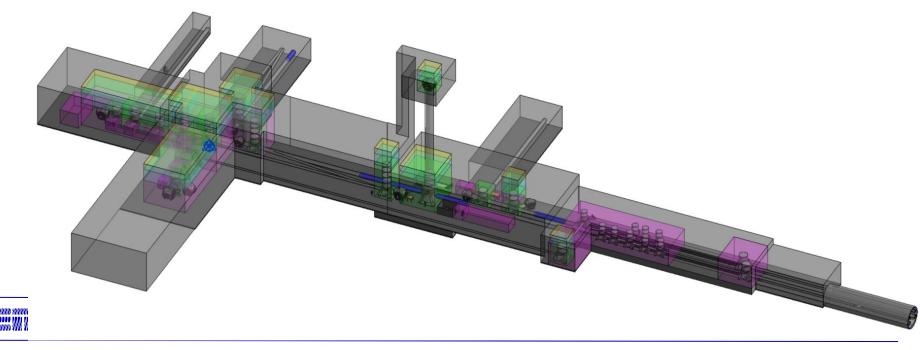




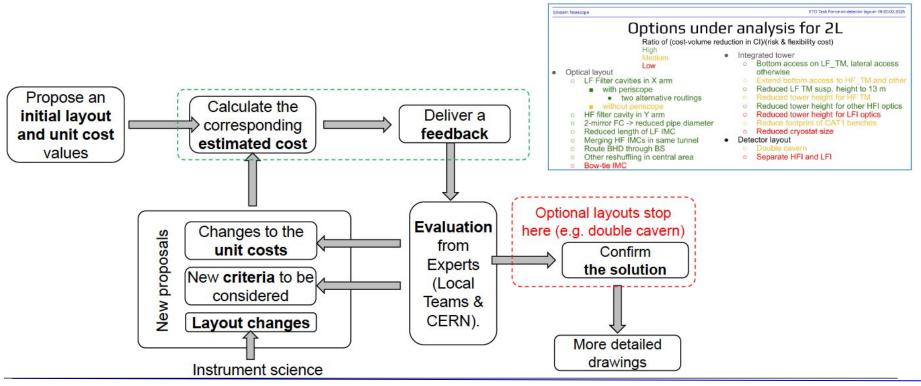
Double Cavern @KAGRA. source O. Miyakawa

Double cavern concept

- Advantage: risk mitigation by removing interface between cryostat and inverted pendulum
- Concept: reduce main cavern height, move inverted pendulum to upper cavern (like KAGRA), all mechanical filters outside of interconnecting shaft (unlike KAGRA)
- Definition of inter-cavern separation & access tunnels would require civil engineering studies



### Interface with civil infrastructure



## Engineering requirements

- Technical infrastructure has several critical interfaces with instrument
  - deserves coordinated technical design
- Task Force identified most critical requirements with impact on infrastructure cost, and provided a preliminary assessment

| REQ_ENG_001 REQ_ENG_002 REQ_ENG_003 | arm tunnel center inclination water tightness humidity in tunnel tunnel cleanliness | mrad<br>m<br>%<br>ISO | <1.5<br>>250<br>40÷60<br>ISO 9     | H/V coupling on TM, same order as minimum inclination from Earth curvature distance from TM: all caverns and part of arm cavity tunnel lifetime of equipment | noise)  Newtonian Noise VAC, electronics, ENV sensors VAC, electronics,  | initial guess<br>initial guess<br>initial guess |
|-------------------------------------|---|-----------------------|------------------------------------|--|--|---|
| REQ_ENG_001 REQ_ENG_002 REQ_ENG_003 | inclination<br>water tightness<br>humidity in tunnel                                | m<br>%                | >250<br>40÷60                      | from Earth curvature<br>distance from TM: all caverns<br>and part of arm cavity tunnel<br>lifetime of equipment  | noise)  Newtonian Noise VAC, electronics, ENV sensors VAC, electronics,  | initial guess<br>initial guess                  |
| REQ_ENG_002                         | water tightness<br>humidity in tunnel   | m<br>%                | >250<br>40÷60                      | distance from TM: all caverns<br>and part of arm cavity tunnel<br>lifetime of equipment  | Newtonian Noise<br>VAC, electronics,<br>ENV sensors<br>VAC, electronics, | initial guess                                   |
| REQ_ENG_003                         | humidity in tunnel  | %                     | 40÷60                              | and part of arm cavity tunnel lifetime of equipment  | VAC, electronics,<br>ENV sensors<br>VAC, electronics,                    |   |
| REQ_ENG_003                         | humidity in tunnel  | %                     | 40÷60                              | lifetime of equipment  | VAC, electronics,<br>ENV sensors<br>VAC, electronics,                    |   |
|                                     |   | 7.650                 |                                    |  | ENV sensors<br>VAC, electronics,   | initial guess                                   |
|                                     |   | 7.650                 |                                    |  | VAC, electronics,  | initial guess                                   |
| REQ_ENG_004                         | tunnel cleanliness  | ISO                   | ISO 9                              |  |  |   |
| REQ_ENG_004                         | tunnel cleanliness  | ISO                   | ISO 9                              |  |  |   |
|                                     |   |                       |                                    | lifetime of equipment  | ENV sensors  | initial guess                                   |
|                                     |   |                       | ISO 9 in general, ISO 7 and better | vacuum contamination, stray  |  |   |
|                                     |   |                       | in cleanrooms as specified in      | light from dust, monolithic  |  |   |
| REQ_ENG_005                         | cavern cleanliness  | ISO                   | detector layout                    | suspension failure from dust   | VAC + SLC + PAY  | initial gues:                                   |
| NAME OF TAXABLE PARTY.              | temperature stability in  |                       | natural undeground stability is    |  | VAC, electronics,  |   |
|                                     | tunnels   | deg                   | sufficient                         |  | ENV sensors  | initial gues                                    |
|                                     |   |                       |                                    | actuation range for  |  | 7.5   |
|                                     |   |                       |                                    | suspensions, misalignment of   | f  |   |
|                                     | temperature stability in  |                       |                                    | in-air optical systems, and  | SUSP + OPT + INJ +   |   |
| REQ_ENG_007                         | caverns   | deg                   | 0.1/hr, 0.2/day, 0.3/month         | many other   | other  | initial gues:                                   |
|                                     |   |                       |                                    | in-air optics, maybe other   |  |   |
| REQ_ENG_008                         | humidity stability in caverns   | 96                    | +/-5 within 40÷60                  | equipment  | Virgo experience   | initial gues:                                   |
|                                     |   |                       |                                    | seismic noise, acoustic noise,   | er mannananananan  |   |
| REQ ENG 009                         | ventilation   | TBD                   | TBD                                | NN   | SUSP, NN   | initial guess                                   |
|                                     |   |                       |                                    | main building blocks, i.e.   |  | Water Table                                     |
| REQ ENG 010                         | logistics   | m                     | as specified in detector layout    | cryostat and pipe sections   | CRYO, VAC  | initial guess                                   |
|                                     |   | #, m^3,               |                                    | safety/escape routes, gate   |  |   |
| REO ENG 011                         | recesses in tunnels   | m^2                   | TBD                                | valves   | CE, VAC  | initial gues                                    |
|                                     |   |                       | 1.55                               | infrastructure lifetime of at  |  | 0   |
| REQ ENG 0                           | expected lifetime   | year                  | >50                                | least 50 years   |  |   |
| -                                   | allowable main arm tunnel   | mm.                   | - 55                               | Maintaining optical axis   |  |   |
|                                     | deformation   | mm/yr                 | TBD                                | alignment.   |  |   |
| nLQ_LNO_0                           | delottidaton  | iiiiii/yi             | 100                                |  |  |   |
|                                     |   |                       |                                    | To limit stress on welding lips<br>(a few mm of differential   |  |   |
|                                     | allowable differential  |                       |                                    |  | VAC, Virgo   |   |
|                                     | deformation   | mm                    | TBD; per 20m beampipe segment      |  | experience   |   |

- Besides the space demands, technical requirements from detector may largely affect the cost of underground infrastructure
- The civil engineering working group in task force collected a list of relevant requirements
  - · to be possibly attached to the baseline detector layout
  - to allow for a technical feasibility study on the ET infrastructure.

| ET / ETD / QA_ETO_LT / tosues  | New York: 0                 |
|--|-----------------------------|
| REQ_END_01t: Recesses in the tunnels     resided weeks ago by yatrick-wereks   | <b>B</b> 1                  |
| © REQ_ENG_010: Logistics requirements<br>#26 - created 2 weeks ago by satistic weenake   | eptimed 2 weeks ago         |
| © REQ_ENG_008: Verbilation requirements #27 - created 2 wome, ago by particle woments  |                             |
| © REQ_ENG_007: Temperature and its stability requirements in the caverns 420: revaled 2 weeks upo by gatest-warrale  | Rh 1<br>updated 3 weeks ago |
| © REQ_EMG_006: Temperature and its stability requirements in the main aren tunnel.  #22 - created 2 weeks up to yate: Aventable Coverse Form 44 (Teverse Windowster)    Coverse Form 44 (Teverse Windowster)   Coverse Form 44 (Teverse Windowster)   Mat Wakkindskidosadden.  | ₽2 2<br>updated 2 weeks ago |
| Tunnel Diameter Requirements for the Triangular and 2L Configuration     #2d - created 2 weeks ago by patick-werrein   | updated 2 weeks ago         |
| []* REQ_ENG_008: Relative Humidity requirement in the cavern<br>#32 - crosted 2 weeks ago by patick-wereine  | © 3<br>updated 2 weeks ago  |
| D REQ_EMG_DOS: Relative Humidity requirement in the funnet size created a week size by parks were all the control of the fundament of the control of the con | apdaled 1 week ago          |
| © REQ_ENG_005: Cleanliness requirement for the cavern<br>#21- created 2 words ago by particl wantels   | 码 2<br>upobited 2 weeks ago |
| REQ_ENG_004: Cleanliness requirement for the tunnel. 22: created 2 weeks upo by jamick-warrania.   | updated 2 weeks ago         |
| DY REQ_ENO_002: Which part of the infrastructure chould be water Sight? #10 : resided events got by apartick wereals (Contain Fillacing) Contain (Erection Contain Fillacing) (Contain Fillacing) Contain Fillacing) (Contain Fillacing) (Contain Fillacing) (Contain Fillacing)   | ez-totad 1 visek ágo        |
| ()* REQ_ENG_001: What is the maximum allowed inclination of the tunnels?  *III: crashed 2 weeks up to particl extension (Corests Corested 2006; Crest-41)* (Metabloout Assurance)  *III: crashed 2 weeks up to particl extension  *III: crashed 2 weeks up to particle extensi | PC 4 updated 2 weeks ago    |

## Noise budget and science case

- Use python code from ET scientific collaboration (Py-GWINC) to generate noise budget for different configurations
- Analyse sensitivity curves against ET science case with a set of figures of merit
  - Compare 2024 reference with new baseline layout from Task Force
  - Derive scientific requirements on critical design parameters

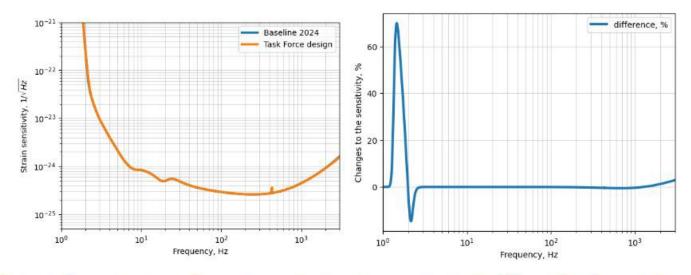


Figure 6: Comparison between the baseline sensitivity and the task force result (left) and the difference between the old and new sensitivities, expressed in % (right). Positive represents better sensitivity in the task force design. The difference at very low frequency comes from the new suspension design, and the difference at the very high frequency from slightly reduced HF signal extraction cavity length.

## Background information

- Terms of Reference of External Review Committee (shared);
- ETO Task Force mandate (shared);
- Optical layout 2024 document for Triangle (pdf);
- Optical layout 2024 document for 2L (pdf);
- 2D drawing of optical layout 2024 for Triangle (pdf);
- 2D drawing of optical layout 2024 for 2L (pdf);
- Detector layout 2024 document for the Triangle (pdf);
- Detector layout 2024 document for the 2L (pdf);
- 3D model of 2024 detector layout (trimble connect) for the Triangle;
- 3D model of 2024 detector layout (trimble connect) for the 2L;
- Trimble guideline;
- ESFRI proposal: 2020 CDR;
- Tunnel diameter requirements (pdf);
- Reference document for cryogenic system (pdf);
- LF TM suspension document (draft pdf);
- Suspension system classification (pdf) completely changed in ETO Task Force work, highlight relevant sections;
- Science case: COBA paper (pdf);
- ET noise budget: sensitivity curve update (pdf);
- Reference on Civil Engineering (TBD);
- Guideline how to read the documents.



## International review committee

- Following the mandate by ET Coordinators, ETO directorate set up an international review committee
  - composition: 8 members from LIGO Lab, LSC, KAGRA, PSI, CERN
  - o terms of reference: review will focus on
    - Mandate Compliance & Infrastructure Feasibility
    - Clarity, Consistency & Supporting Information
    - Scientific & Design Justification
    - Risk, Flexibility & Decision Support
- Output documents by task force were delivered to review committee on 23/05
- Draft documents shared with ETC since 12/05 periodically updated until delivery
- Review outcome expected in ~2 weeks
- Task force to provide final version after review by end of June
- ETC EB is organising an independent review